

True coincidences and a Decent Currie

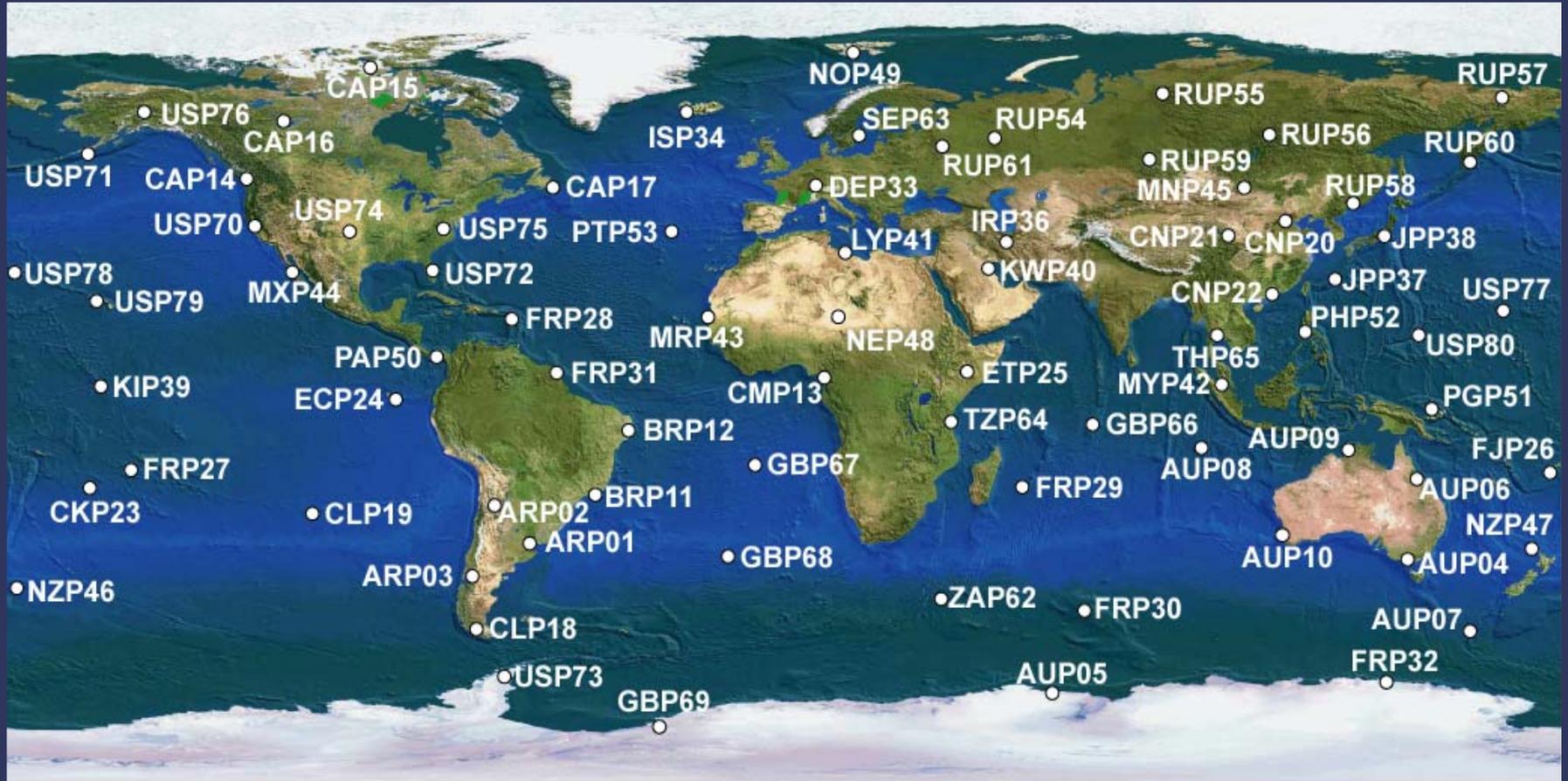
Lars-Erik De Geer

FOI, Swedish Defence Research Agency, Stockholm, Sweden

ledg@foi.se

NKS GammaSem2010, Lillestrøm, Norway, 28 September 2010

Radionuclide Particulate Station Network 79 + 1 TBD





SAUNA – Swedish Automatic Unit for Noble gas Acquisition



Acquires, Analyses and Disseminates Data and Products to Support States' Need for Nuclear-Test-Ban Monitoring in Multiple Environments



5 Geostationary Satellites

GLOBAL COMMUNICATIONS INFRASTRUCTURE

INTERNATIONAL DATA CENTRE

National Authorities

Radionuclide (80, 1/2 Xe)

Infrasound (60)

Seismic (50 Pri + 120 Aux)

Hydroacoustic (6 hydro, 5T)

INTERNATIONAL MONITORING SYSTEM

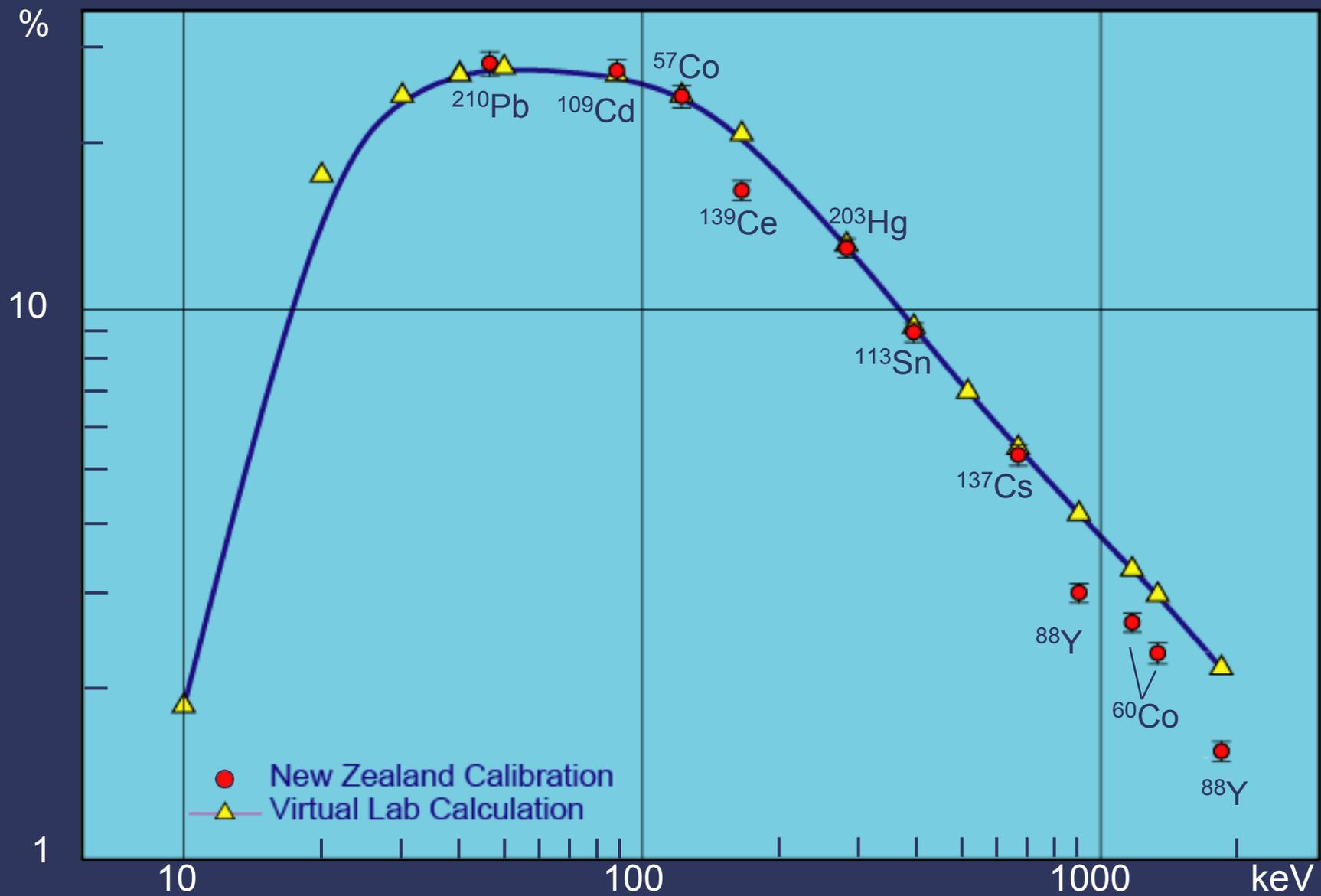


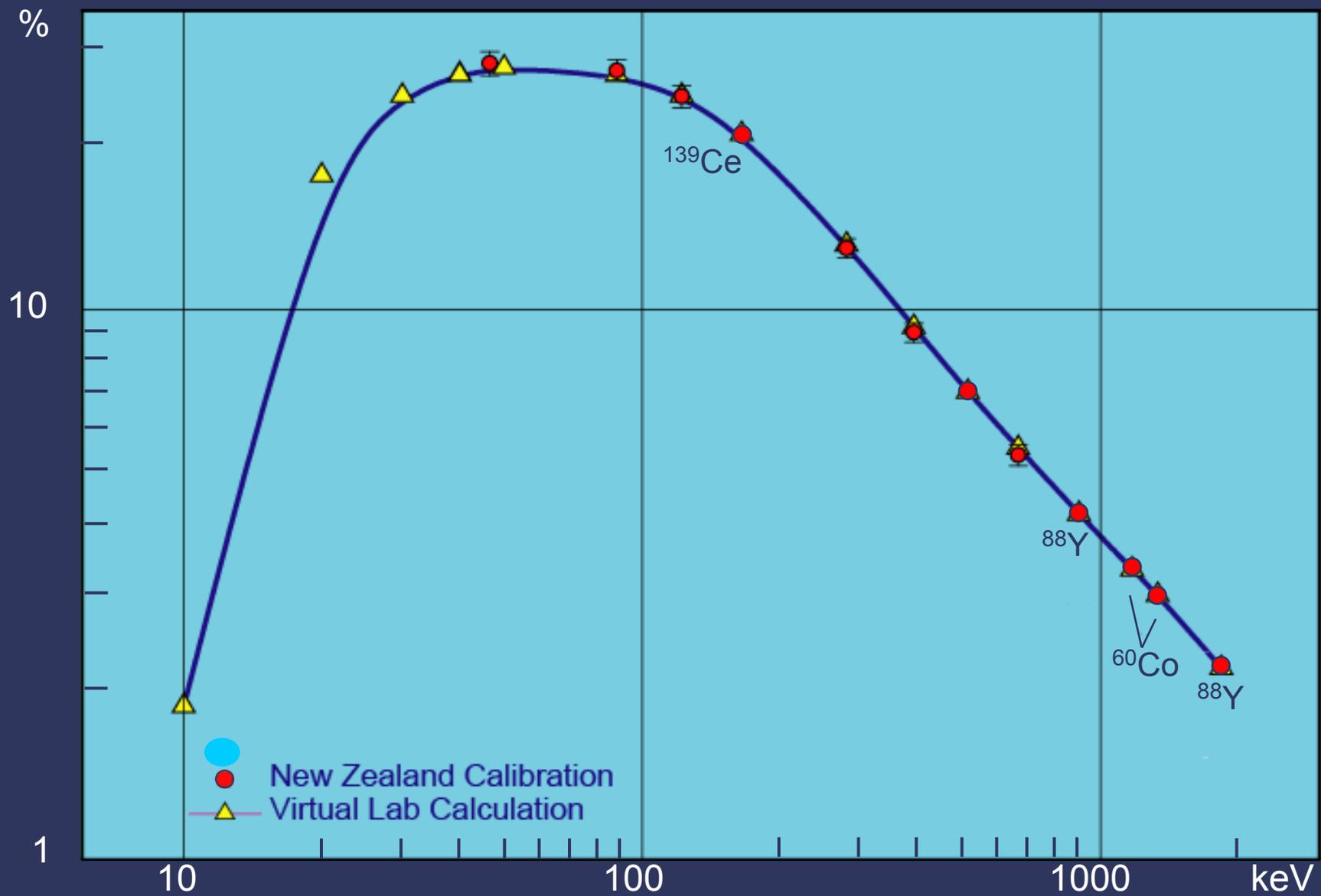
FOI

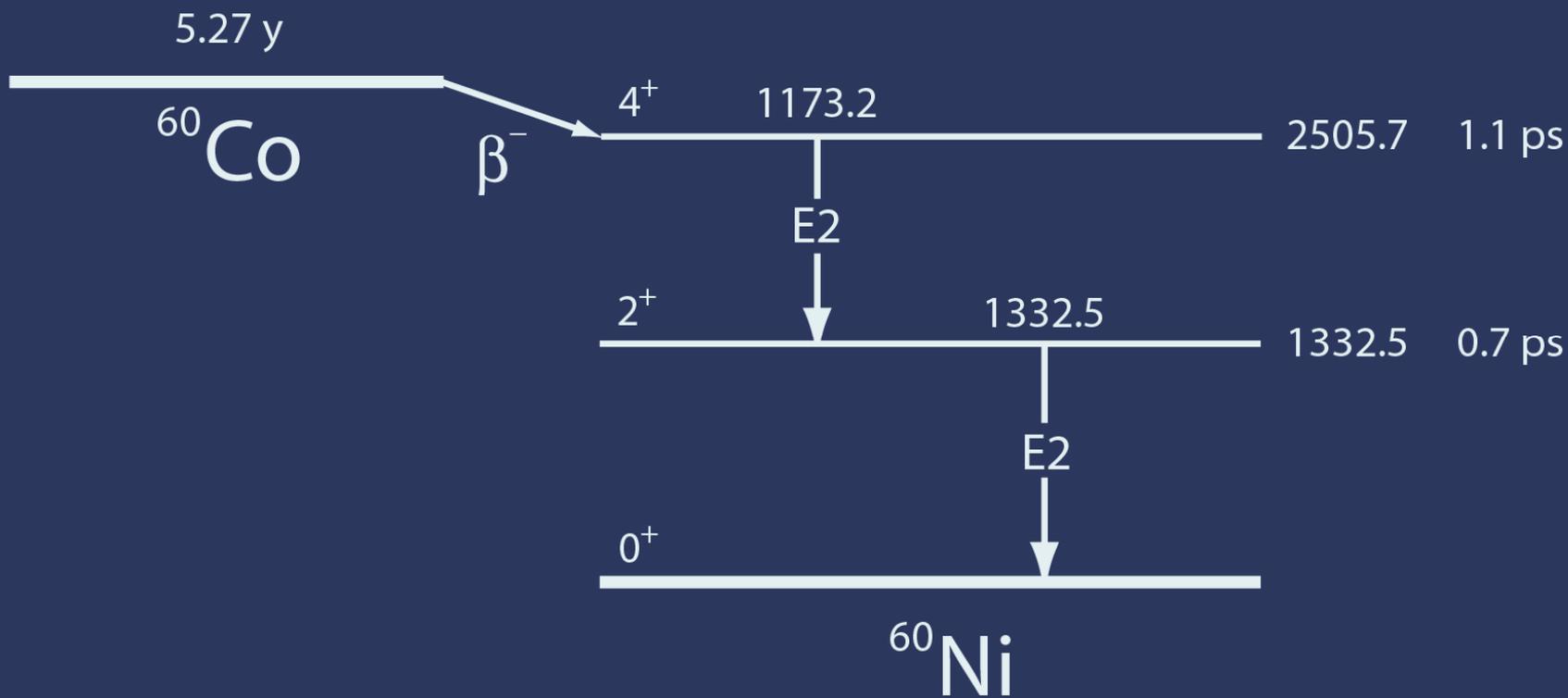
The CTBT Verification Regime

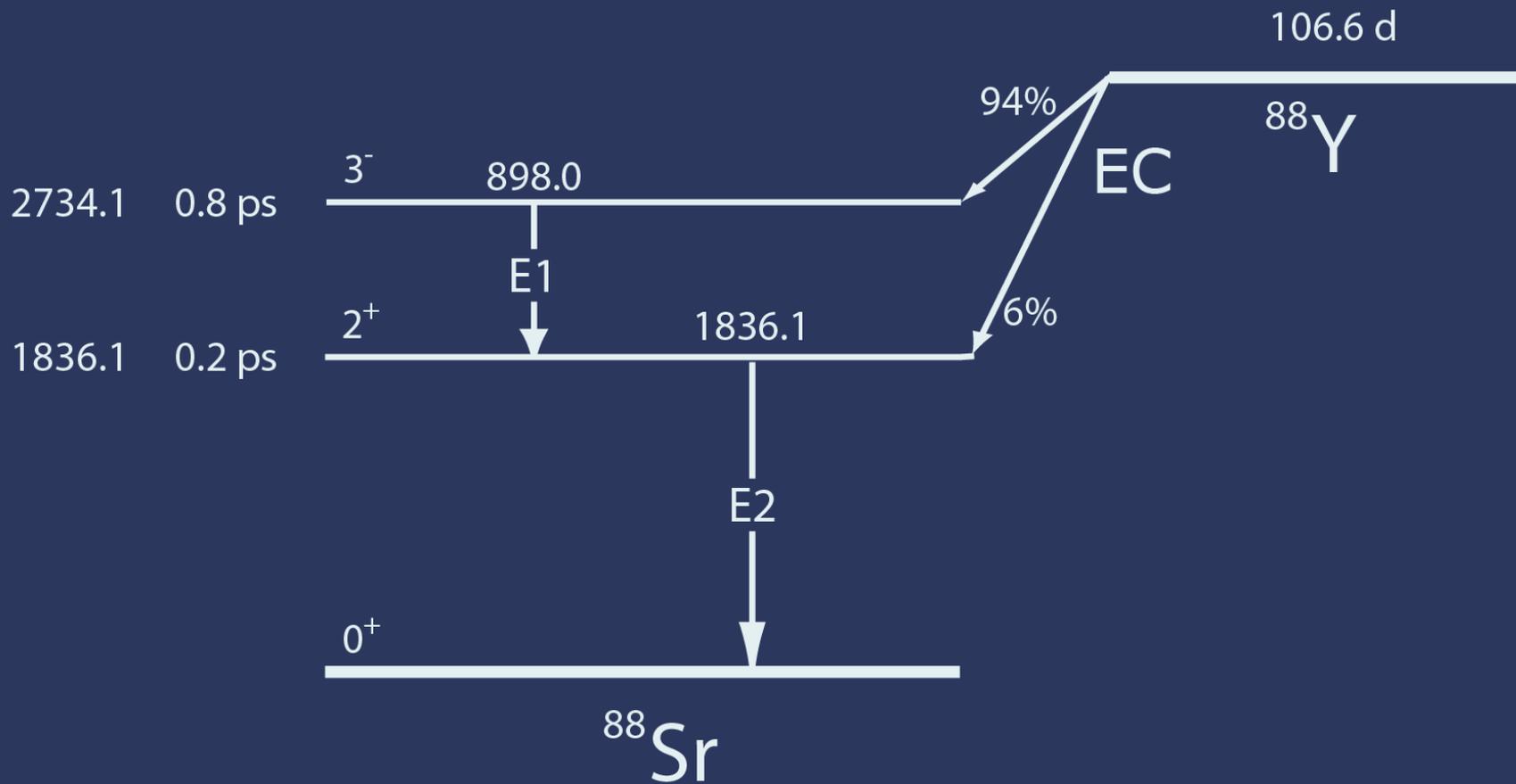
TCS

True Coincidence Summation

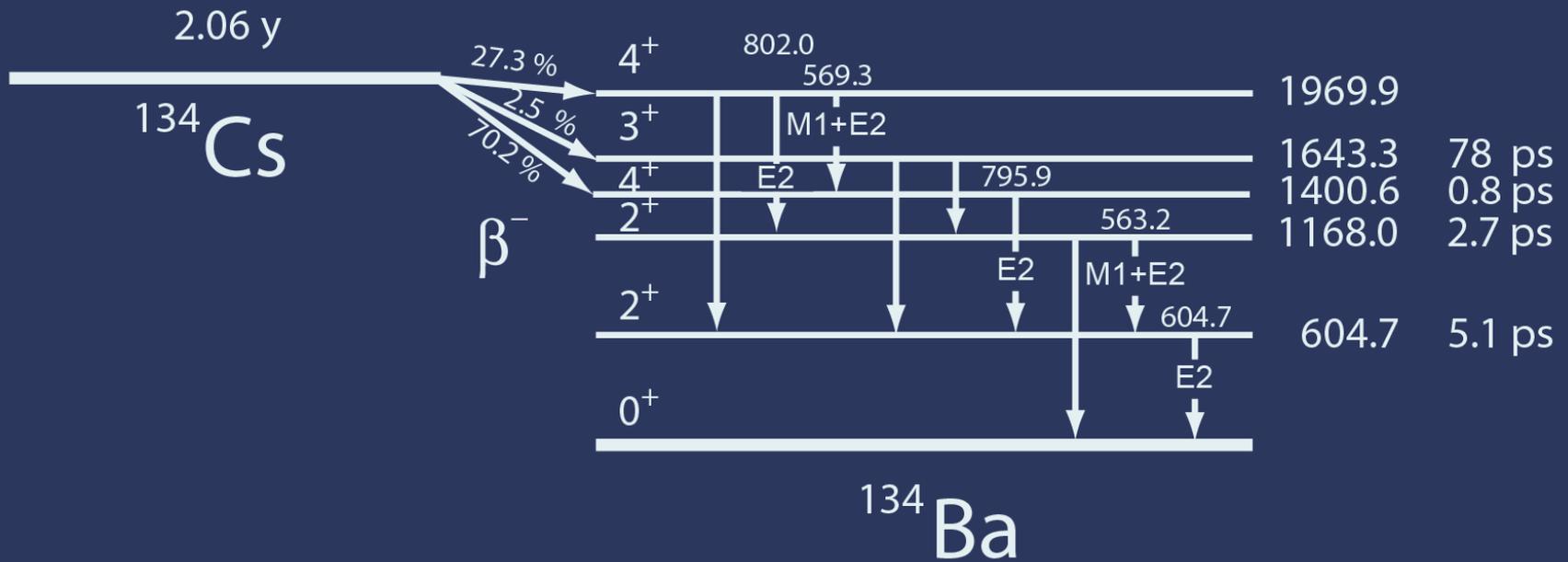


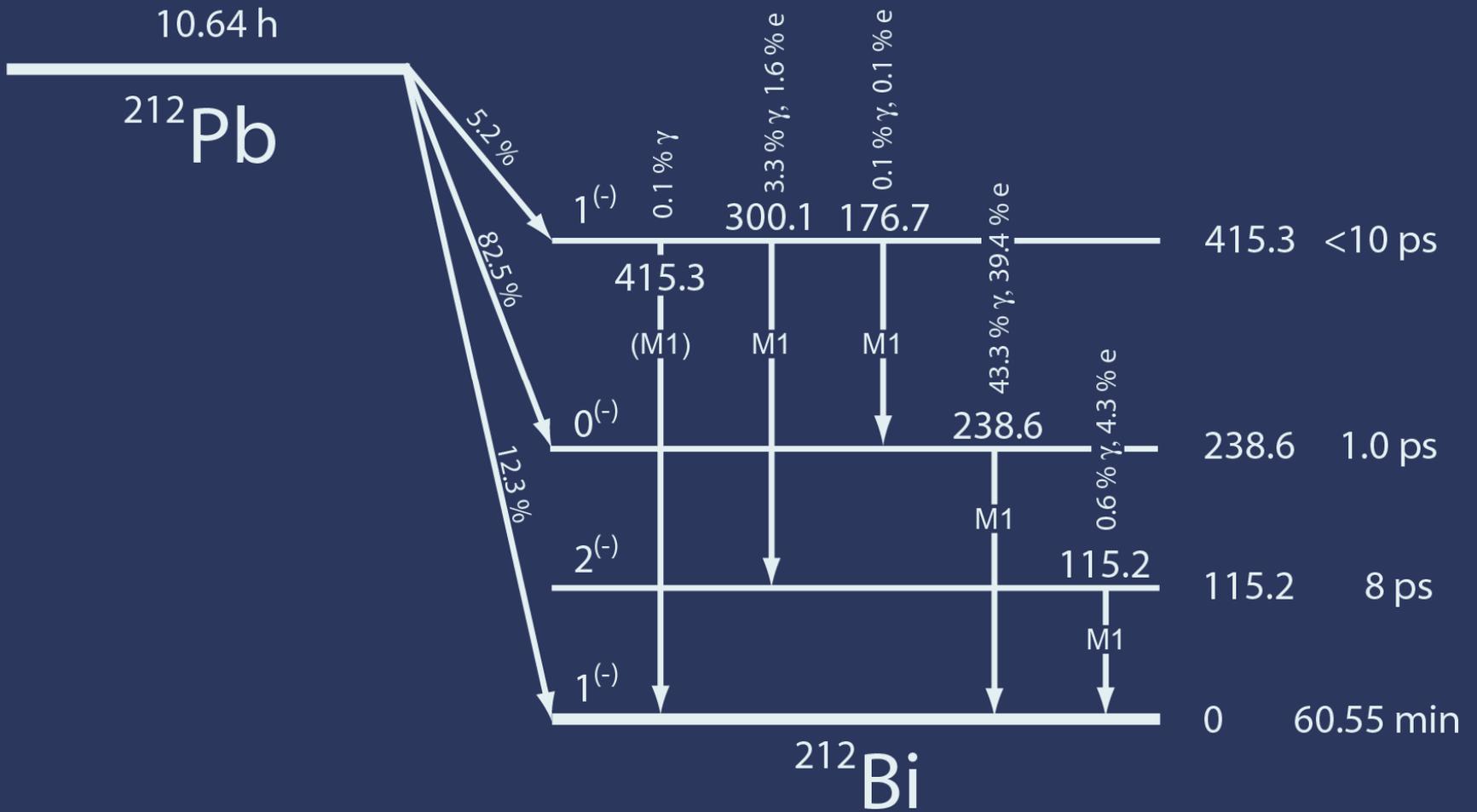


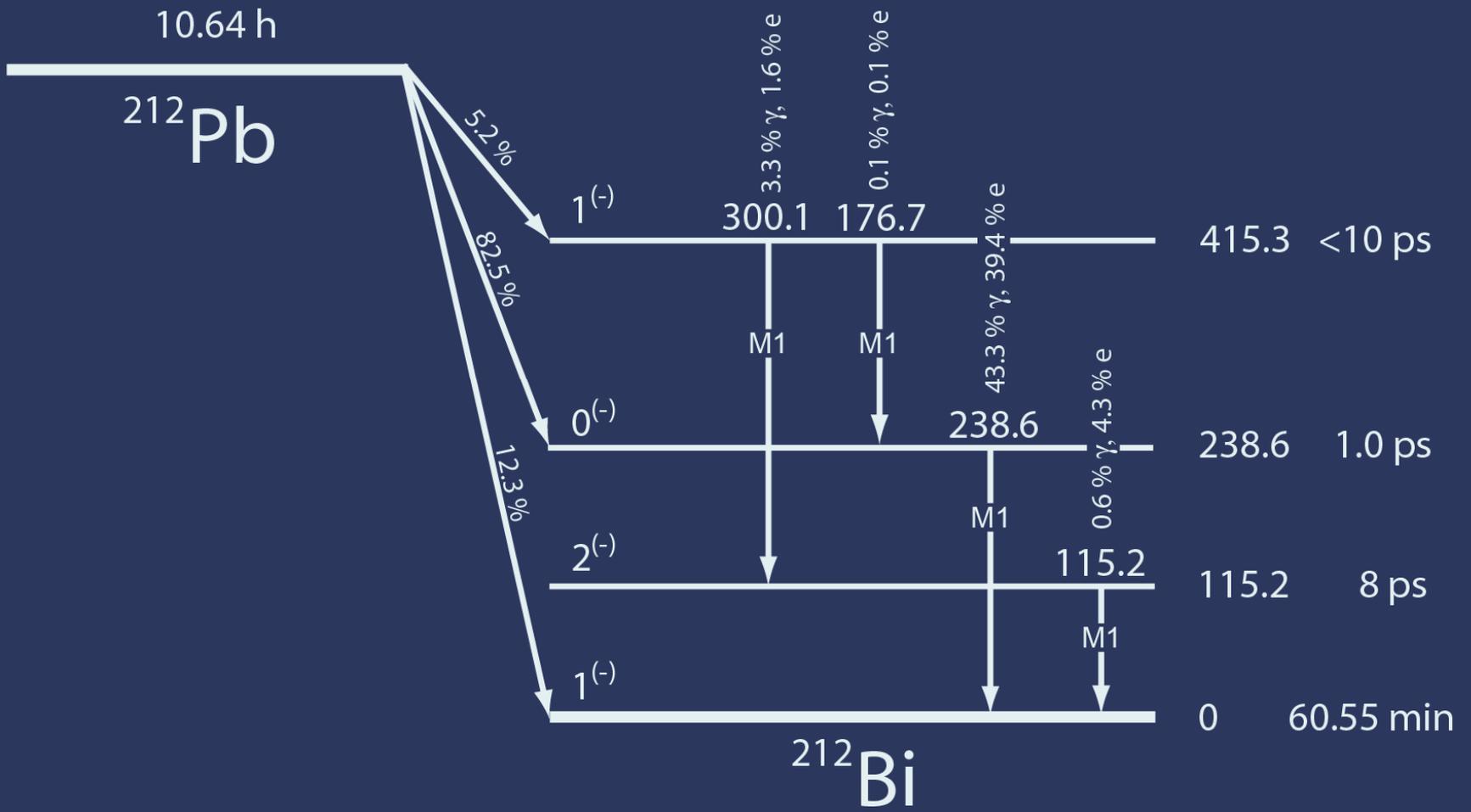




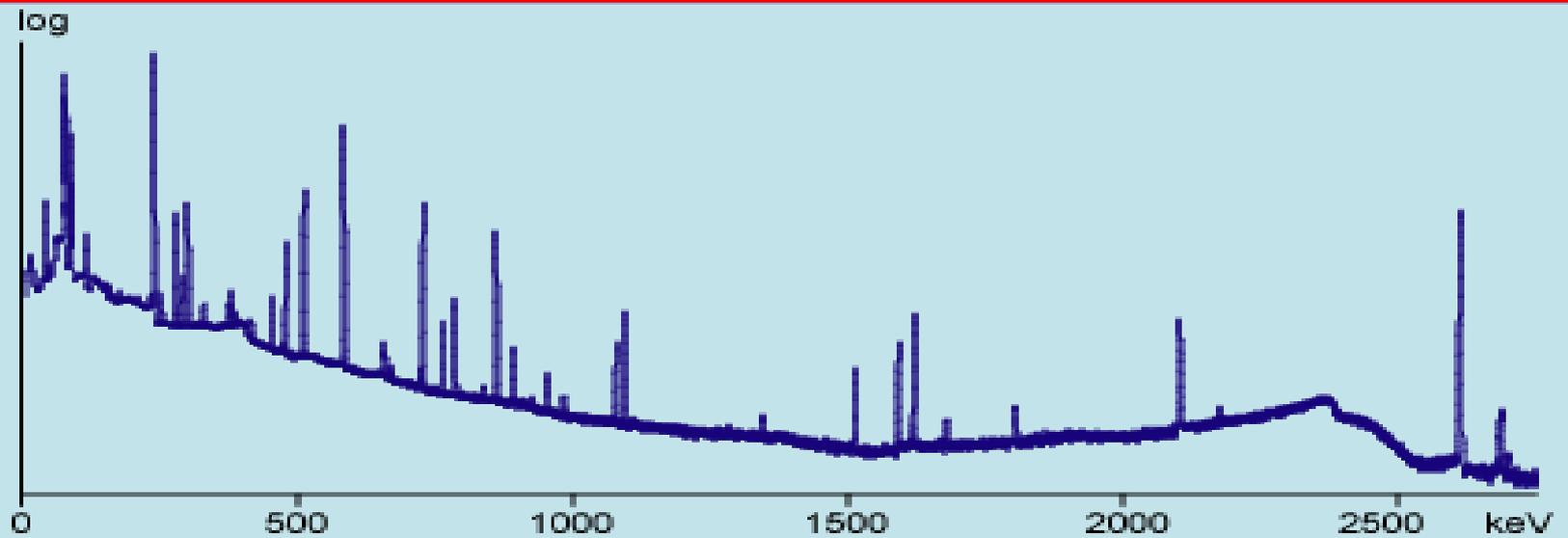




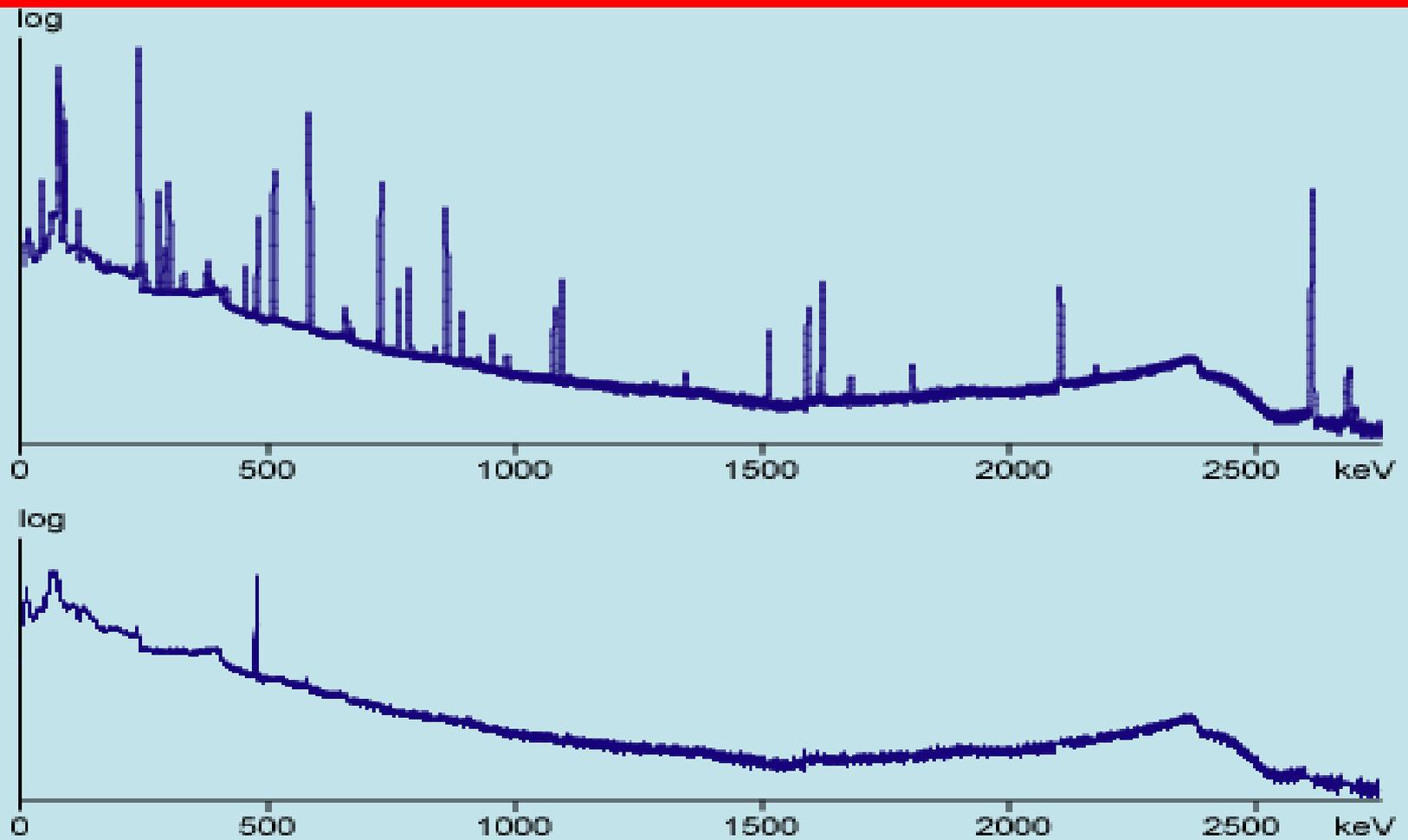




The Lead Picker



The Lead Picker



Sinnko

Debertin

Aaltonen, Ugletveit

Korsum

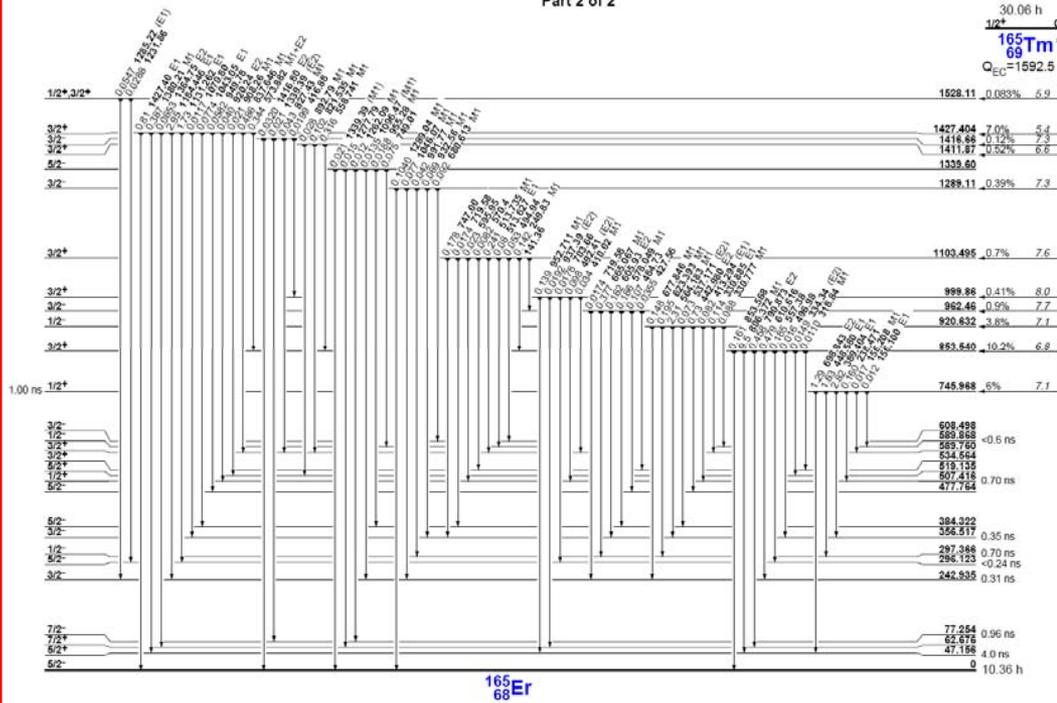
Shaman

GESPECOR

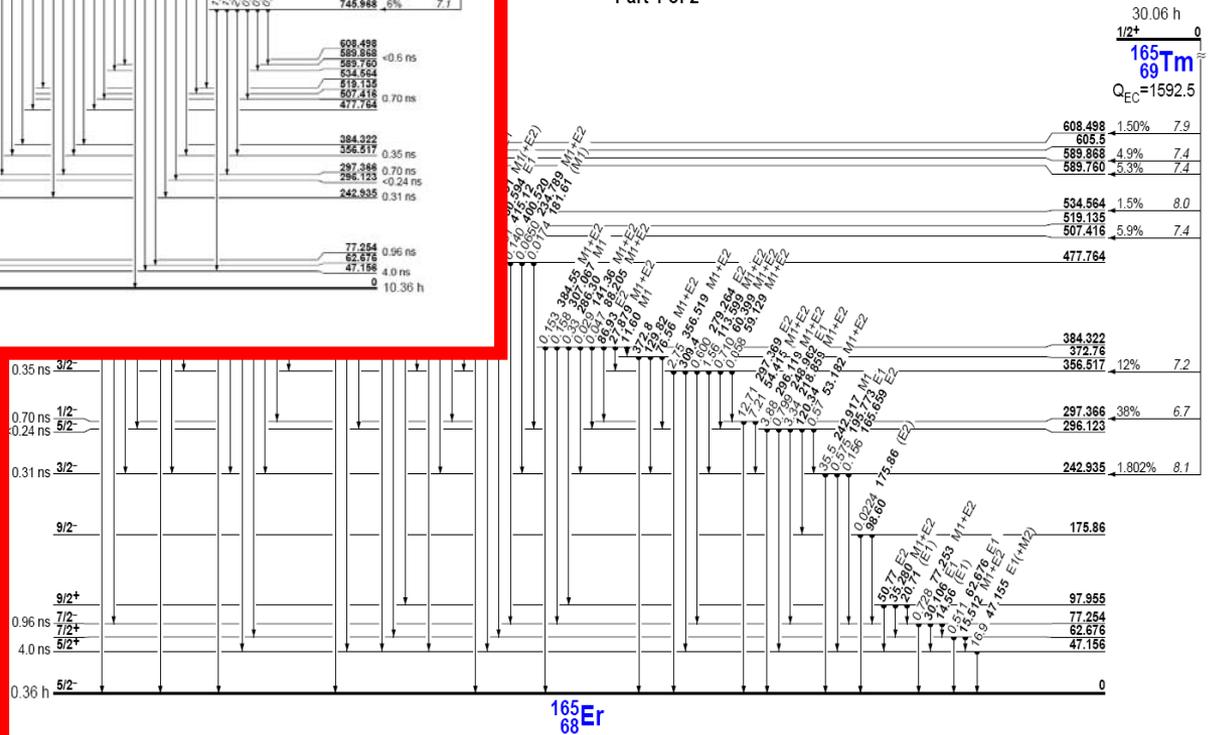
ISOCS, LabSOCS

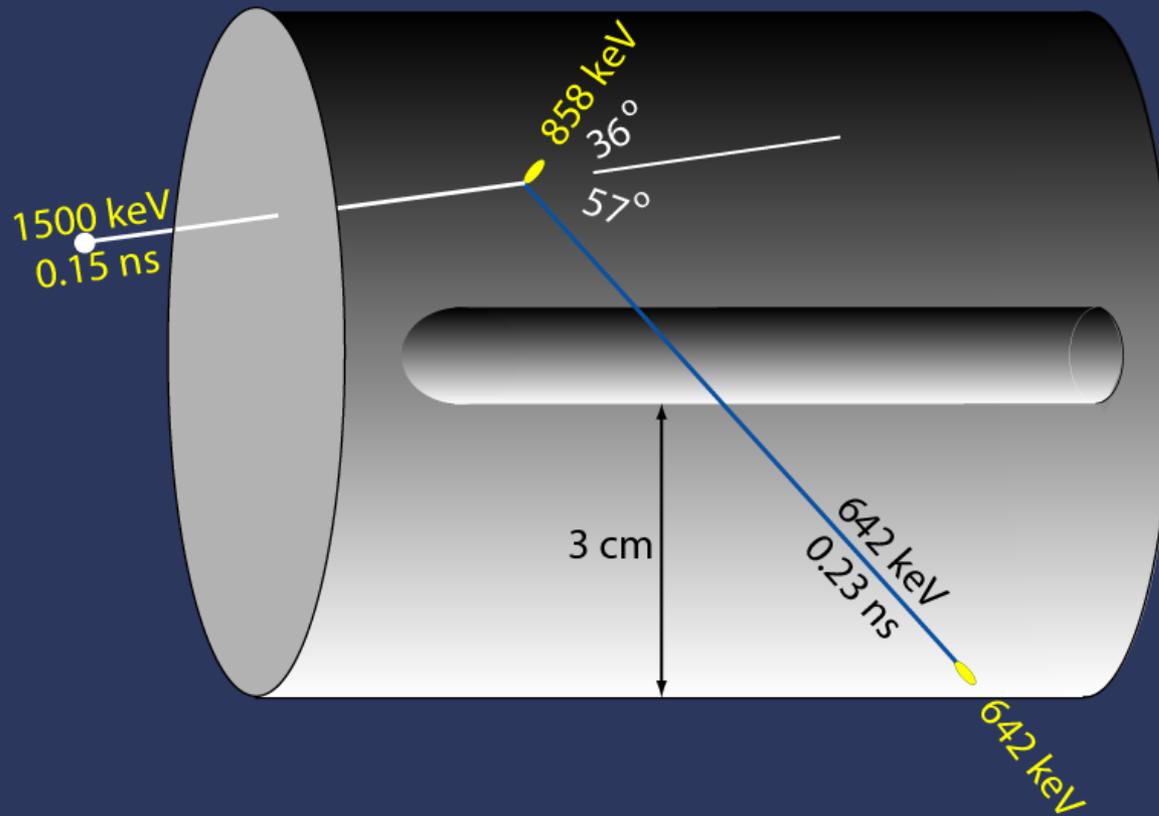
VGSL

Part 2 of 2



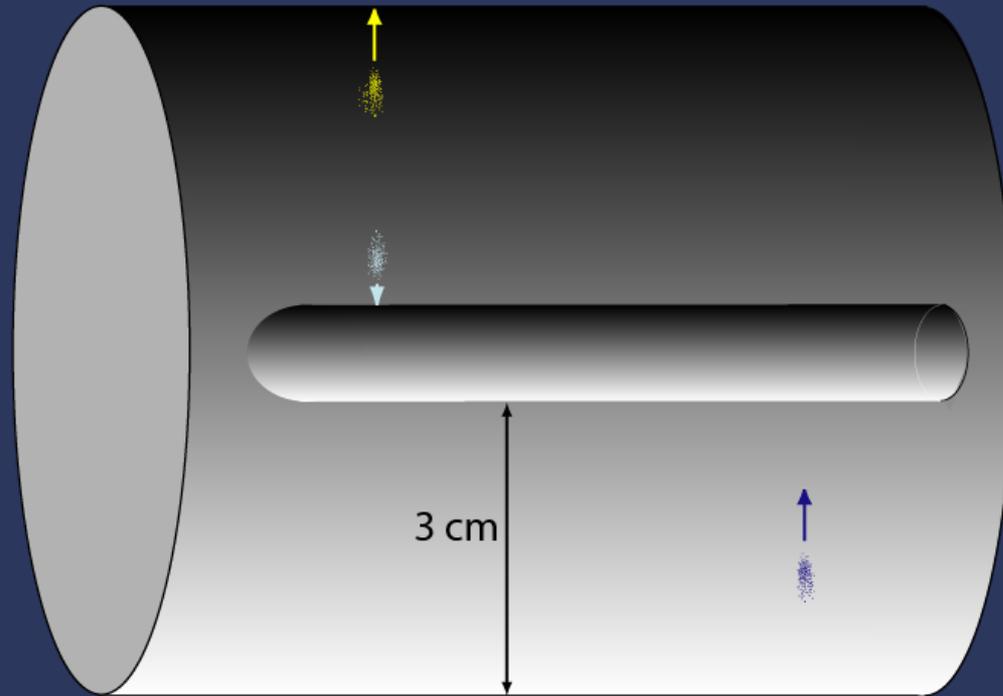
Part 1 of 2



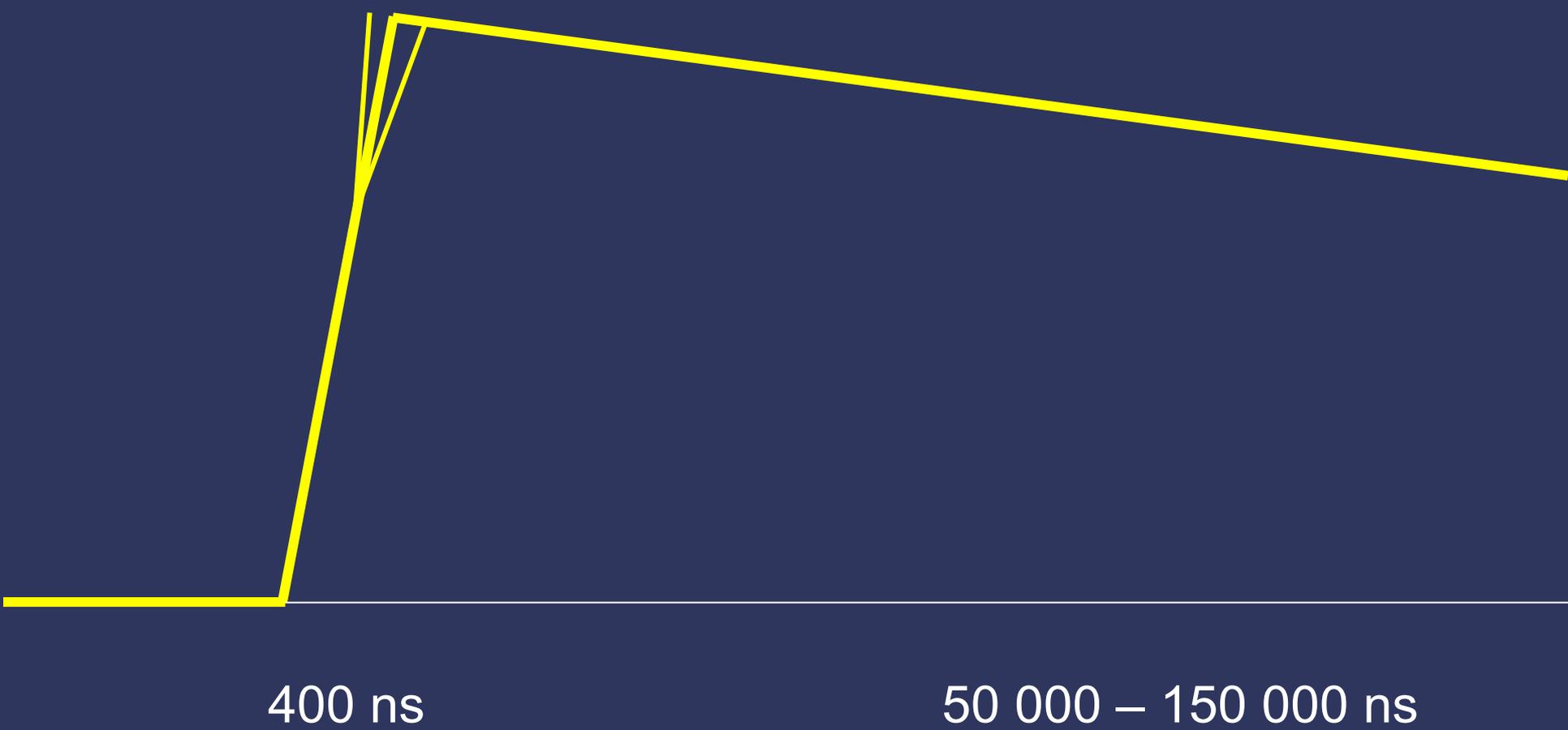


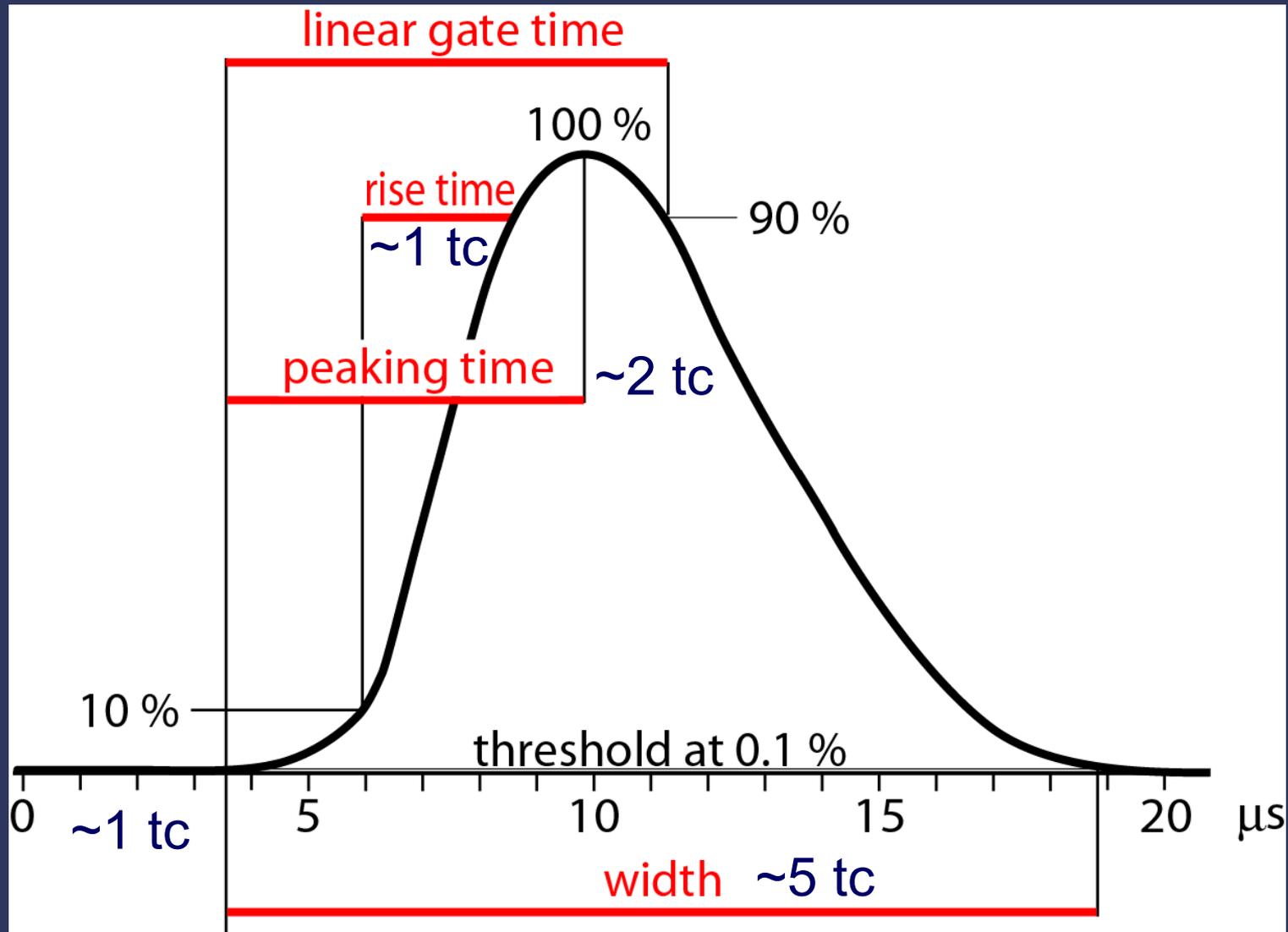
p-type cylindrical crystal, positive on the mantle and negative on the inner surface

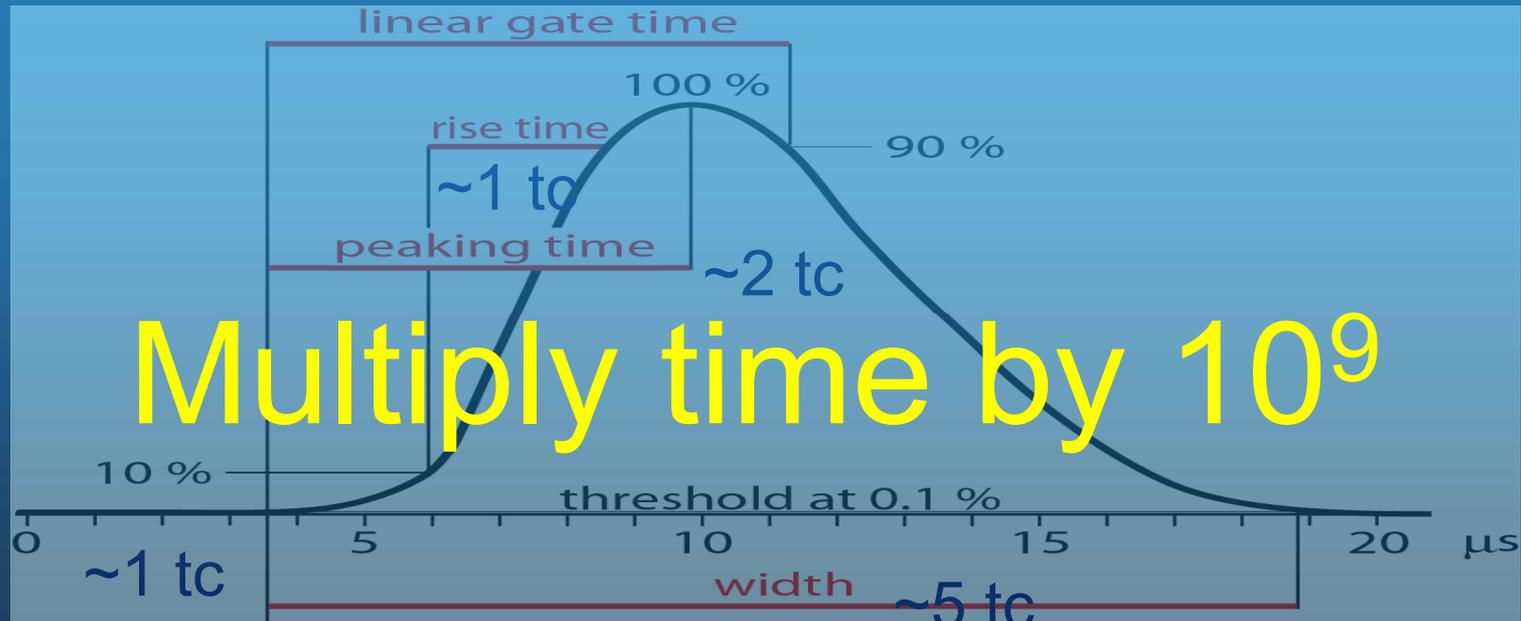
120 ns later



Saturated drifting velocities for both holes and electrons around 7.5 cm per μs imply a preamplifier rise time of about 400 ns. About 300 000 and 200 000 secondary electrons drift towards the mantle and the same number of holes drift towards the center hole [sic].







Everything nuclear happened in a 1 sec flash at breakfast this morning

Charge collection finished after eating a sandwich for less than 10 minutes

The useful part of the main amplifier pulse will be ready by lunch time

The ADC will deliver the count to the spectrum late tonight

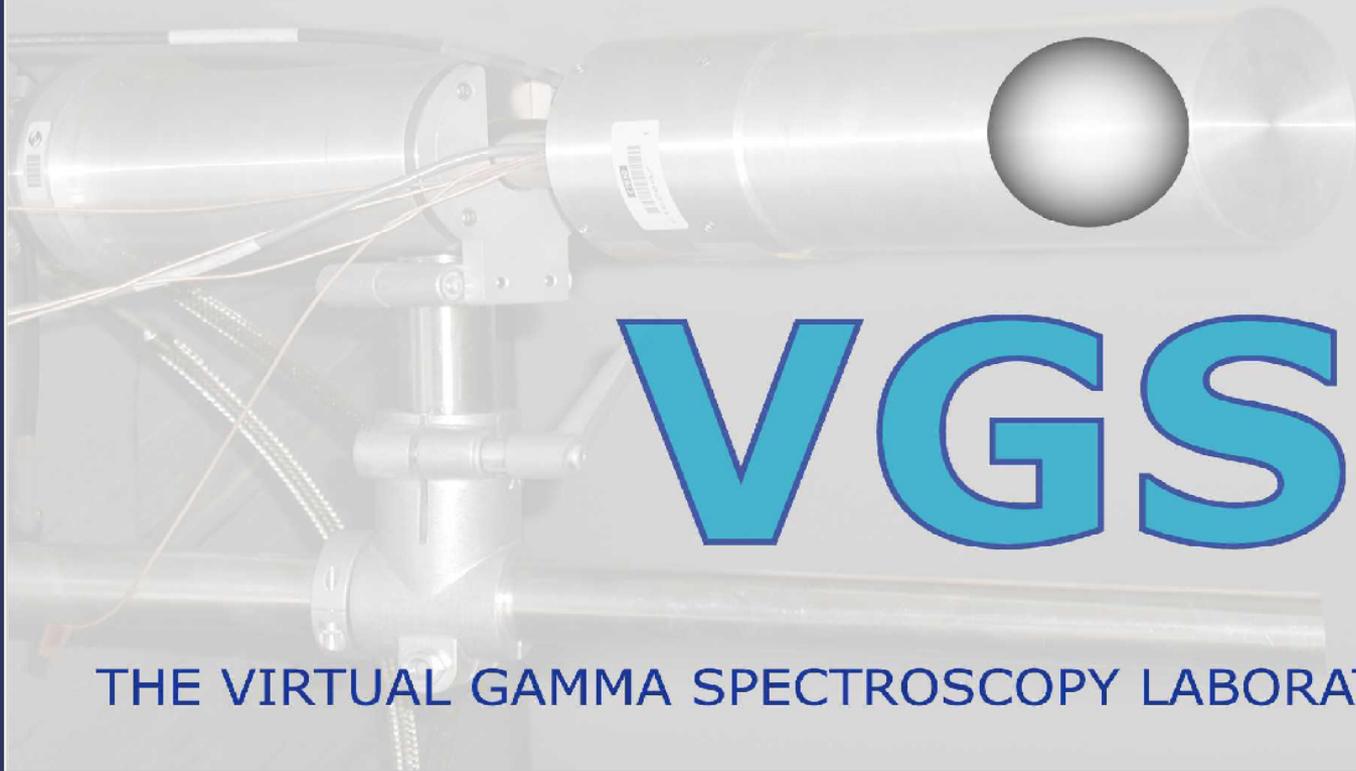
At 10 cps we can expect next event in November 2013

At 100 cps we can expect next event in mid-February next year.



Time constant $t_c = 3 \mu s$

Welcome to



VGSL

THE VIRTUAL GAMMA SPECTROSCOPY LABORATORY

Investing in the laboratory

Detector definition

Material

planar

Crystal

radius (cm) Hole radius (mm) length (cm) Hole length (mm)

Top fl of curvature (mm) Hole fl of curvature (mm)

Dead Layer

Top outside (mm) Inside the hole (mm)

Side outside (mm)

Crystal Holder

Radial thickness (mm)

End Cap Mantle

Inner radius (mm) Thickness (mm)

End Cap Window

Thickness (mm) Distance to Crystal (mm)

material? density(g/cm3)

Electronic Characteristics

$E(ch) = A + B(ch) + C(ch)^2 + D(ch)^3$

A B C D

$FWHM(E) = A + B(E)$

A B

Channels number

Existing Detectors

Detector name

Remove selected

CLEAR

Detector definition

Material

coaxial

Crystal

Radius (mm) Hole radius (mm) Length (mm) Hole length (mm)

Top fl of curvature (mm) Hole fl of curvature (mm)

Dead Layer

Top outside (mm) Inside the hole (mm)

Side outside (mm)

Crystal Holder

Radial thickness (mm)

End Cap Mantle

Inner radius (mm) Thickness (mm)

End Cap Window

Thickness (mm) Distance to Crystal (mm)

material? density(g/cm3)

Electronic Characteristics

$E(ch) = A + B(ch) + C(ch)^2 + D(ch)^3$

A B C D

$FWHM(E) = A + B \sqrt{E + C + E^2}$

A B C

Number of channels

Existing Detectors

Detector name

Remove selected

CLEAR

Shielding Definition

Material

Inner radius (mm) Inner height (mm)

First Layer

Thickness (mm) material? Density (g/cc)

Second Layer

Thickness (mm) material? Density (g/cc)

Third Layer

Thickness (mm) material? Density (g/cc)

Fourth Layer

Thickness (mm) material? Density (g/cc)

Existing

Shielding Name

Remove selected

CLEAR

Source Geometry Definition

Material

Bar

Existing Sources

petradisk point

Remove selected

Filter

Radius (cm) Thickness (cm)

Petradisk Density (g/cm3)

Container thickness

Top(cm) Bottom(cm) Side(cm)

PVC Density (g/cm3)

Source geometry name

CLEAR

Material Definition

Material name Standard density (g/cm3)

Select component

| | | | |
|----|-----------------|---|-----------------|
| He | Atomic Fraction | C | Atomic Fraction |
| Li | | | |
| Be | | | |
| B | | | |
| C | | | |

Existing Materials

PVC

Remove Selected

Lab Setup Definition

Detector Shielding Source Geometry Material

SETUP CONFIGURATION



Laboratory setup

Detector =

Shielding =

Source geometry = 

Distance from Detector window to inner Shielding cup surface

D (cm)

Shift of the source center to the detector axis (z=0 at end cup window)

Radial (cm) Axial (cm)

Existing Setups

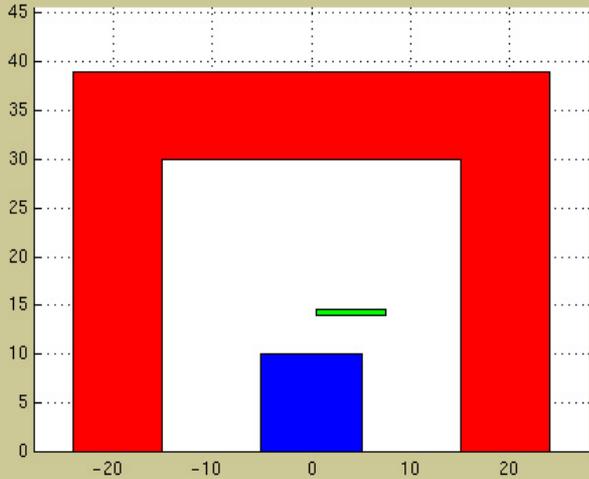
| |
|------------|
| NZP46_001 |
| NZP46_002 |
| NZP46_003 |
| Rarotonga3 |

Remove selected

Efficiency Table

| | |
|----|-----------|
| 10 | 0.0074737 |
| 20 | 0.1622 |
| 30 | 0.25765 |
| 40 | 0.29047 |
| 50 | 0.30329 |
| 60 | 0.30671 |

PLOT EFFICIENCY

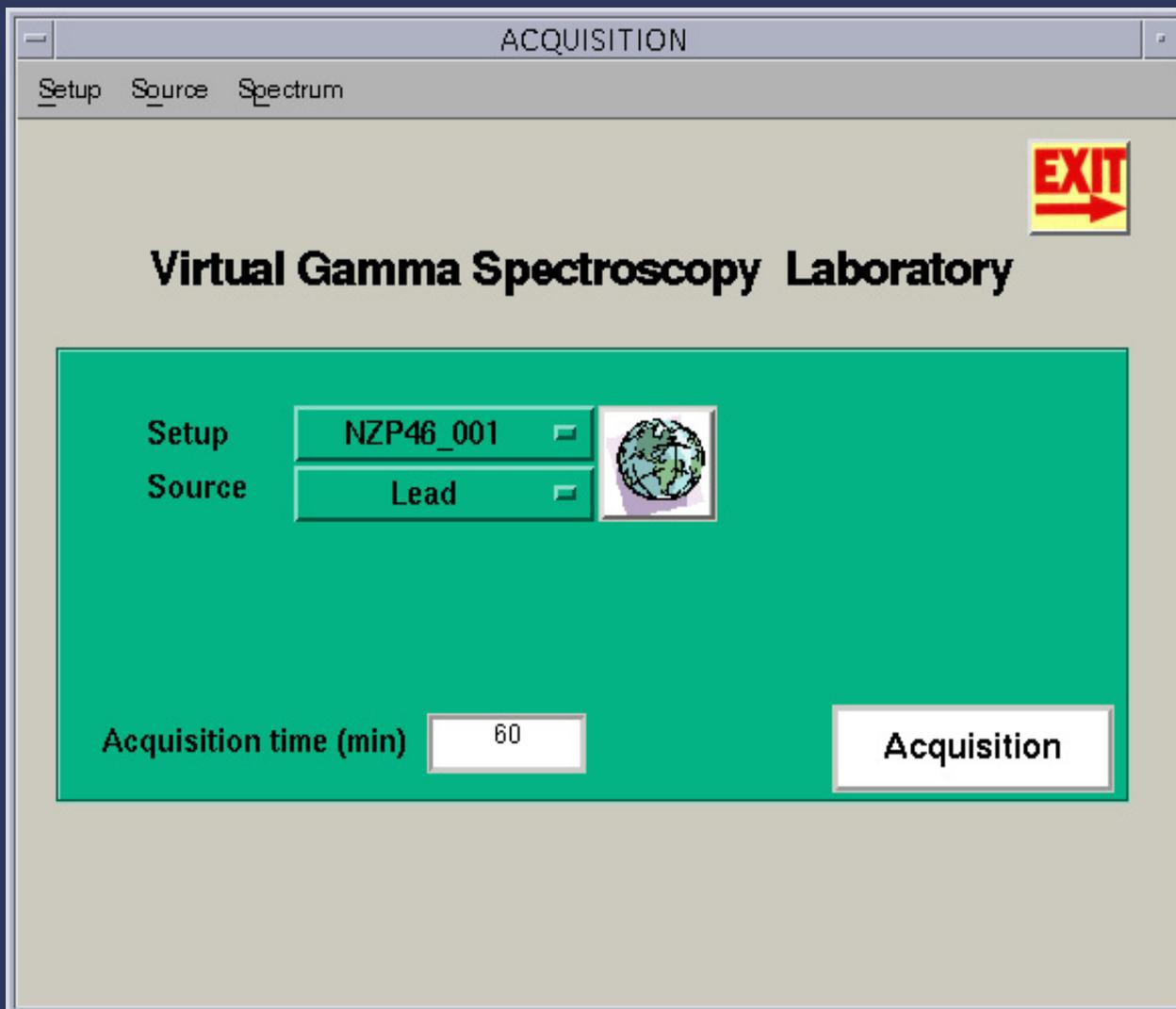


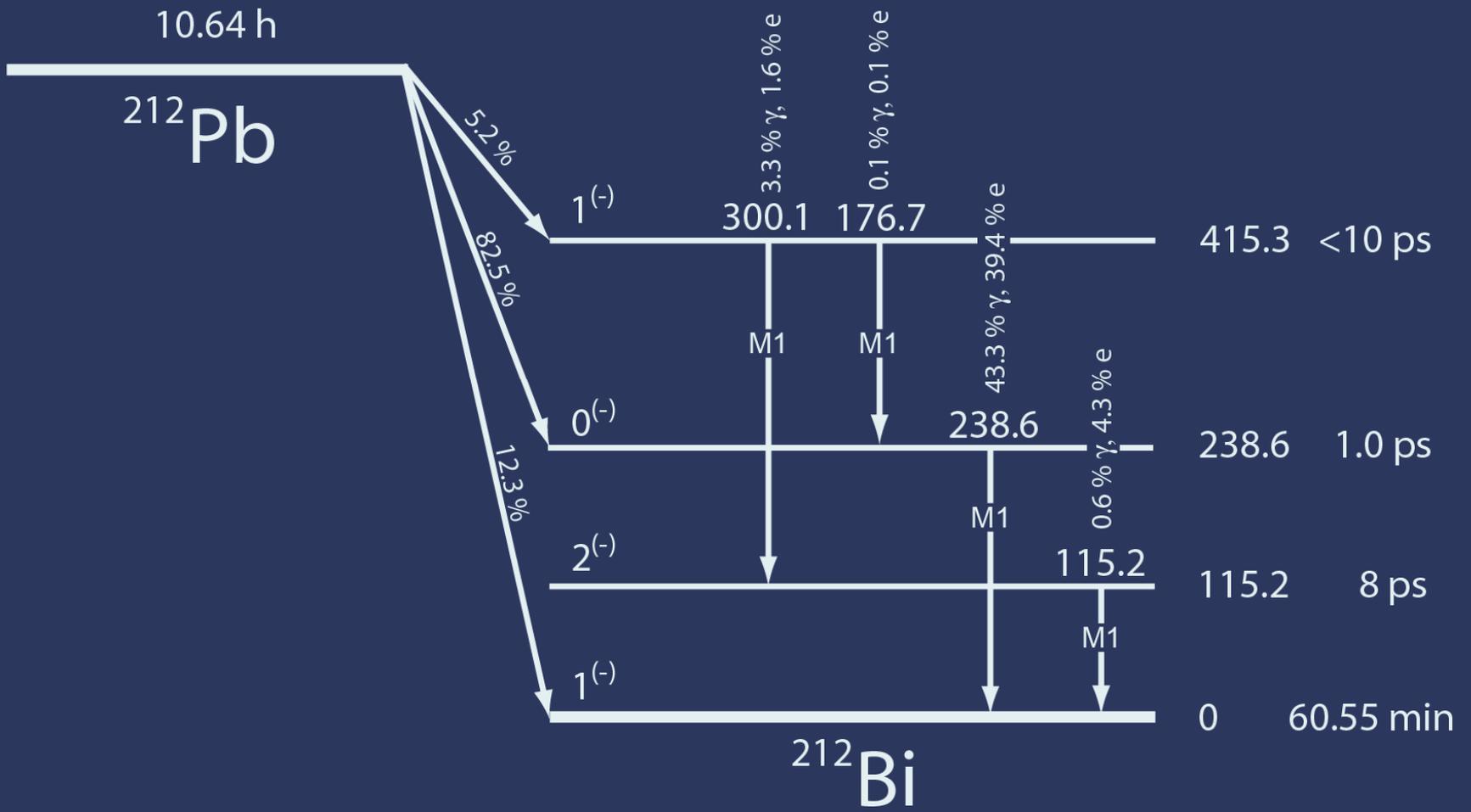
Efficiency table written in file NZP46_001.eff

test

Setup name







Spectrum

Save

Acquisition Time =
5691 sec

Gamma shoot =
10271933

Setup
Source

NZP46_001
Lead



Normal Spectrum

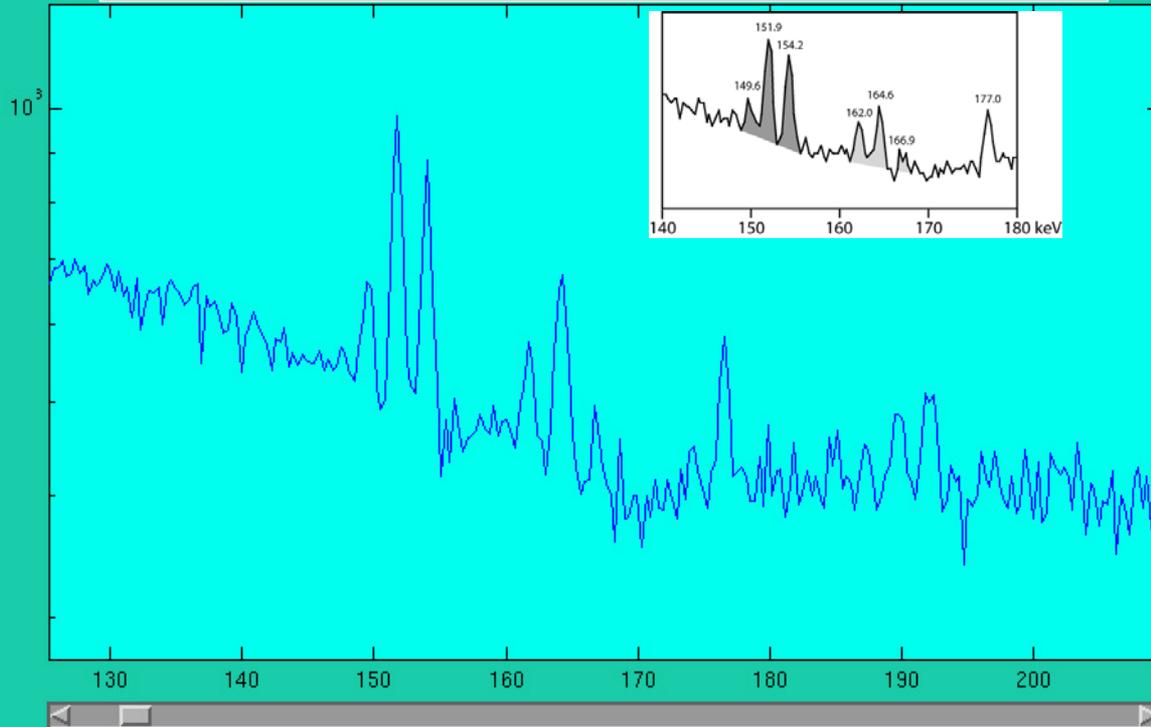
Energy (KeV) =

Channel Count =

?

Summing correction
factor =

+
-



Autoscale

semilog



Spectrum

Save

Acquisition Time =
5691 sec

Gamma shoot =
10271933

Setup
Source

NZP46_001
Lead



Normal Spectrum

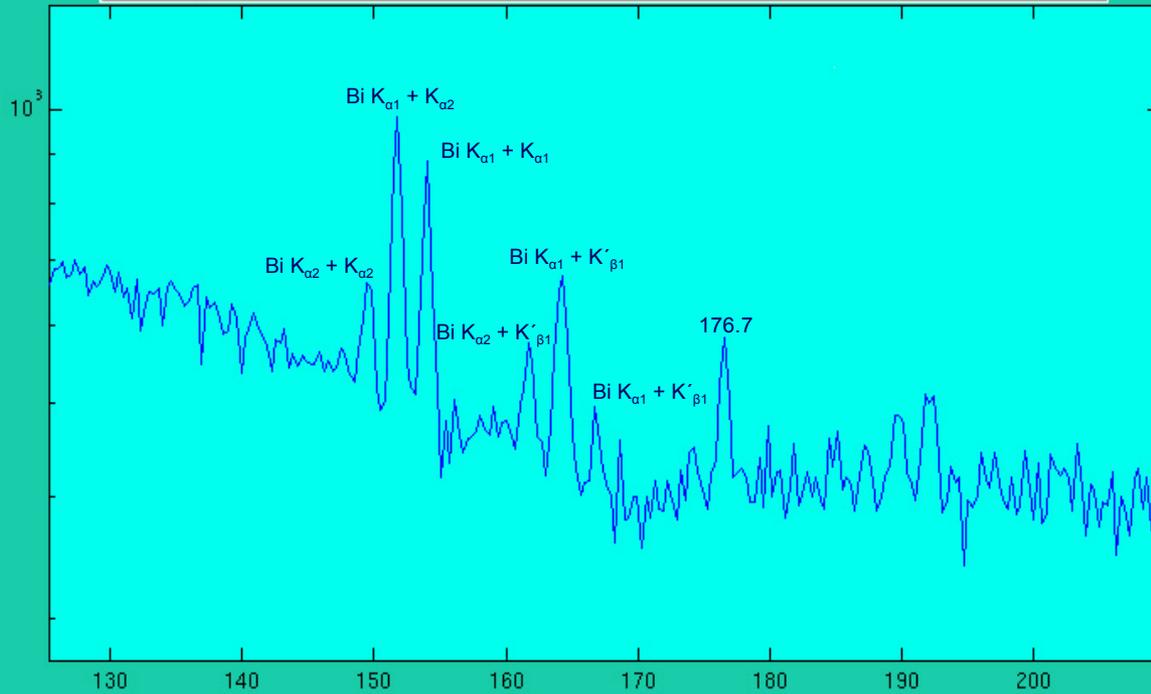
Energy (KeV) =

Channel Count =

?

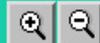
Summing correction
factor =

+
-



Autoscale

semilog



Spectrum

Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead

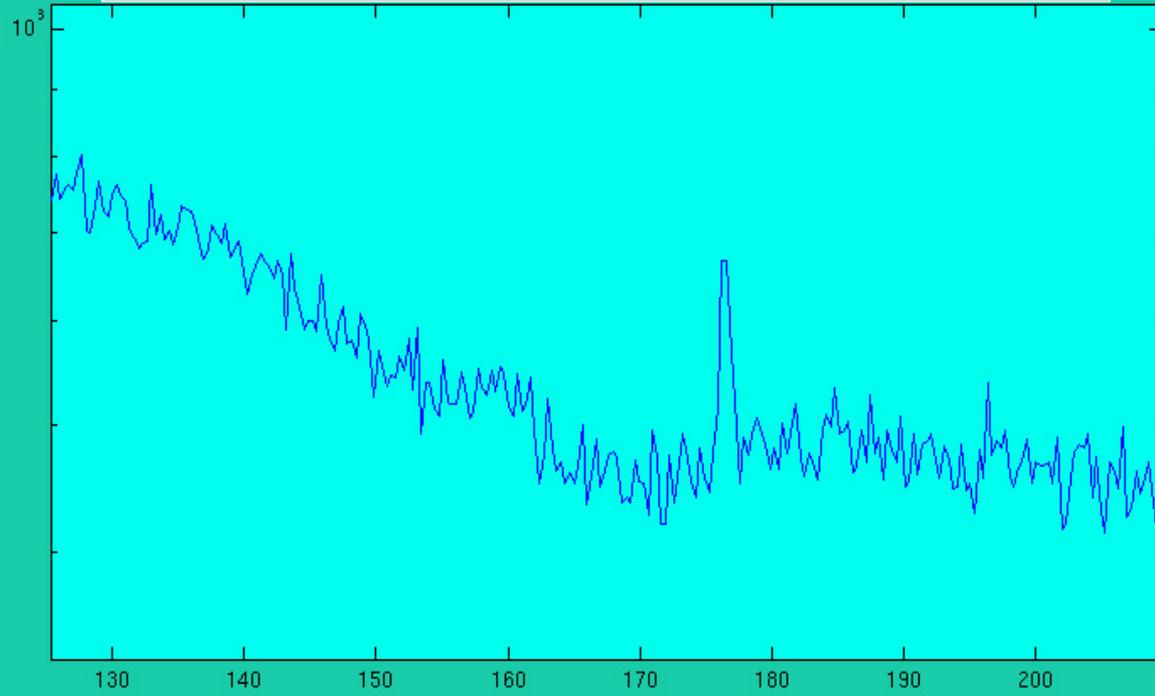


Normal Spectrum with decay coincidence summation turned off

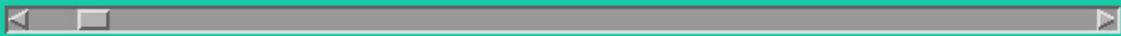
Energy (KeV) =
Channel Count =

?

Summing correction
factor =
+
-



Autoscale
 semilog



Spectrum

Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead



Normal Spectrum with sharp resolution (i.e. with FWHM = 1 channel)

Energy (KeV) =

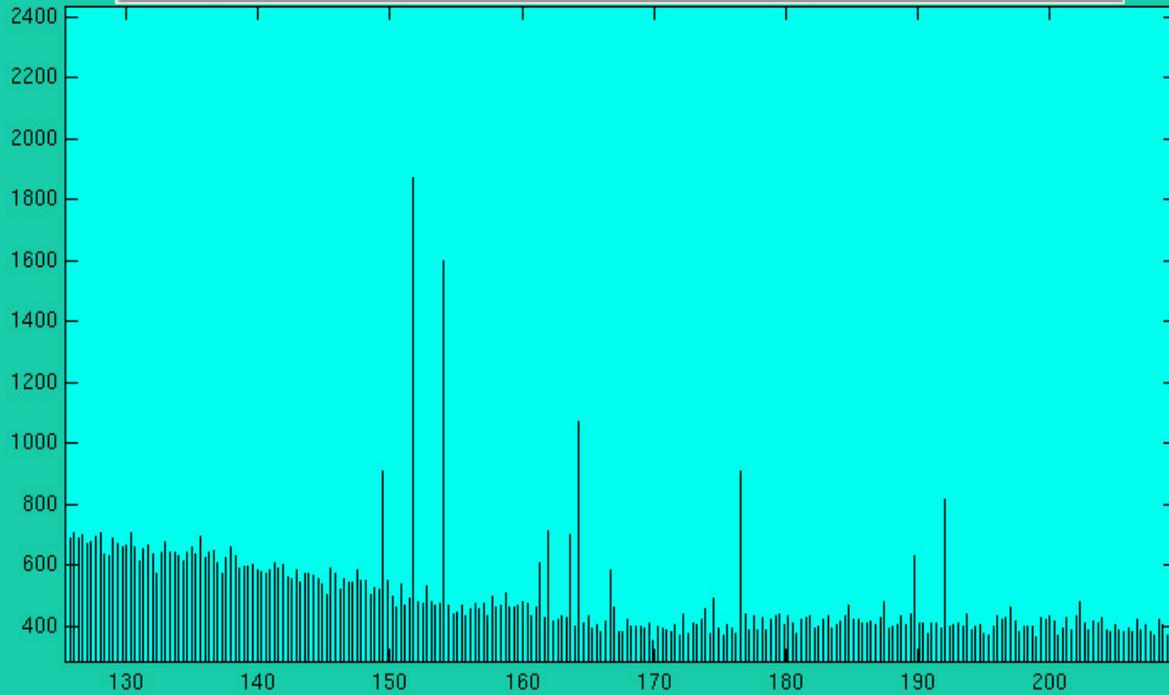
Channel Count =

?

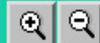
Summing correction
factor =

+

-



Autoscale



Spectrum

Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead



Normal Spectrum with decay coincidence summation turned off and with sharp resolution

Energy (KeV) =

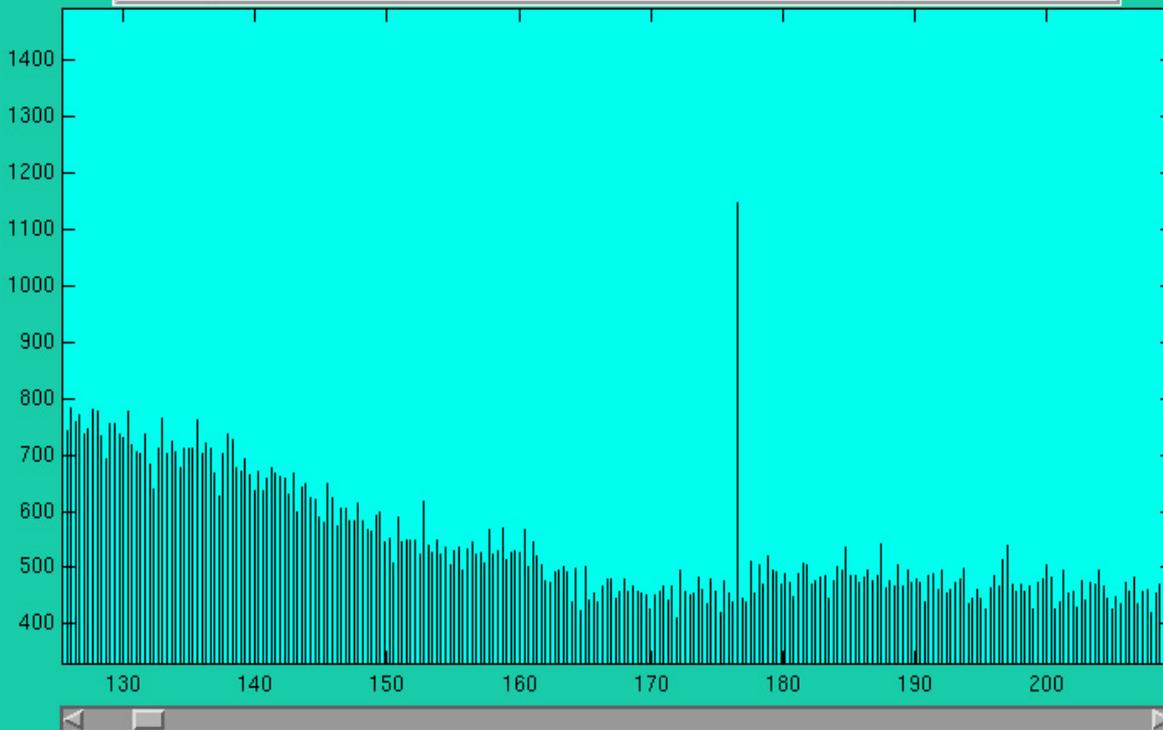
Channel Count =

?

Summing correction
factor =

+

-



Autoscale

Spectrum

Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead



Peaks only (no background line, escape peaks not currently implemented)

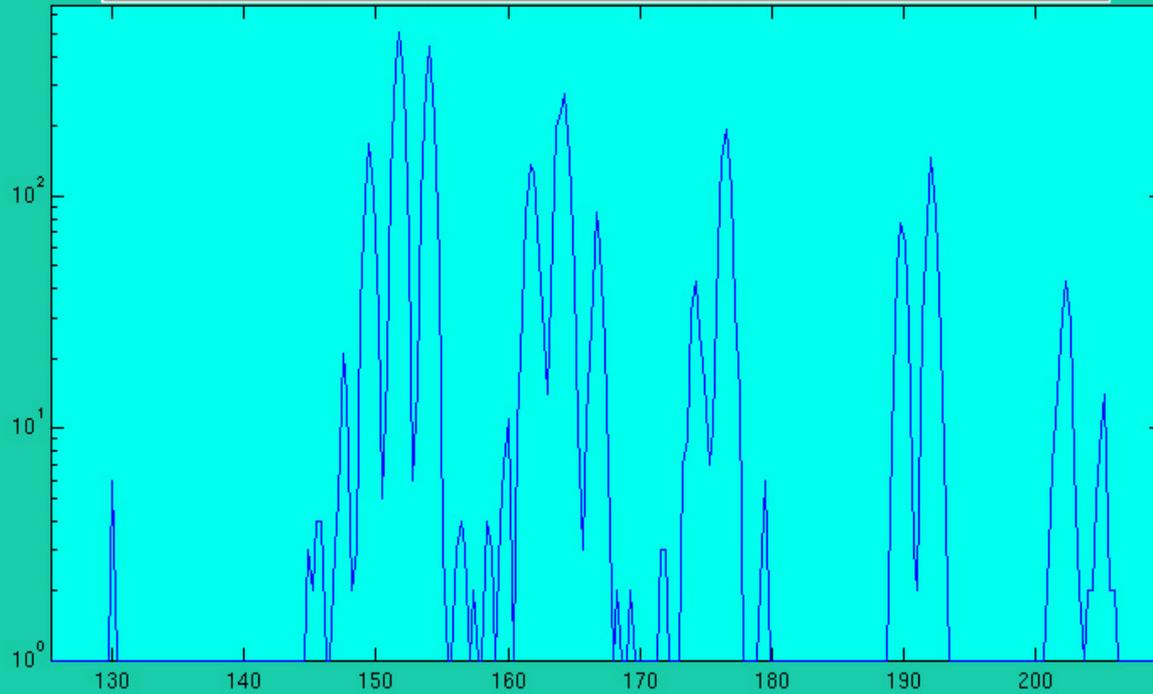
Energy (KeV) =

Channel Count =

?

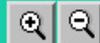
Summing correction
factor =

+
-



Autoscale

semilog



Save

Acquisition Time =
5691 sec

Gamma shoot =
10271933

Setup
Source

NZP46_001
Lead



Peaks only with decay coincidence summation turned off

Energy (KeV) =

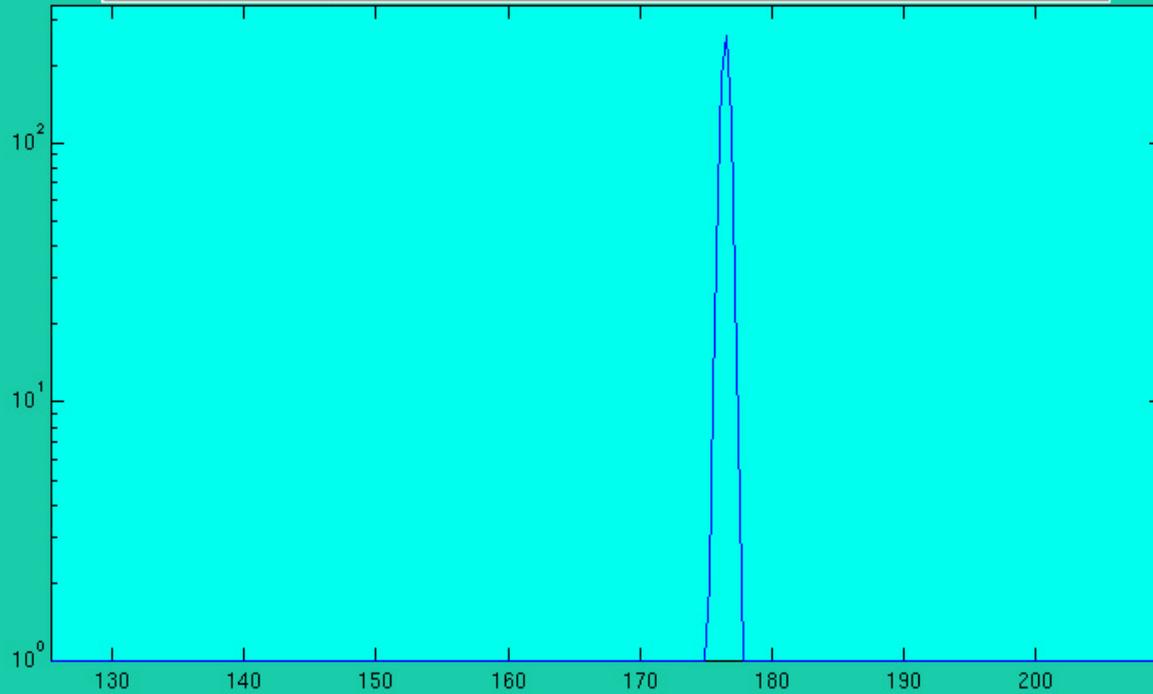
Channel Count =

?

Summing correction
factor =

+

-



Autoscale

semilog



Spectrum

Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead



Peaks only with sharp resolution

Energy (KeV) =

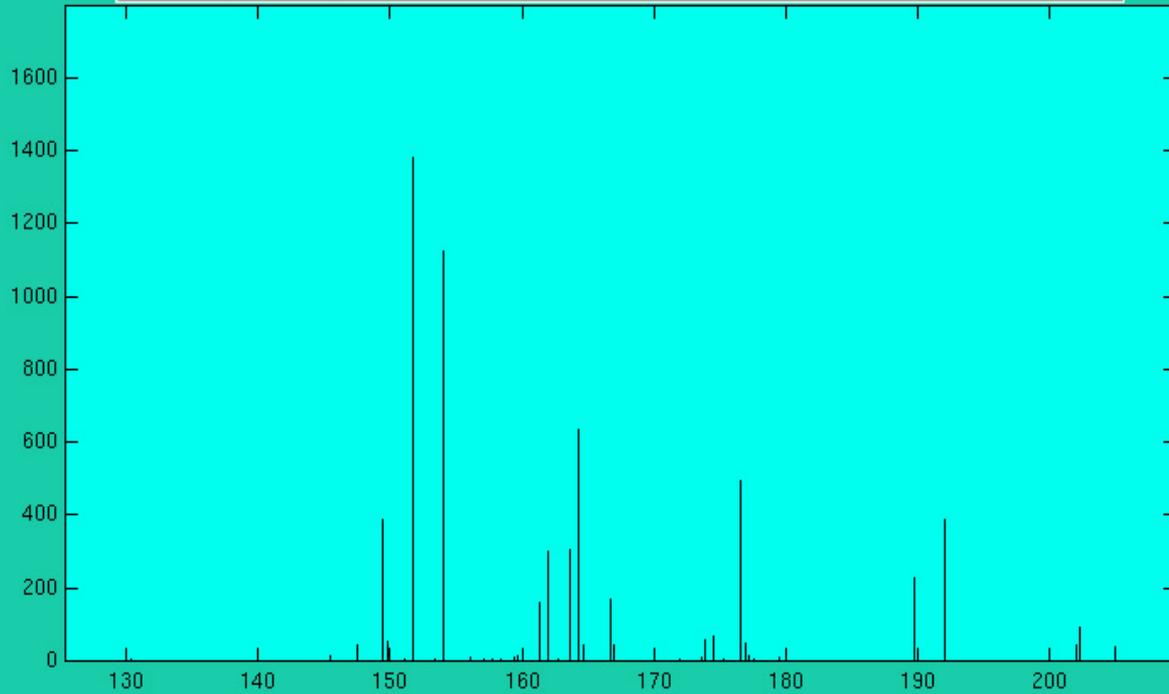
Channel Count =

?

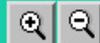
Summing correction
factor =

+

-



Autoscale



Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead



Peaks only with decay coincidence summation turned off and with sharp resolution

Energy (KeV) =

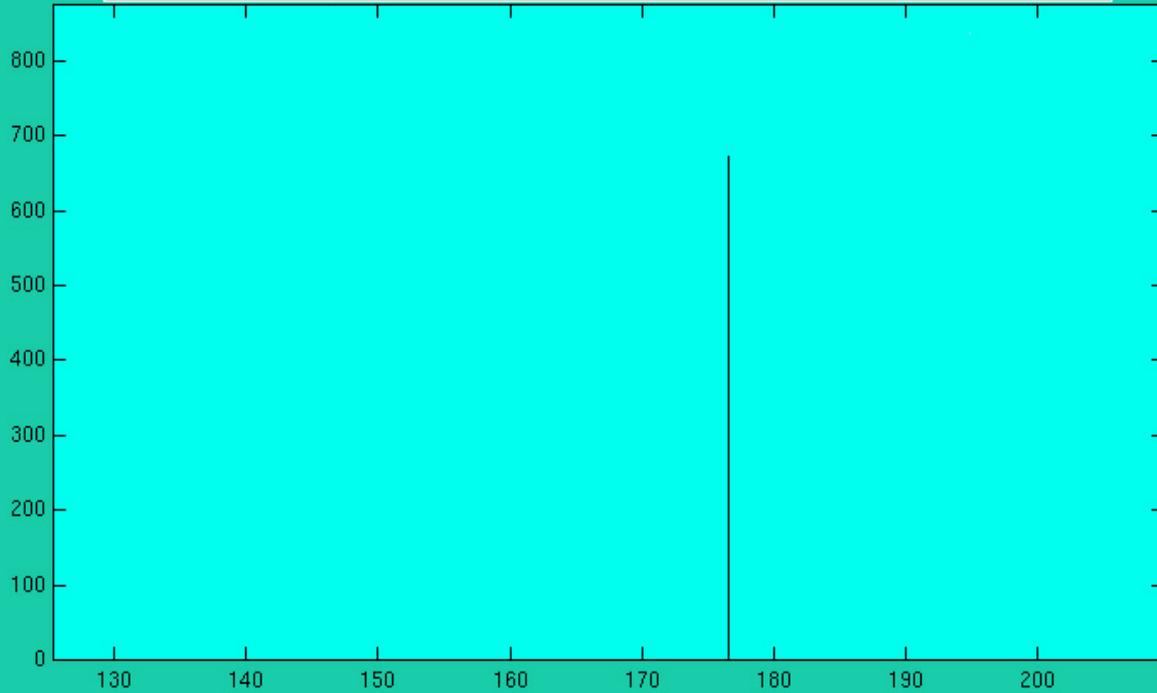
Channel Count =

?

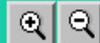
Summing correction
factor =

+

-



Autoscale



Spectrum

Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead



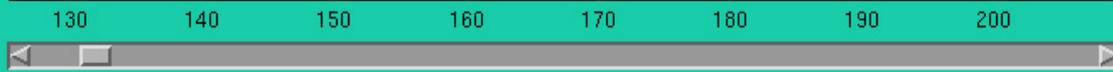
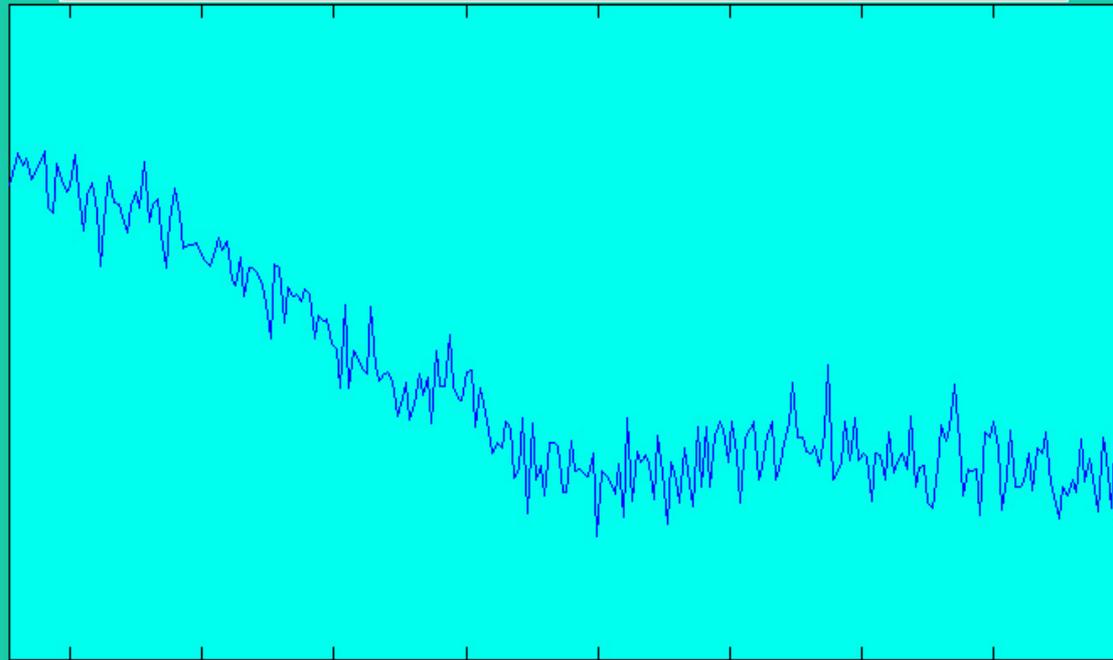
Background line (= Spectrum 1 - Spectrum 5 including escape peaks)

Energy (KeV) =

Channel Count =

Autoscale

semilog



Save

Acquisition Time =
5691 sec

Gammas shoot =
10271933

Setup
Source

NZP46_001
Lead

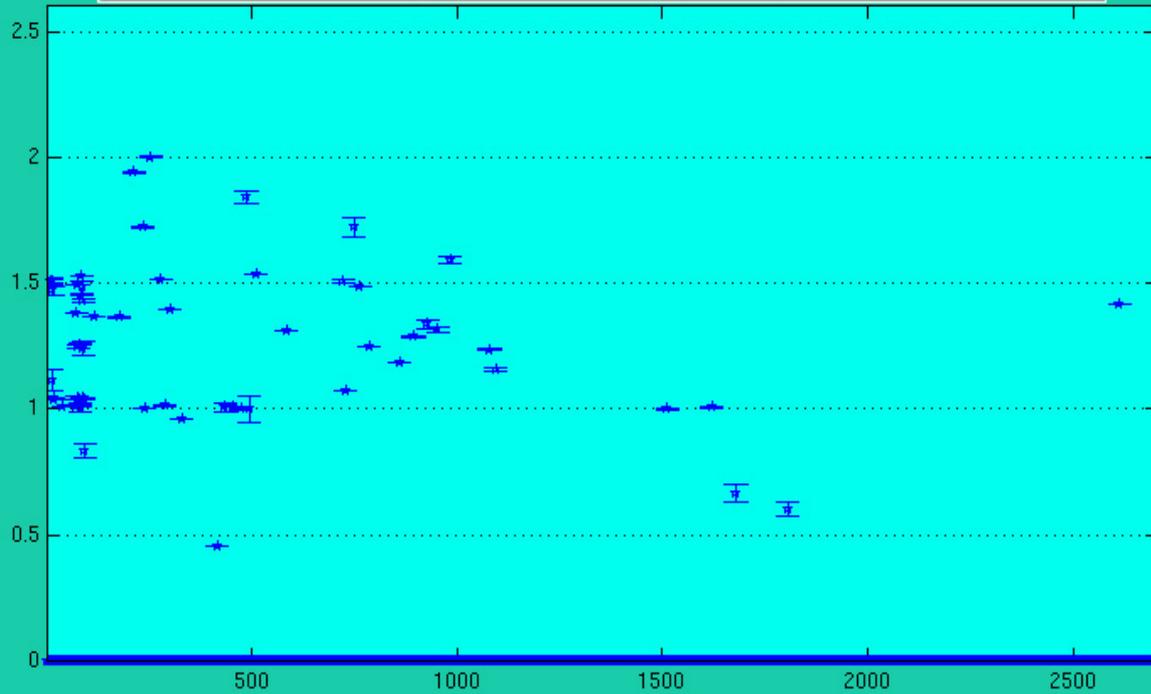


Coincidence Correction Factor (Spectrum 8 divided by Spectrum 7)

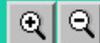
Energy (KeV) =

Summing correction
factor =

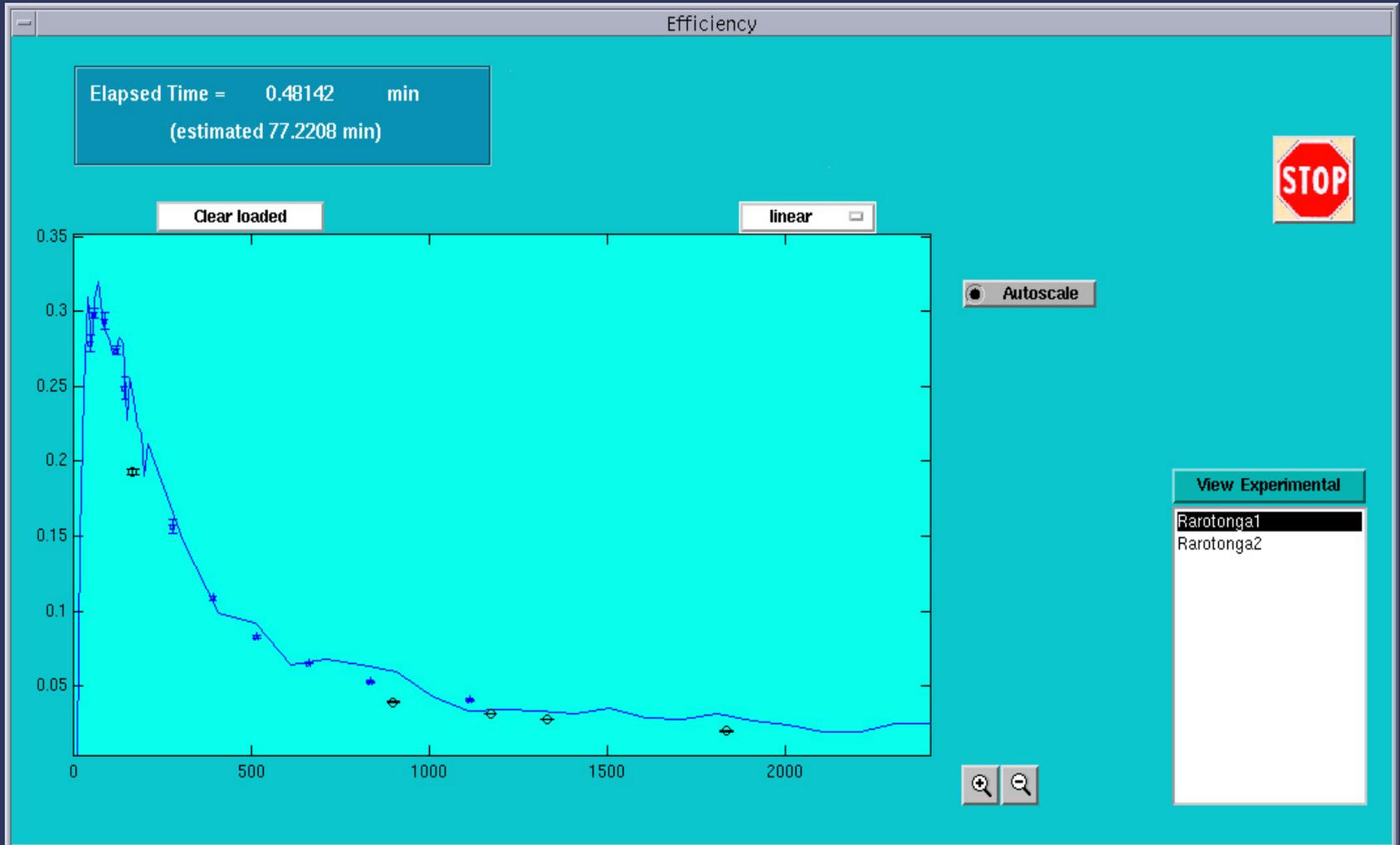
+
-



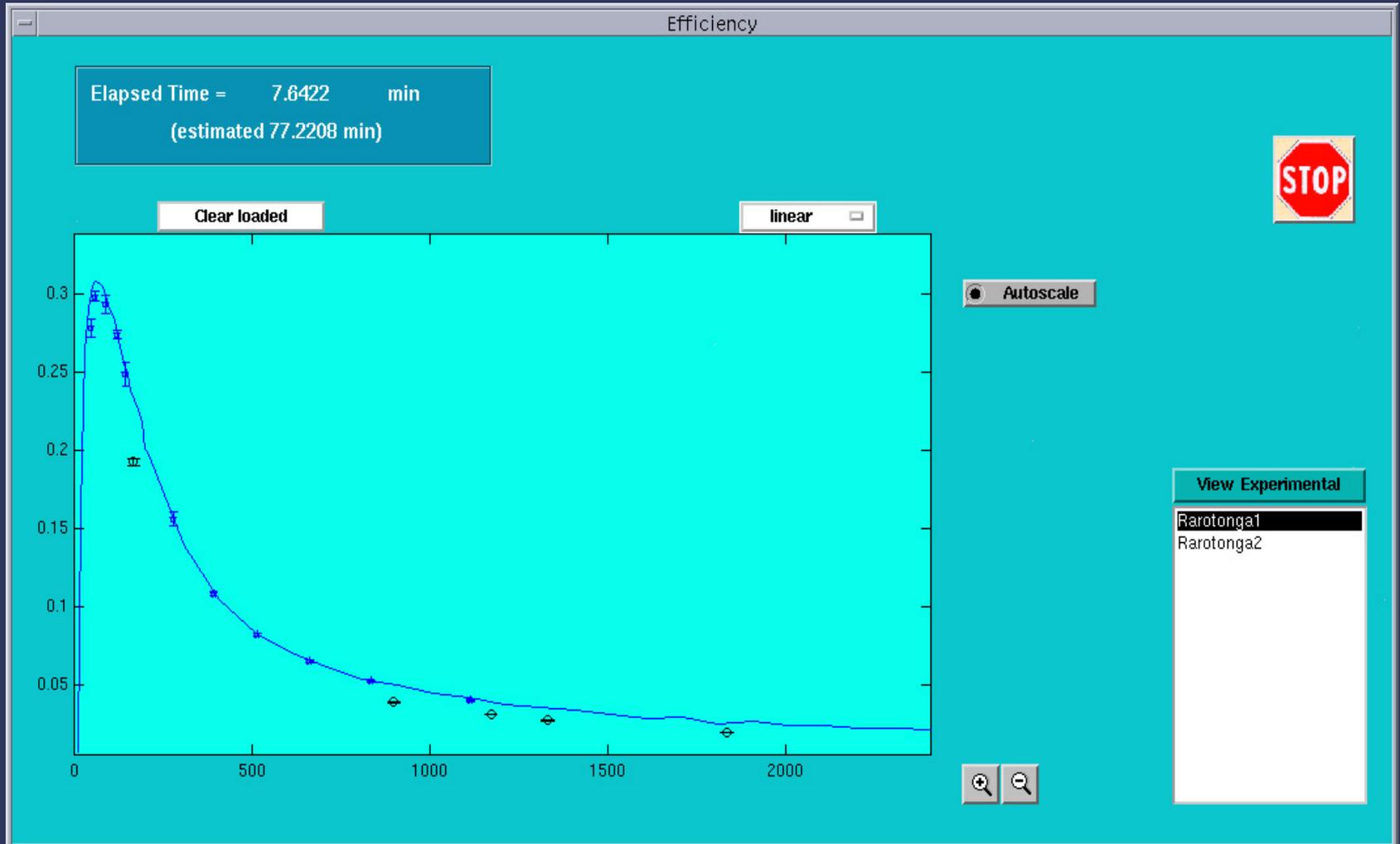
Autoscale



Efficiency curve after half a minute



Efficiency curve after 7.6 minutes



Running an experiment

Source Definition
Database (ENSDF)

Source Content Definition

EXIT

Bq

Refresh **->**

Bq

Remove **CLEAR**

Existing Sources

Remove selected

Lead

Source content name

Data base creation

File Edit View Insert Tools Window Help

212pb Beta-decay 212bi

415.2 300.00 60

238.632 129.5

115.168

? the data file for this element is already present do you want to download the newest version from the nndc server

Yes **No** **Cancel**

212Bi

<http://dbserv.pnpi.spb.ru/nrd/temp/taiblist/ensdf/ensdf2>



ACQUISITION

Setup Source Spectrum

EXIT

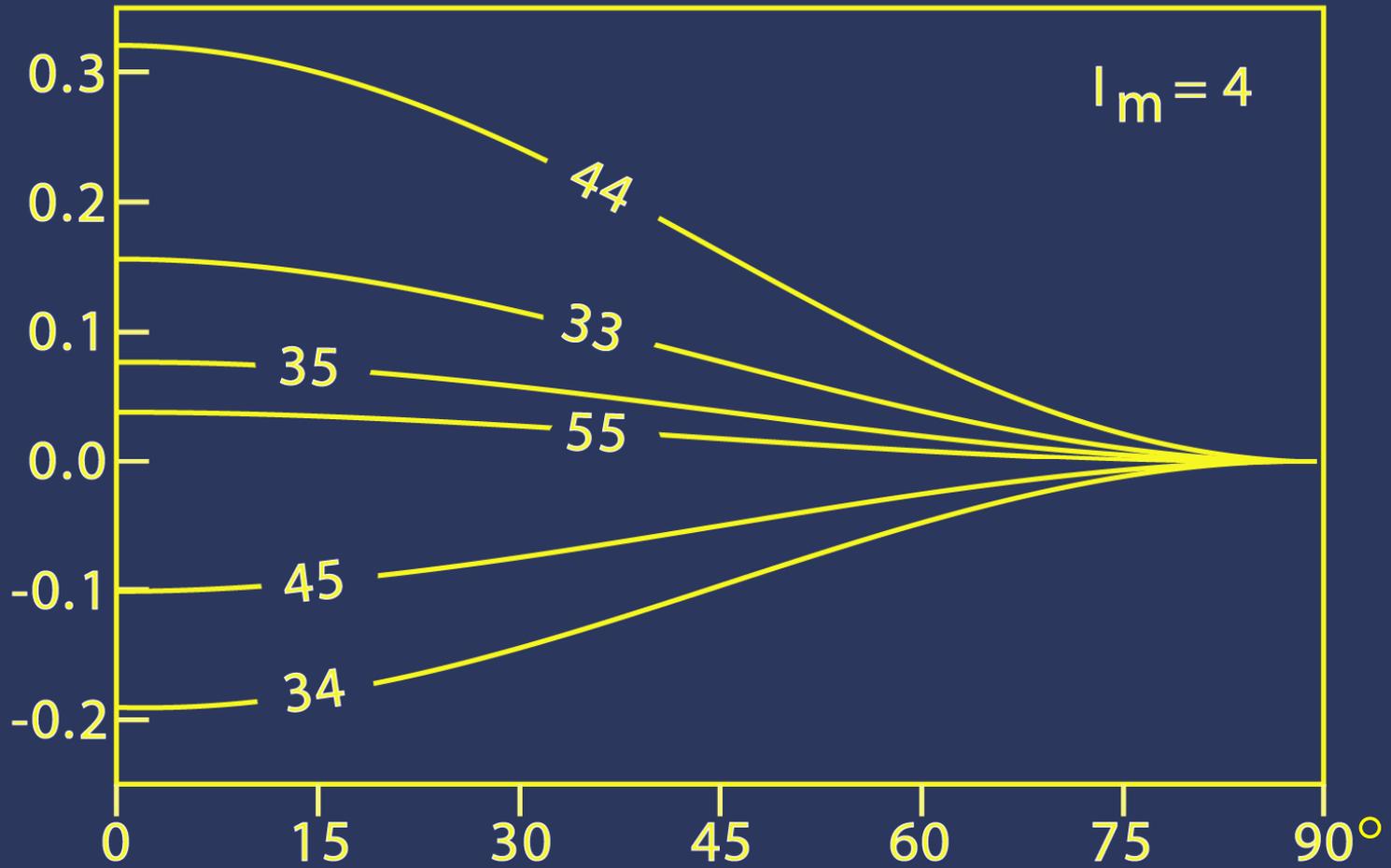
Virtual Gamma Spectroscopy Laboratory

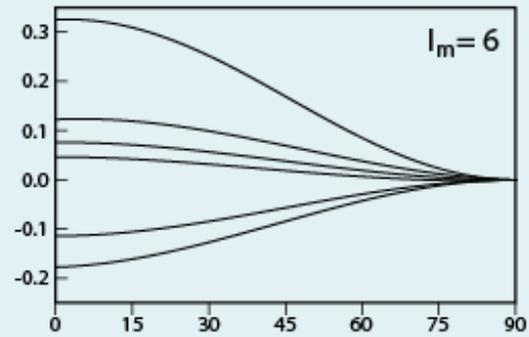
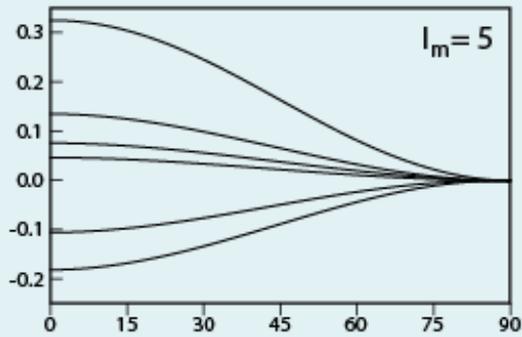
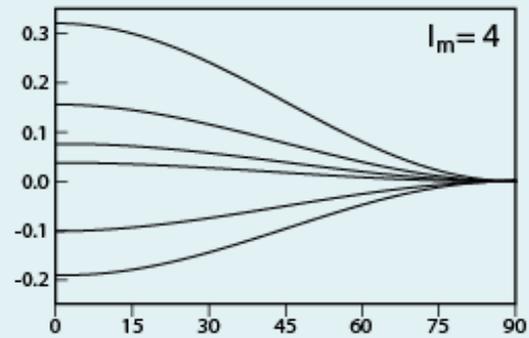
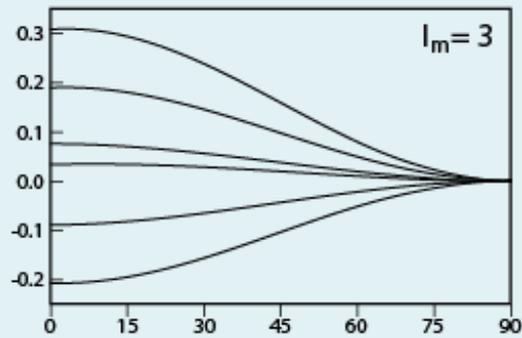
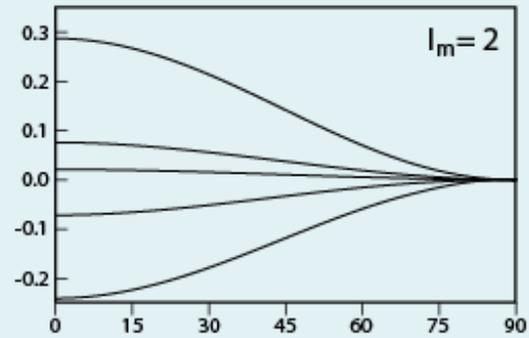
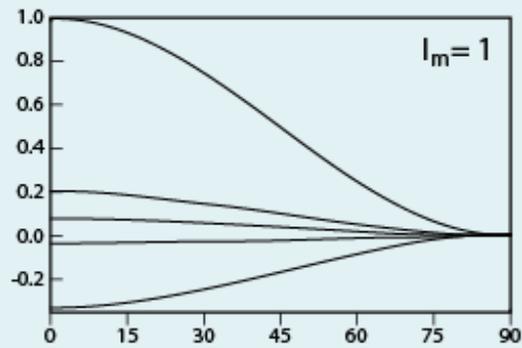
Setup

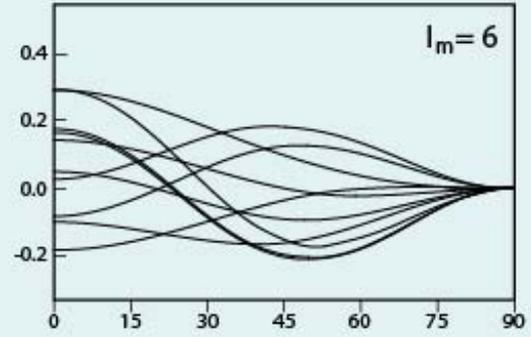
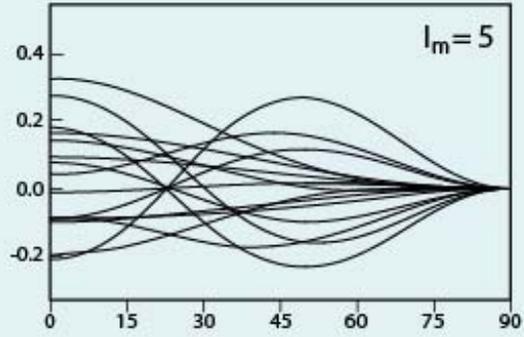
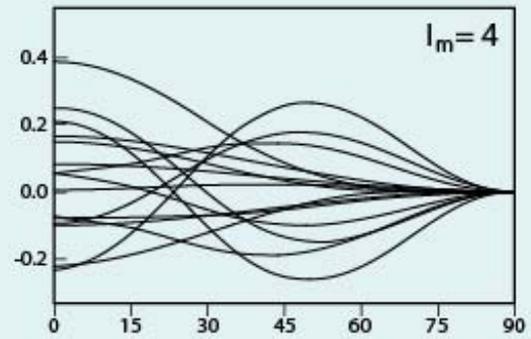
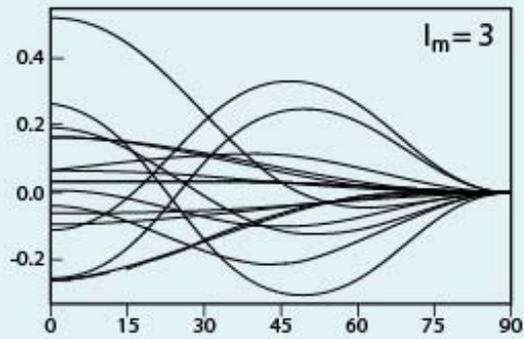
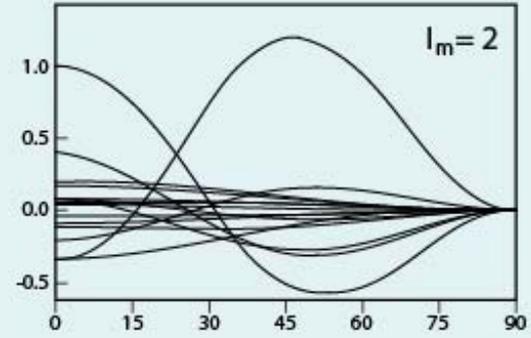
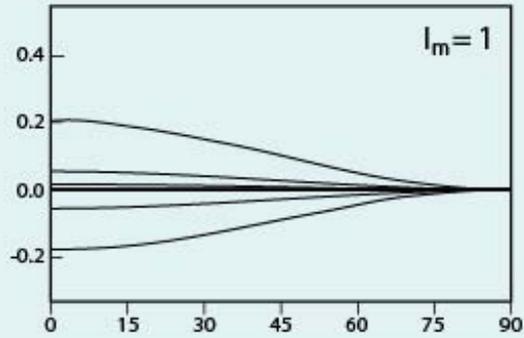
Source

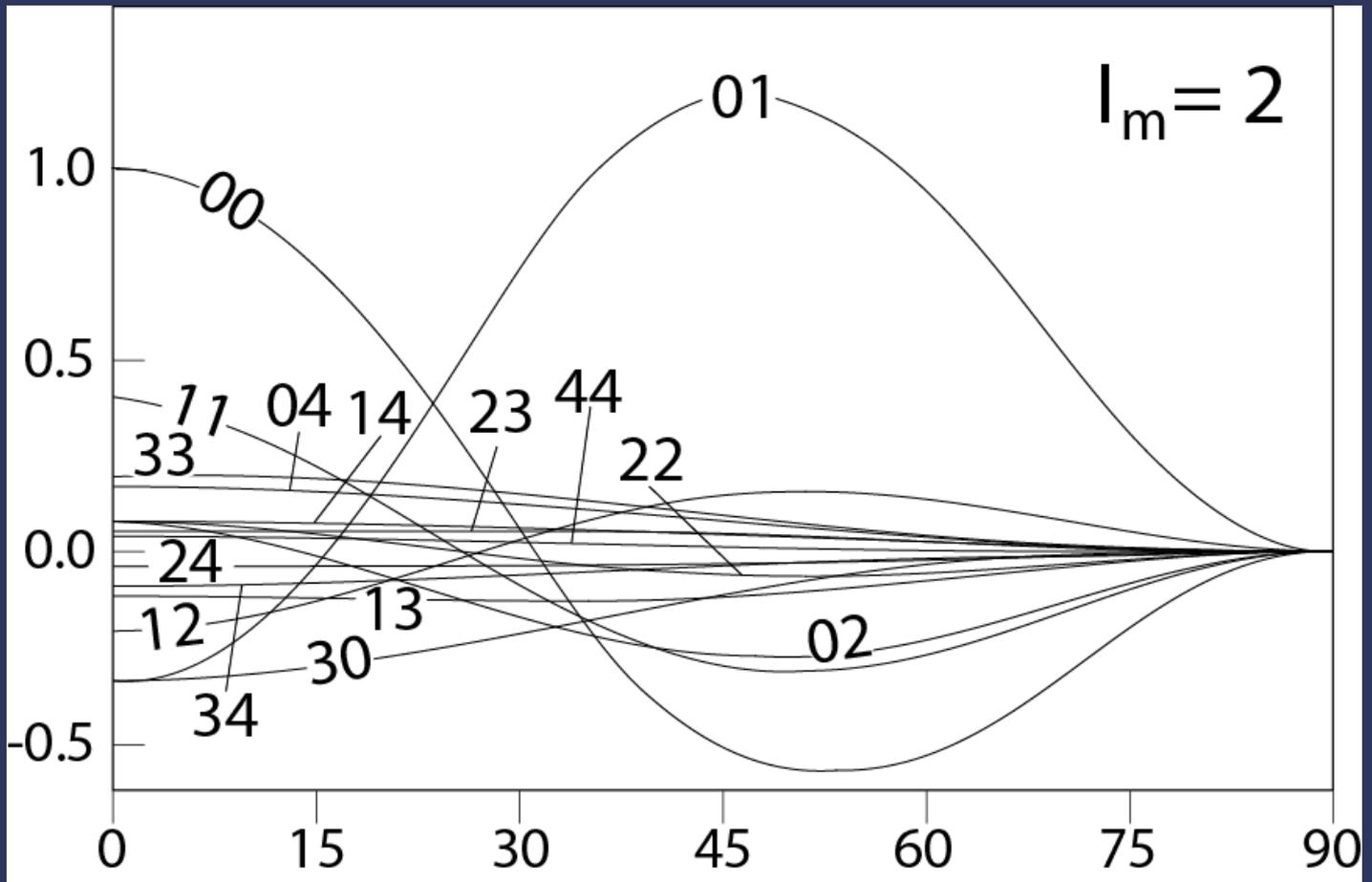
Acquisition time (min)

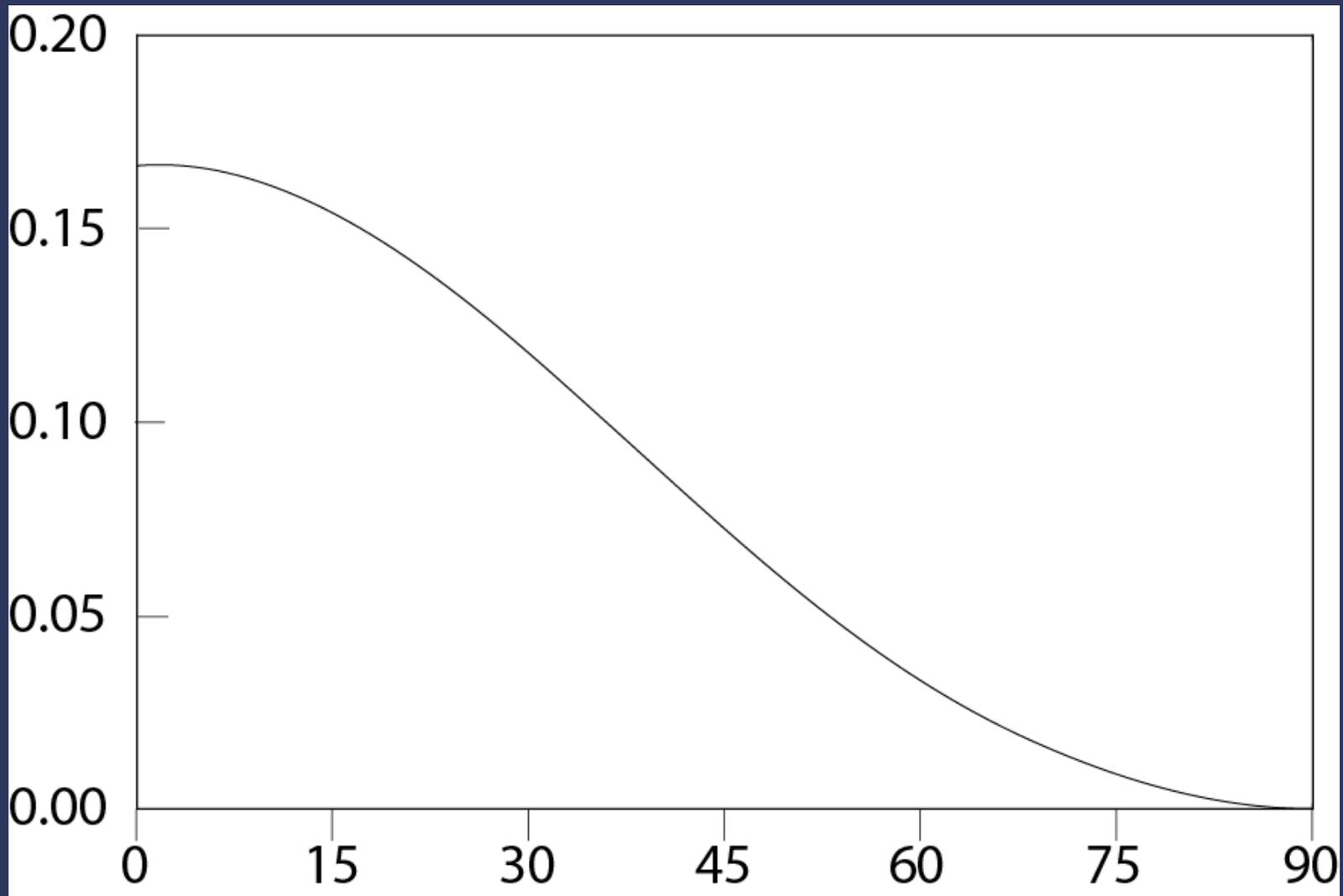
Acquisition











A Decent Currie

Applying the Currie SCA to gamma spectra

$\alpha = \text{e.g. } 5\%$

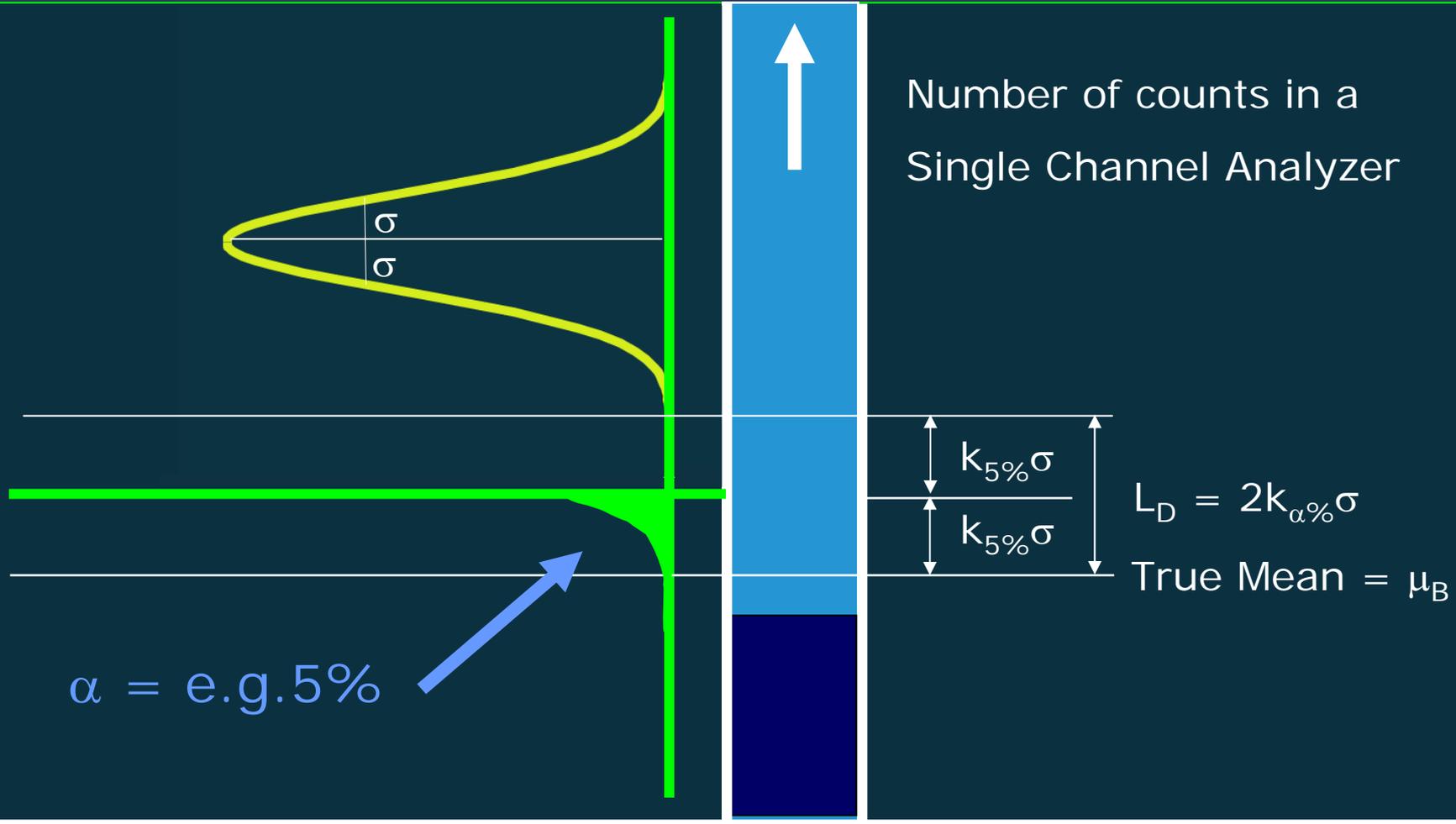


Number of counts in a Single Channel Analyzer

$k_{5\%}\sigma$

$$L_C = k_{\alpha\%}\sigma$$
$$\text{True Mean} = \mu_B$$

Applying the Currie SCA to gamma spectra



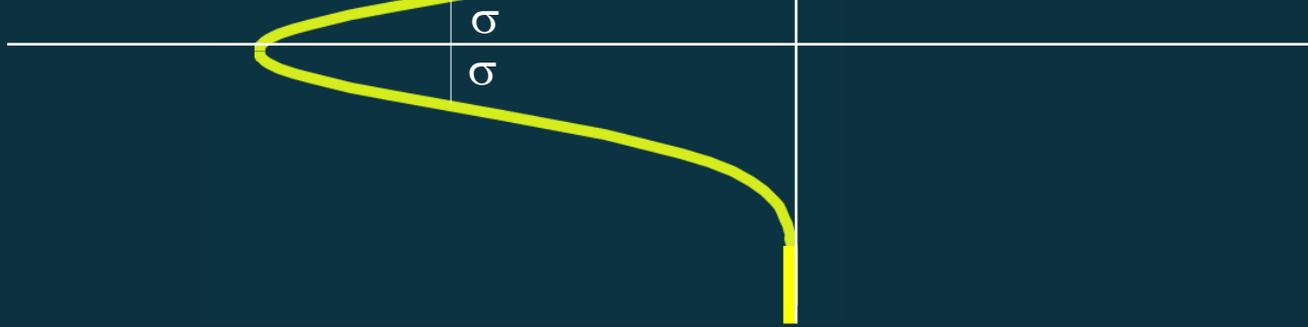
Lloyd Currie's classical paper from 1968



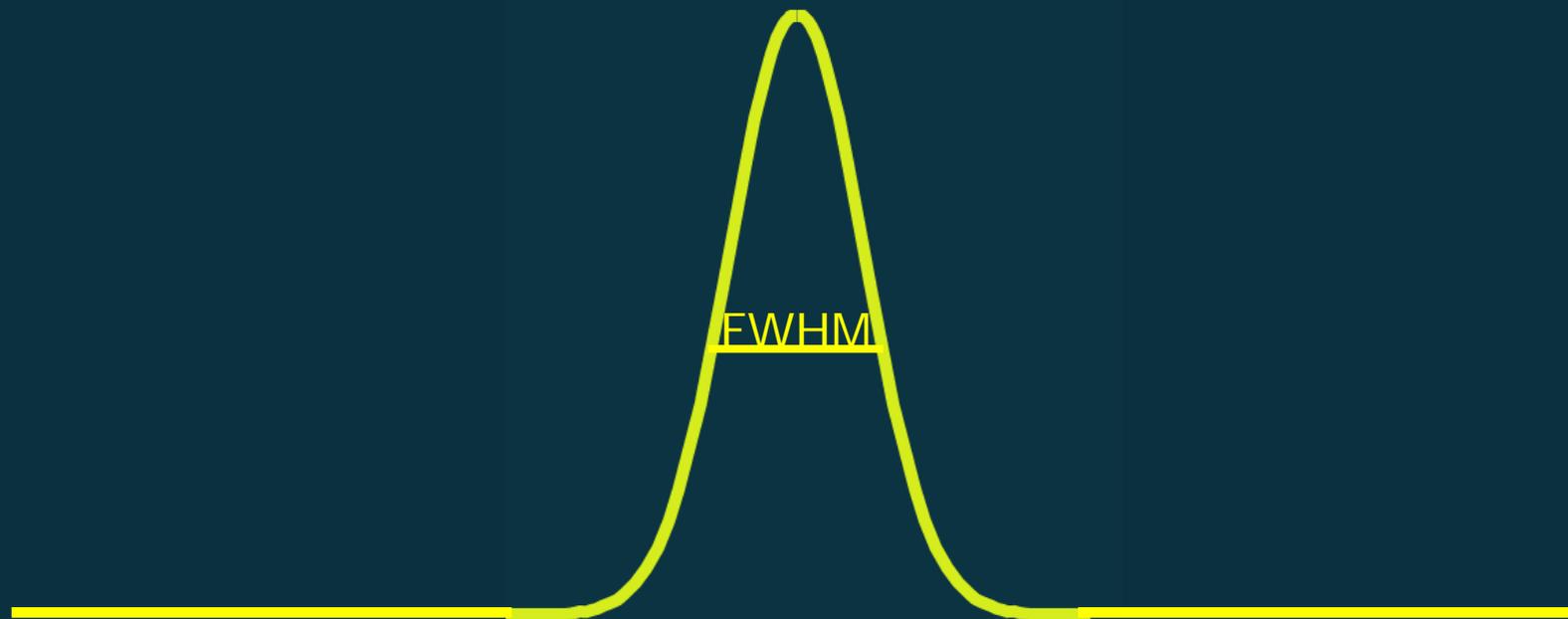
$$L_C = k \sqrt{\mu_B (1 + 1/m)}$$

$$L_D = k^2 + 2L_C$$

Applying the Currie SCA to a gamma spectrum peak



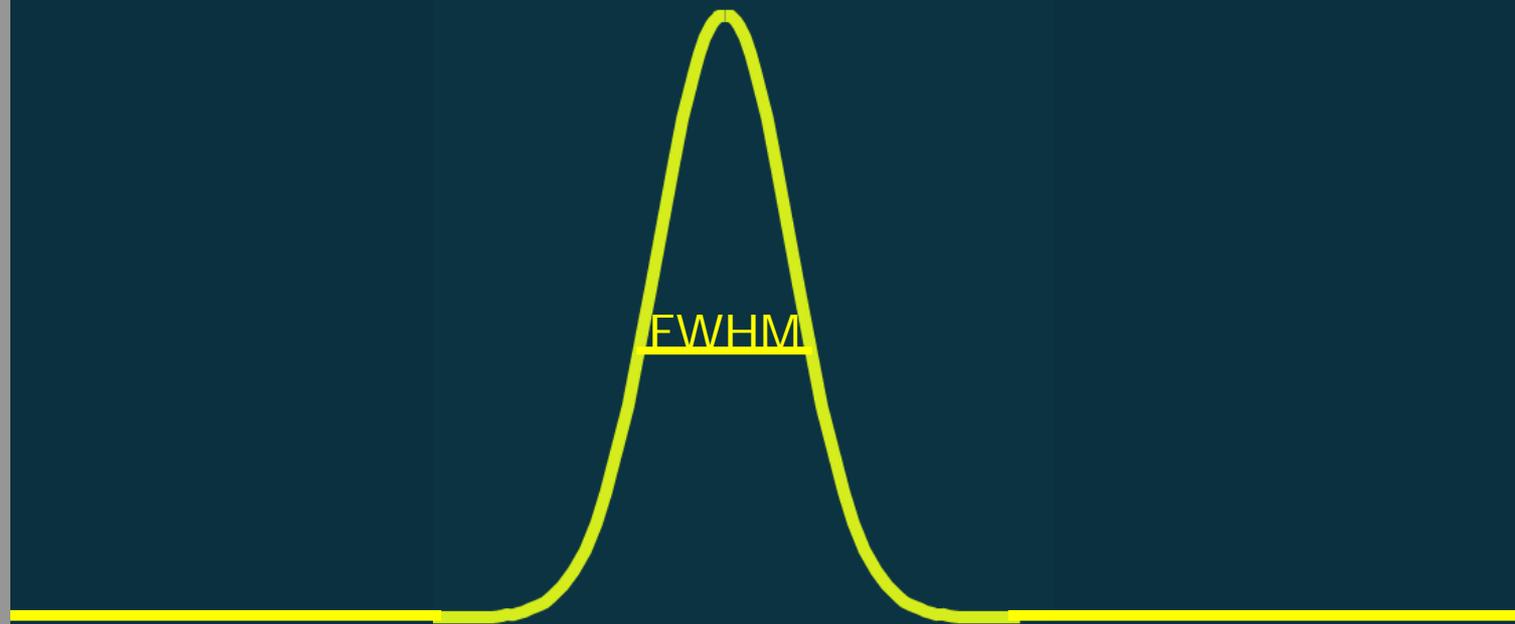
Applying the Currie SCA to a gamma spectrum peak



Applying the Currie SCA to a gamma spectrum peak



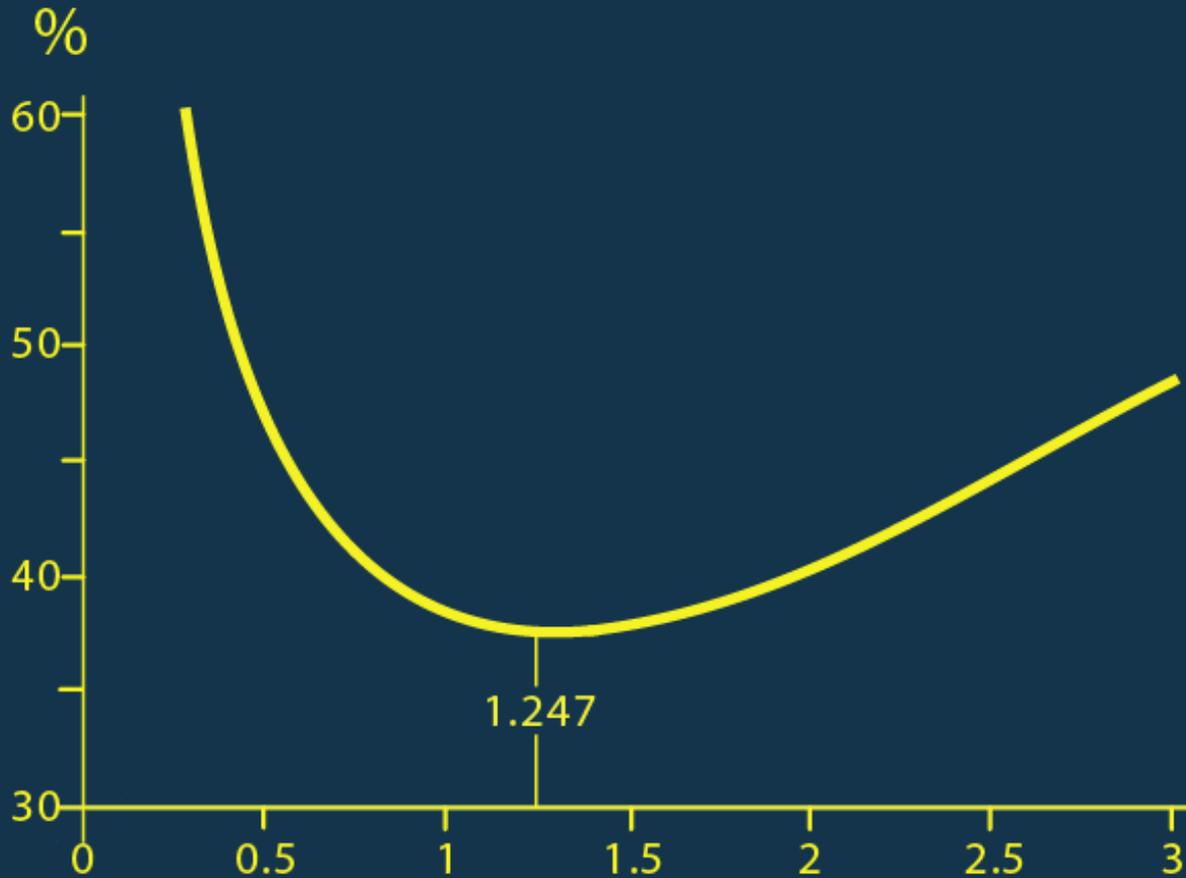
Applying the Currie SCA to a gamma spectrum peak



Applying the Currie SCA to a gamma spectrum peak

Commonly
used width:
2.5 FWHM





Calculating the channel-width that gives the lowest relative uncertainty

$$L_C = k \sqrt{\mu_B (1 + 1/m)}$$

With the SCA-width (and thus μ_B) reduced by a factor of 2 and m considered large instead of 1,

L_C is reduced by a factor of 2.

This can be compensated by doubling k , which is equivalent to reducing α , the risks for errors of the first and second kind, from e.g. 5% to 0.05%.

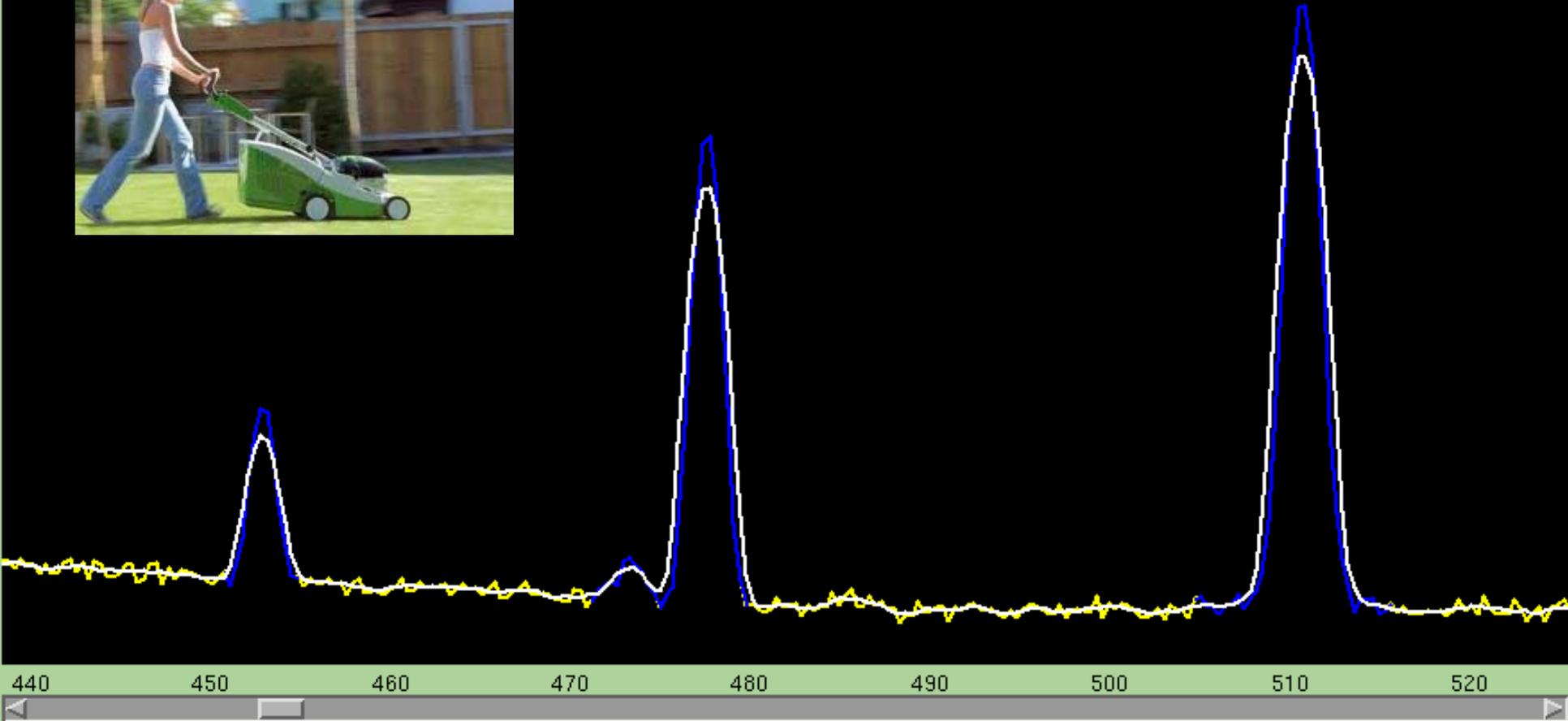
$$\text{Still } L_D = 2L_C$$

This is much more consistent with experience, as we normally see just around five false positives in a 8192 channel spectrum and not hundreds of them which we would if the risk was really 5%.

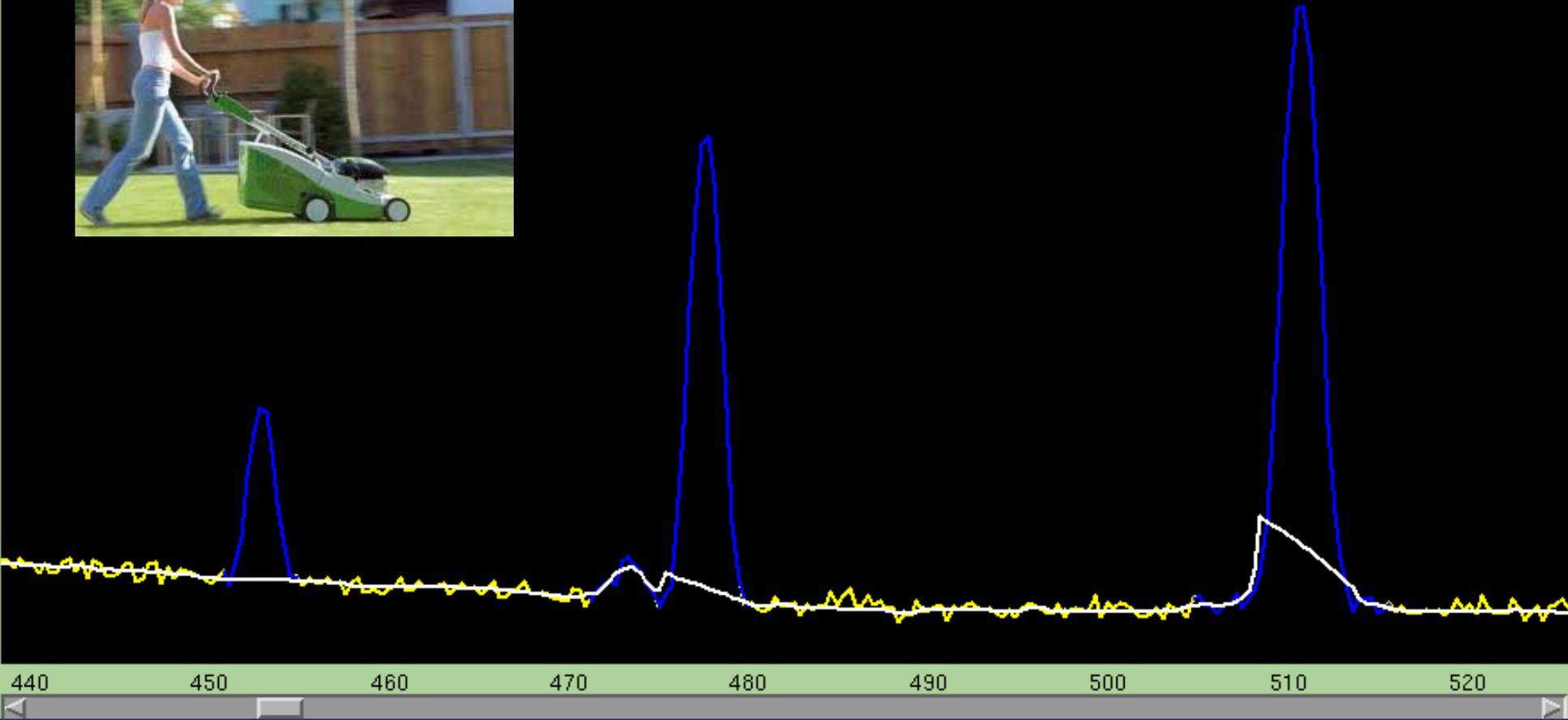
So how do we now utilize this in a spectrum where we a priori don't know where all peaks are?

1. Pick all clear and nice peaks away from the spectrum.
2. We need a good background line and we can now get one by smoothing the picked spectrum.
3. Assume all channels are the centroid of a small peak.
4. Add the channels within 1.25 FWHM of each channel in the spectrum and write the sum into the centroid channel. This involves also partial channels at each end of the SCA.
5. Compare with the L_c criterion. If channel content is $> L_c$ it's part of a peak, otherwise not.

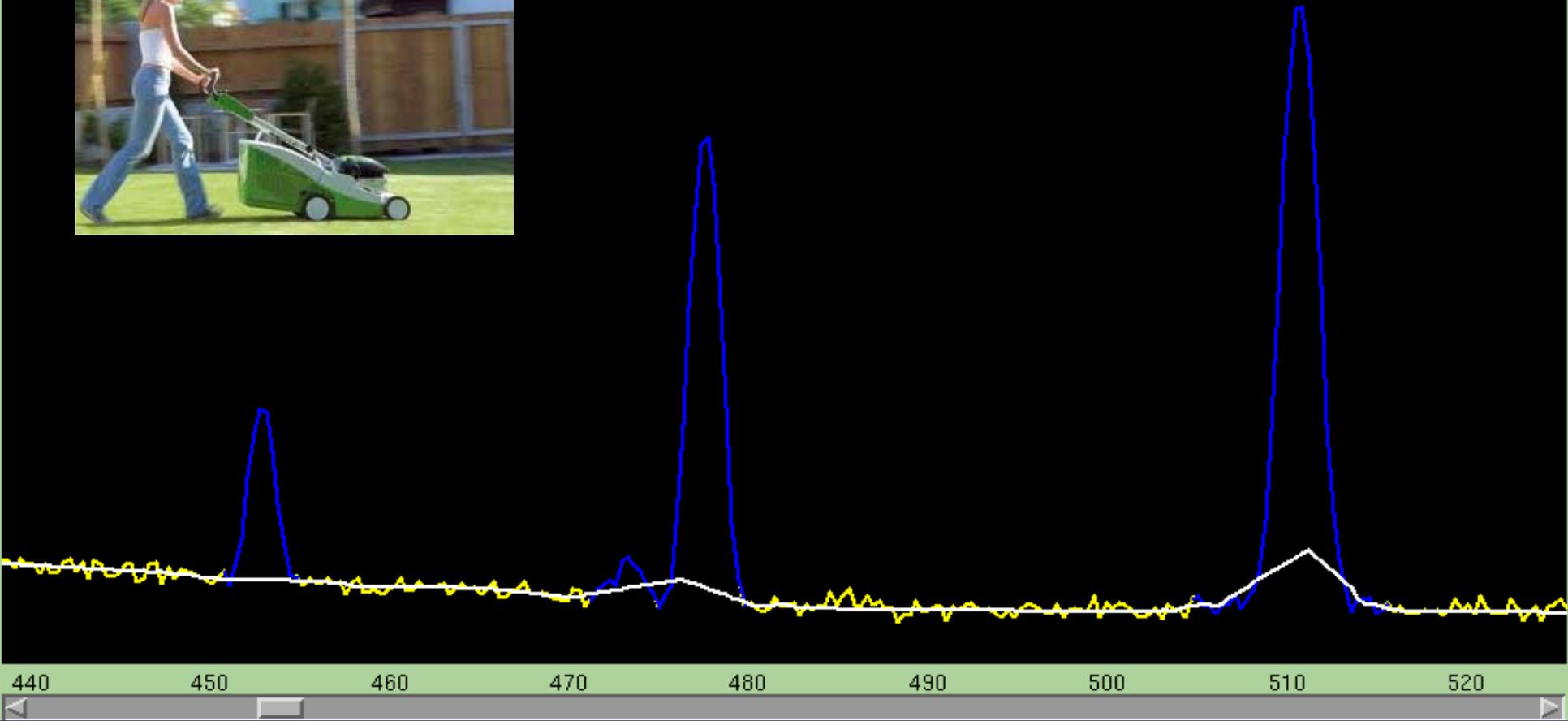
Lawnmower operation



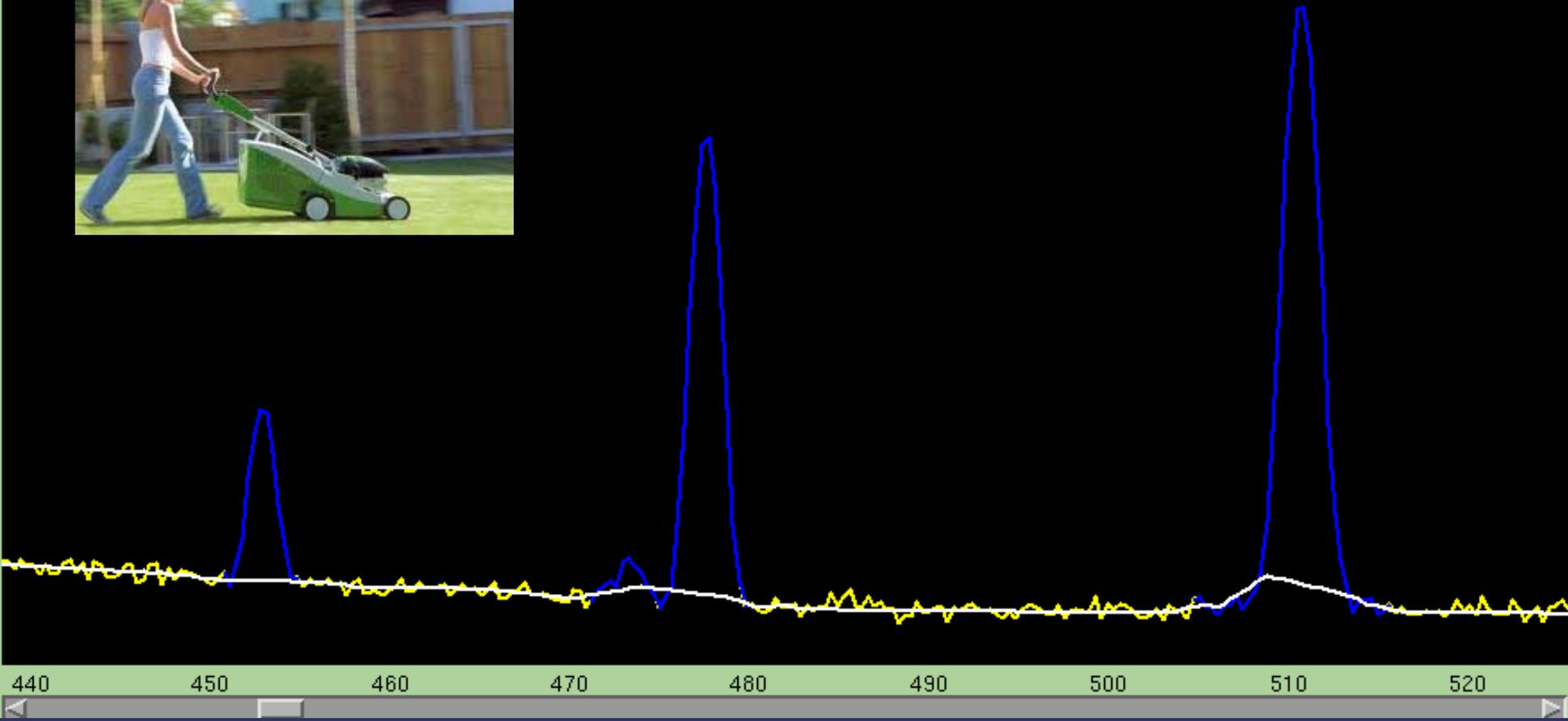
Lawnmower operation



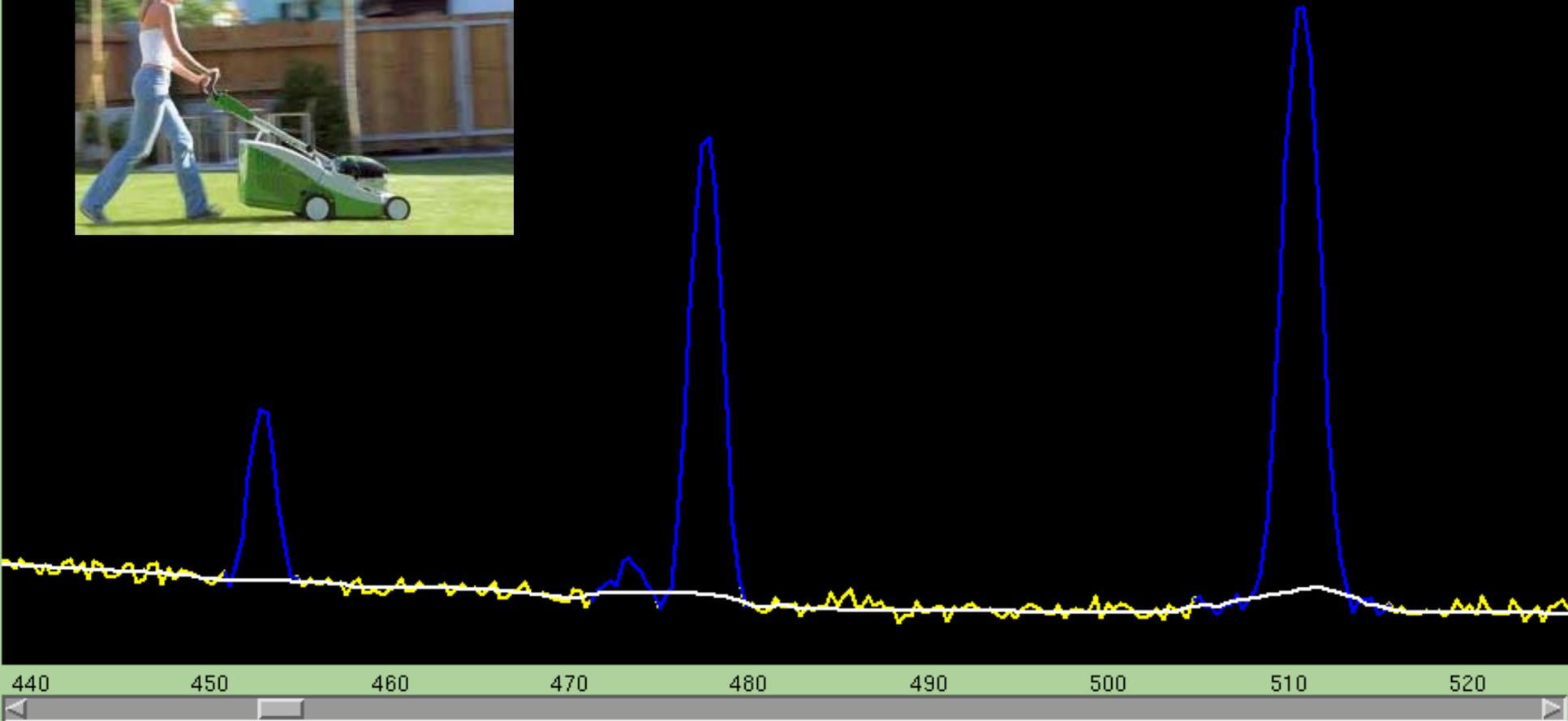
Lawnmower operation



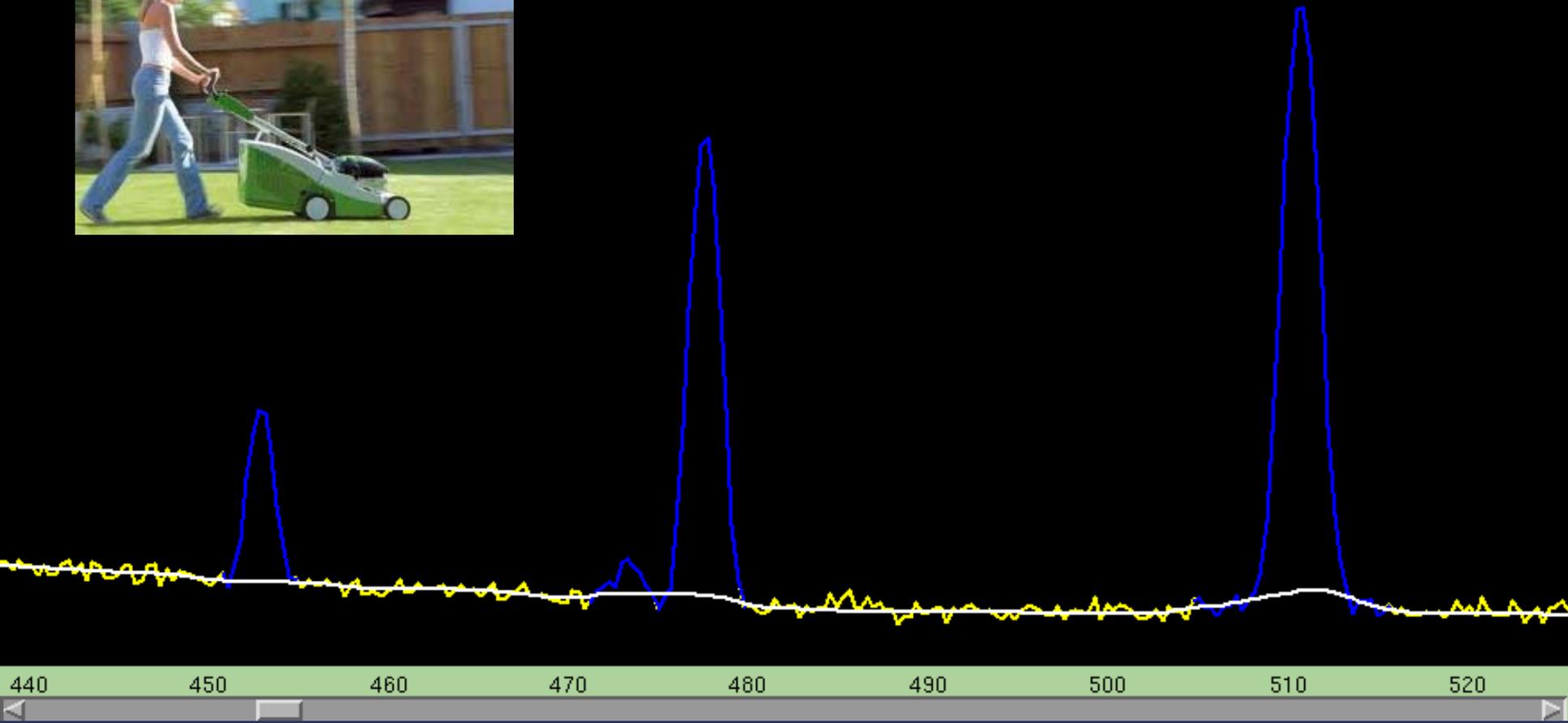
Lawnmower operation



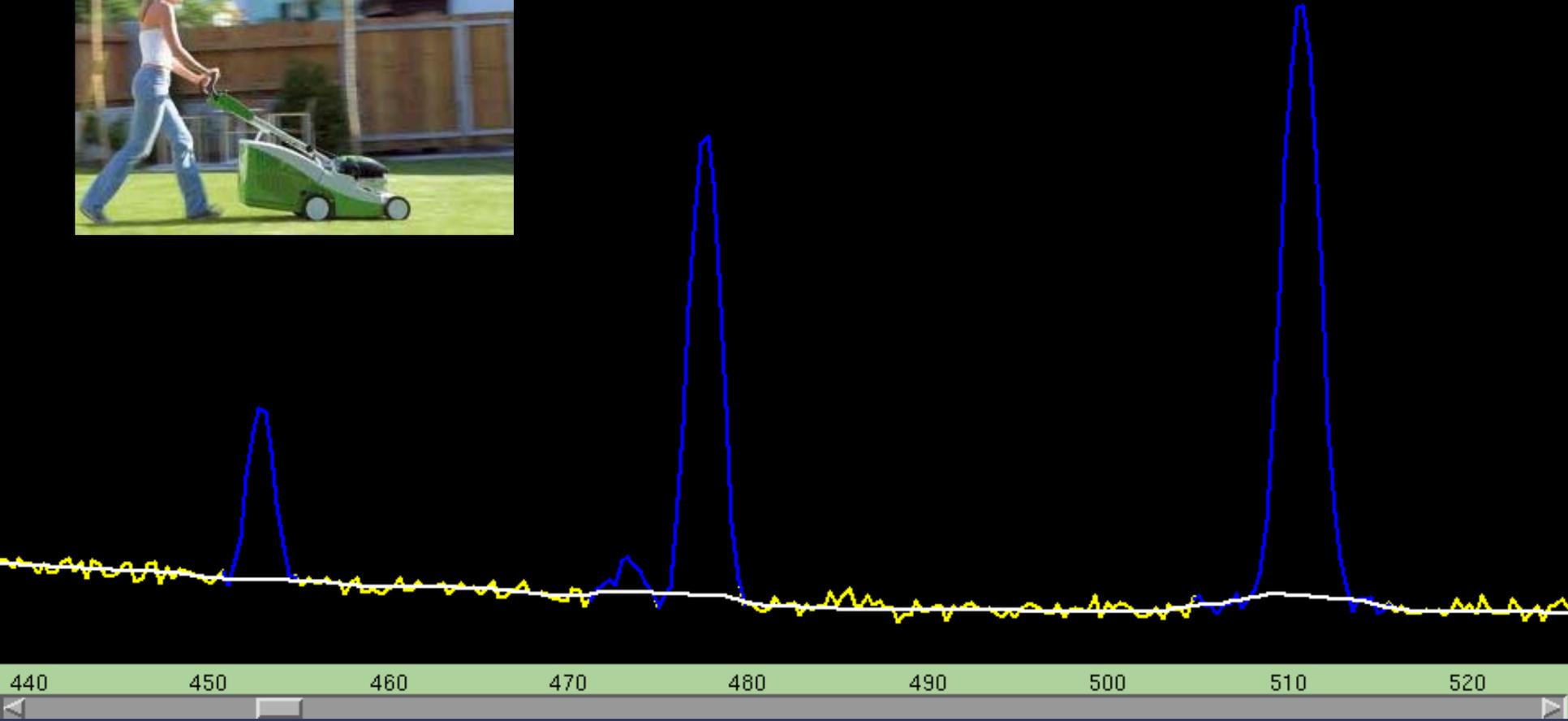
Lawnmower operation



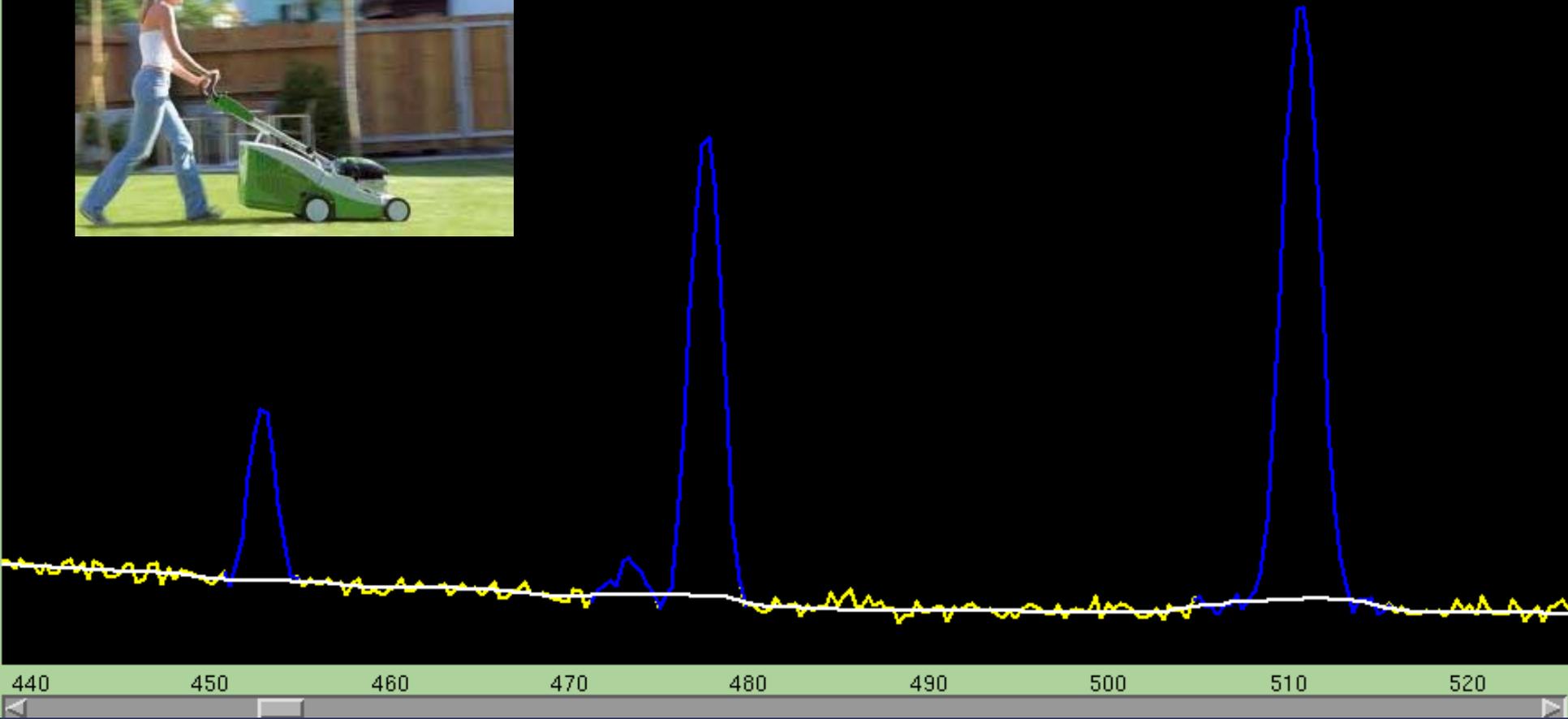
Lawnmower operation



Lawnmower operation

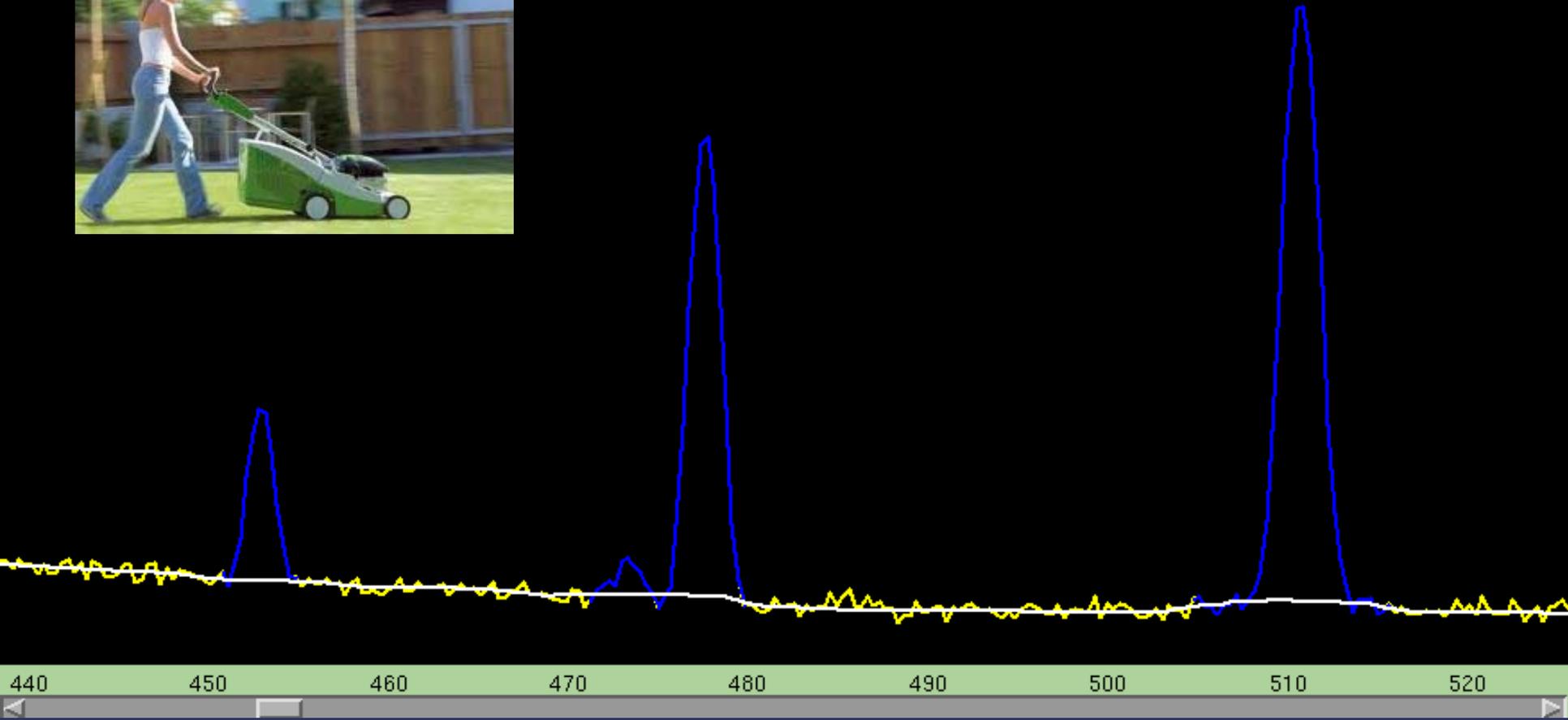


Lawnmower operation

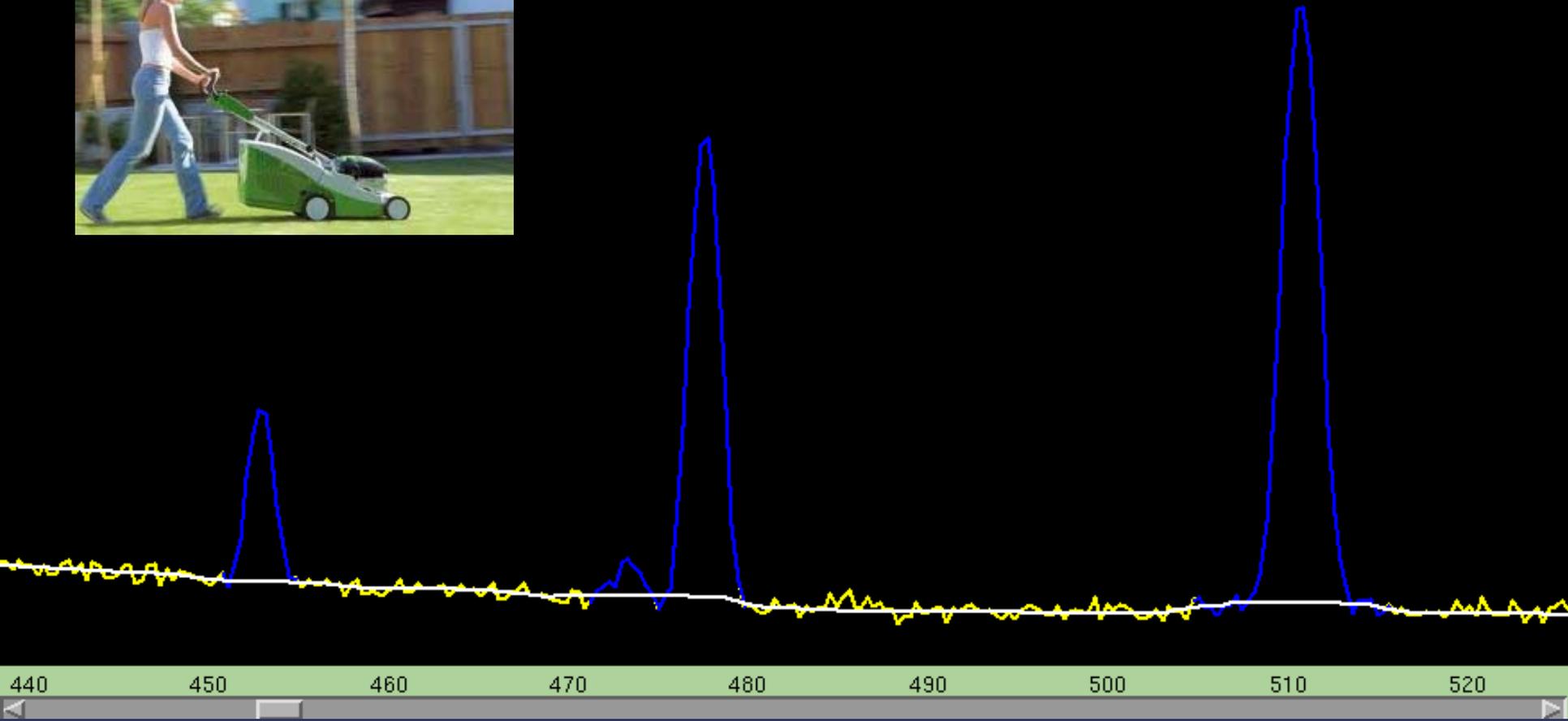




Lawnmower operation

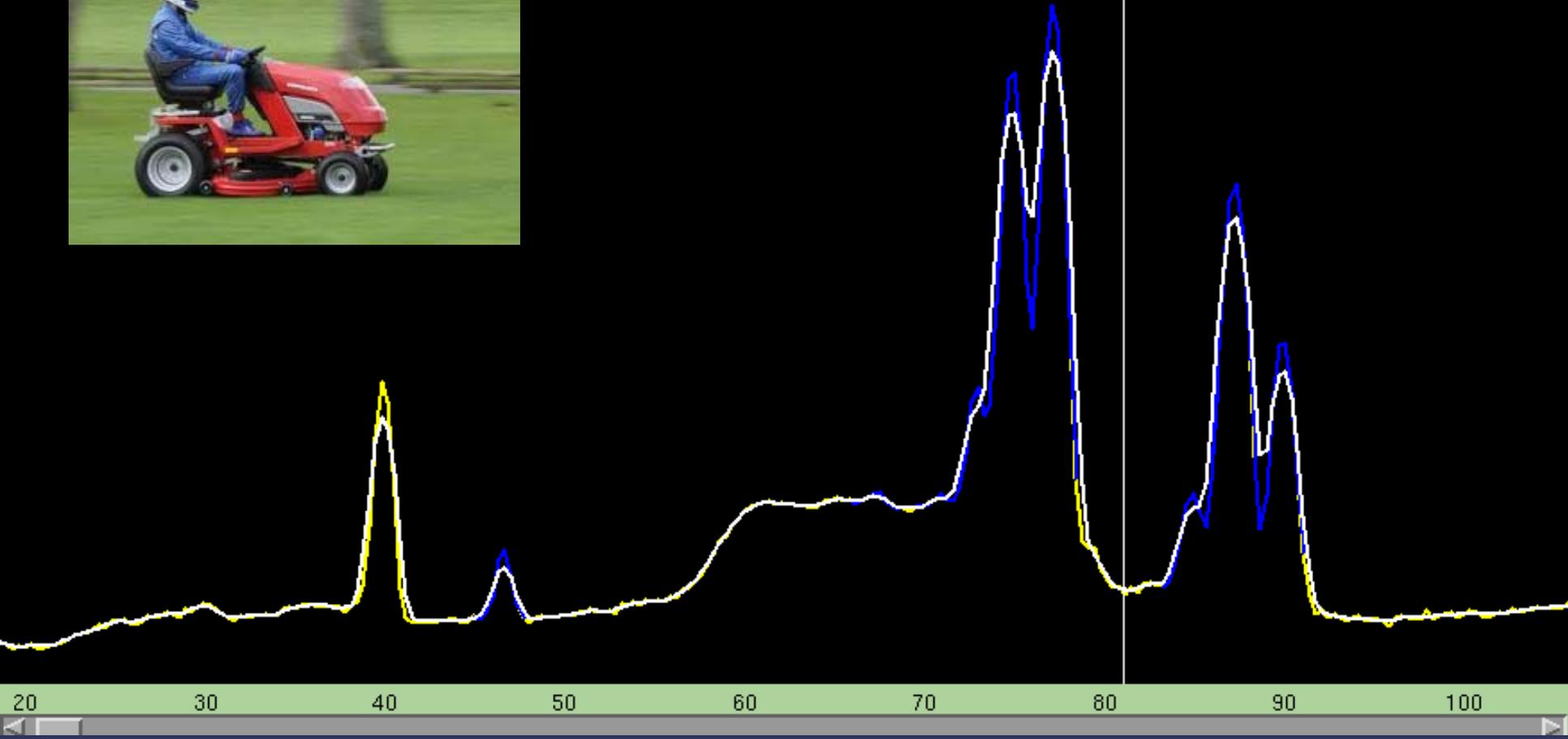


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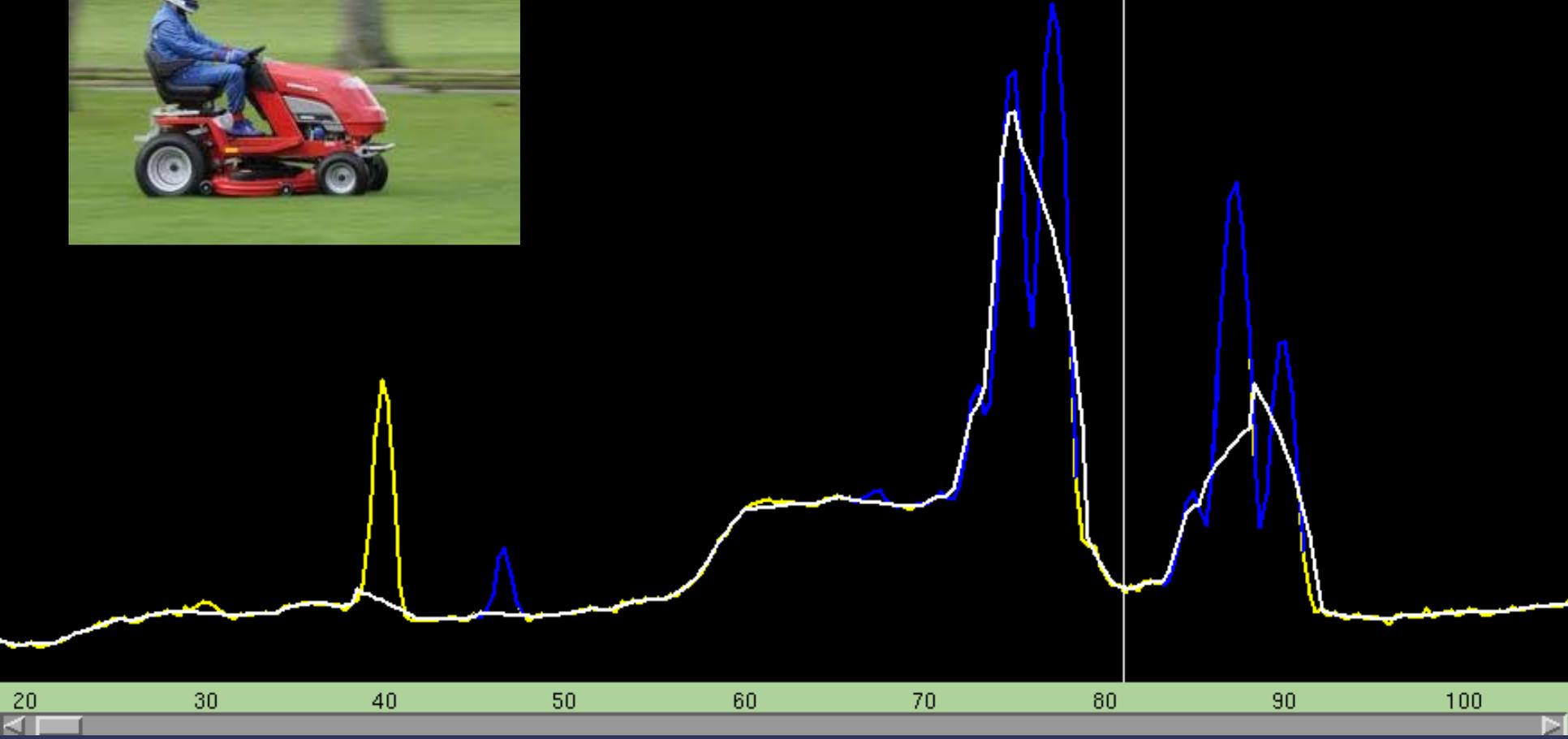


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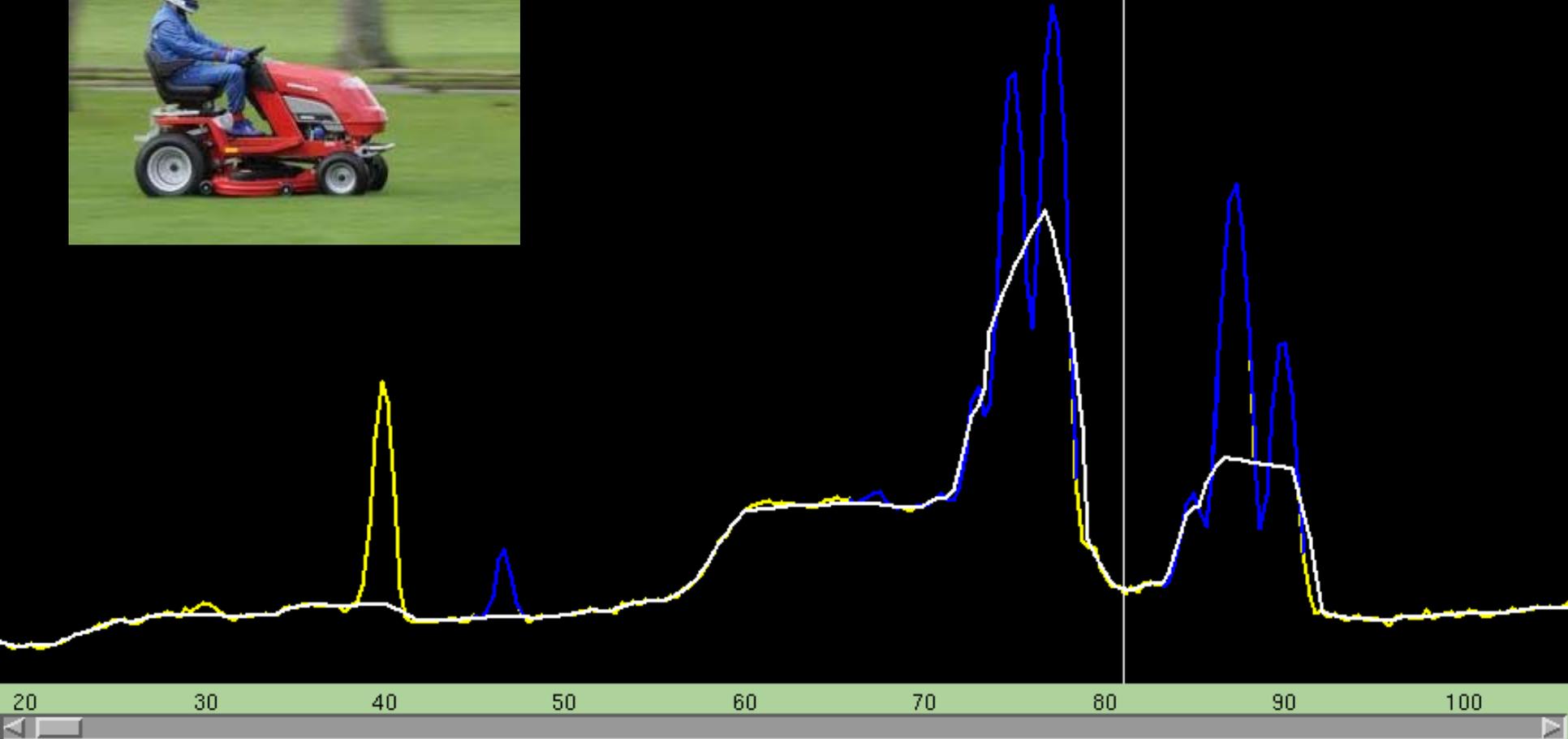


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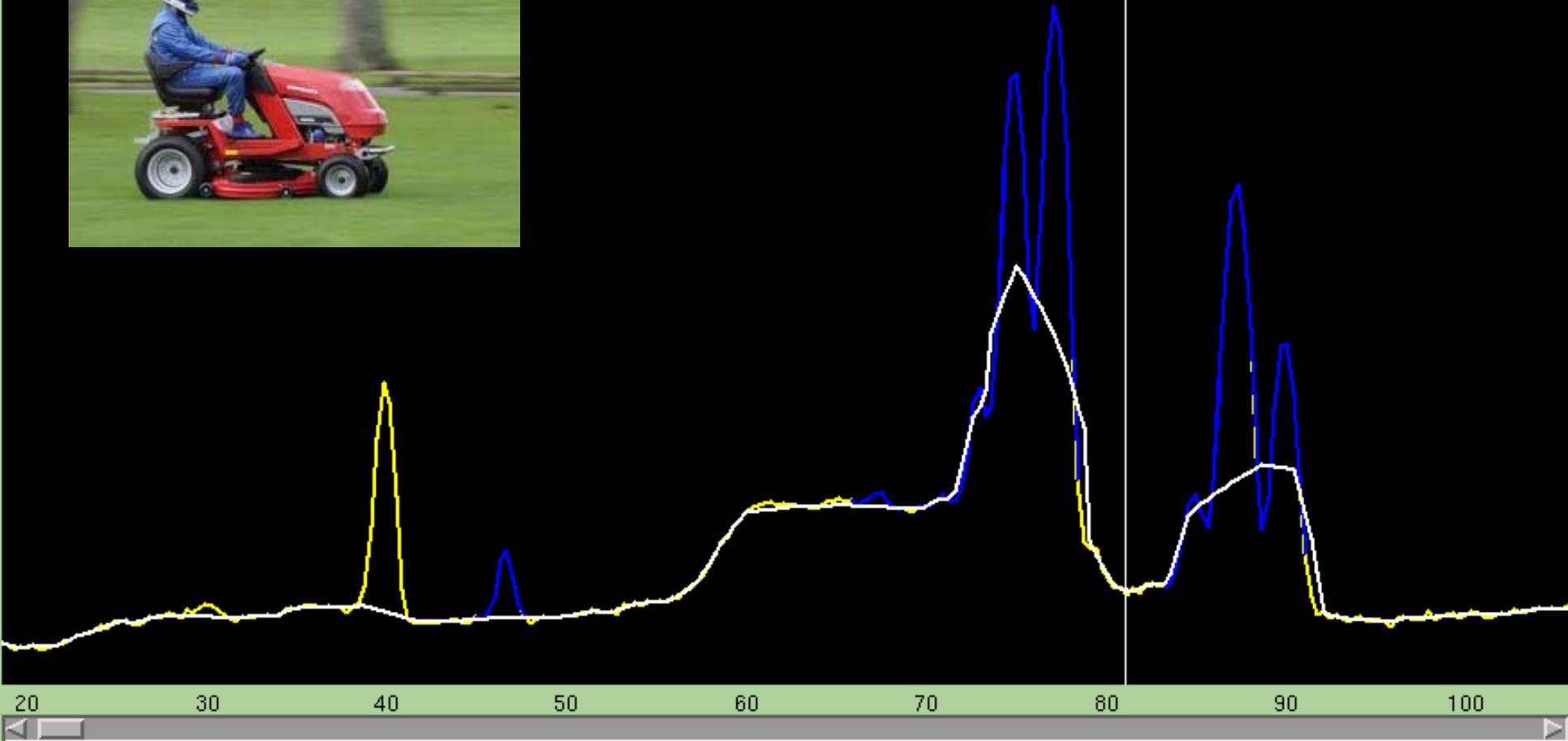




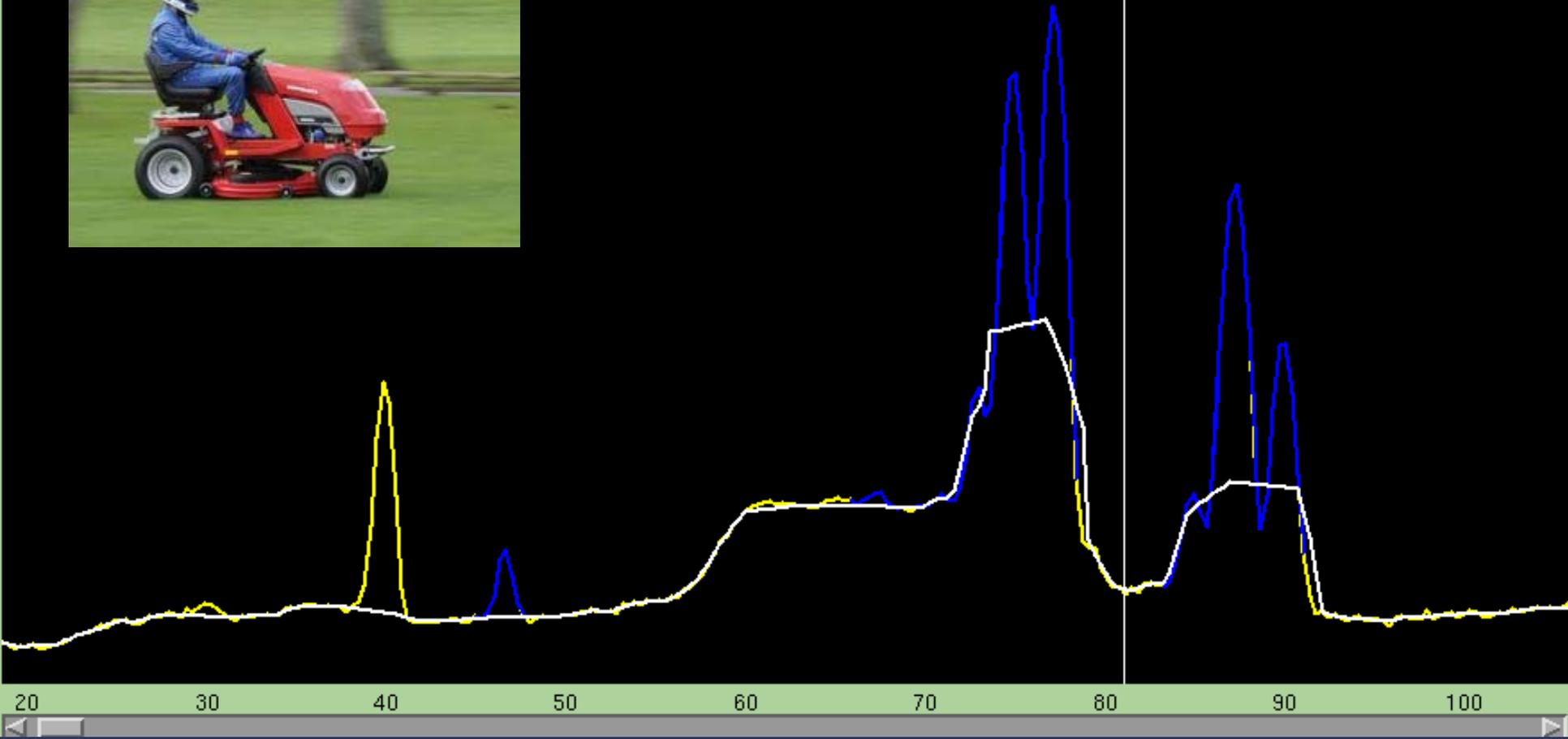
Lawnmower operation



Lawnmower operation

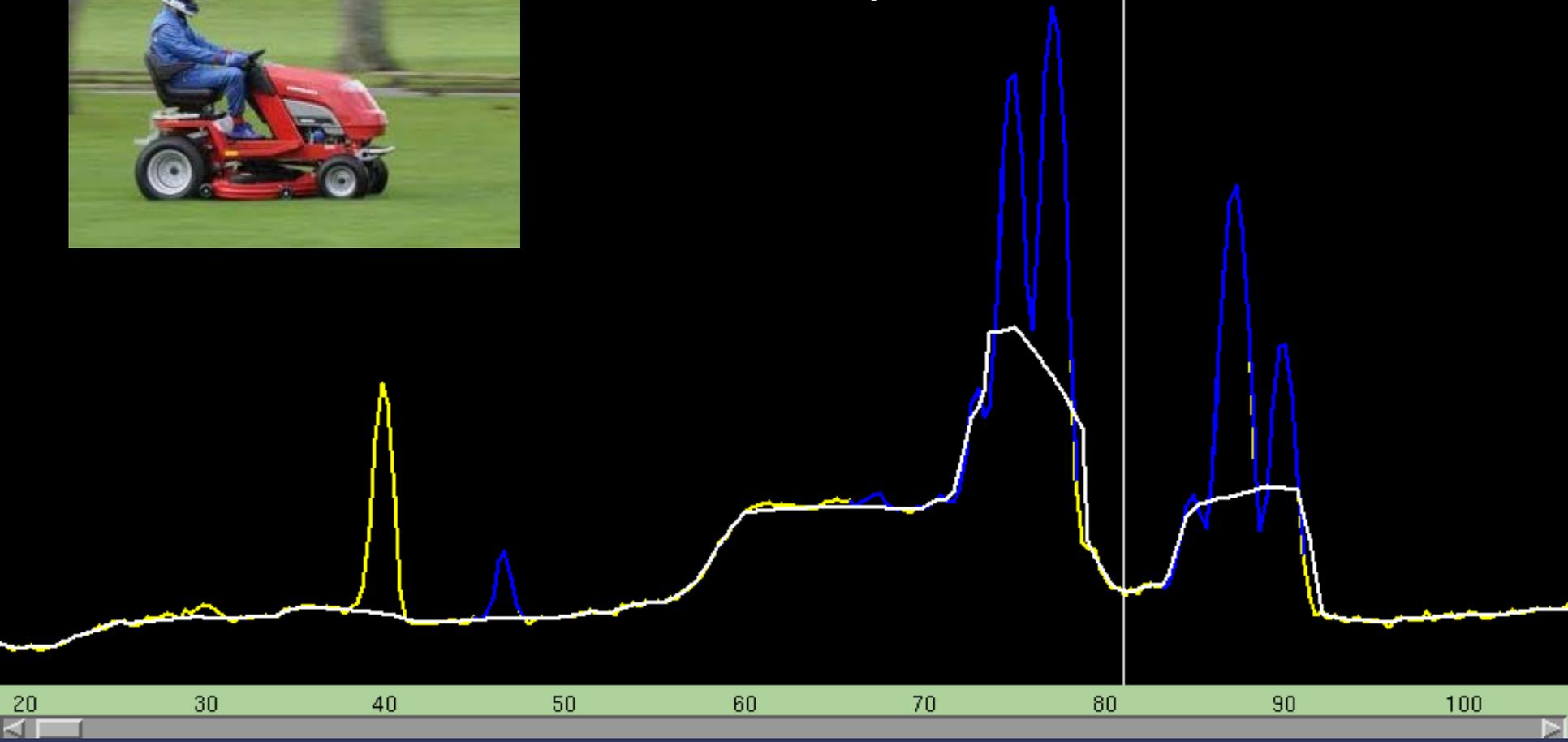


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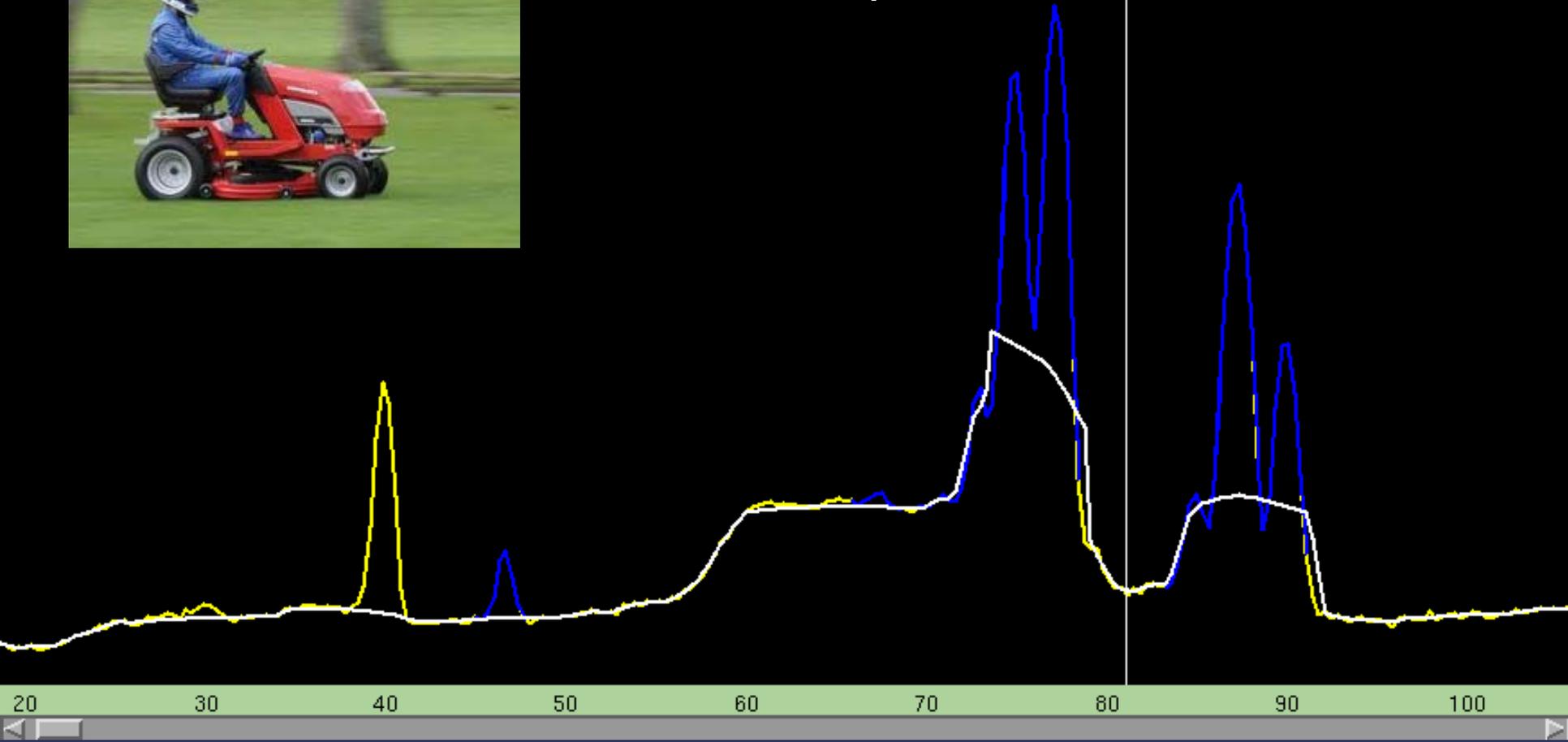


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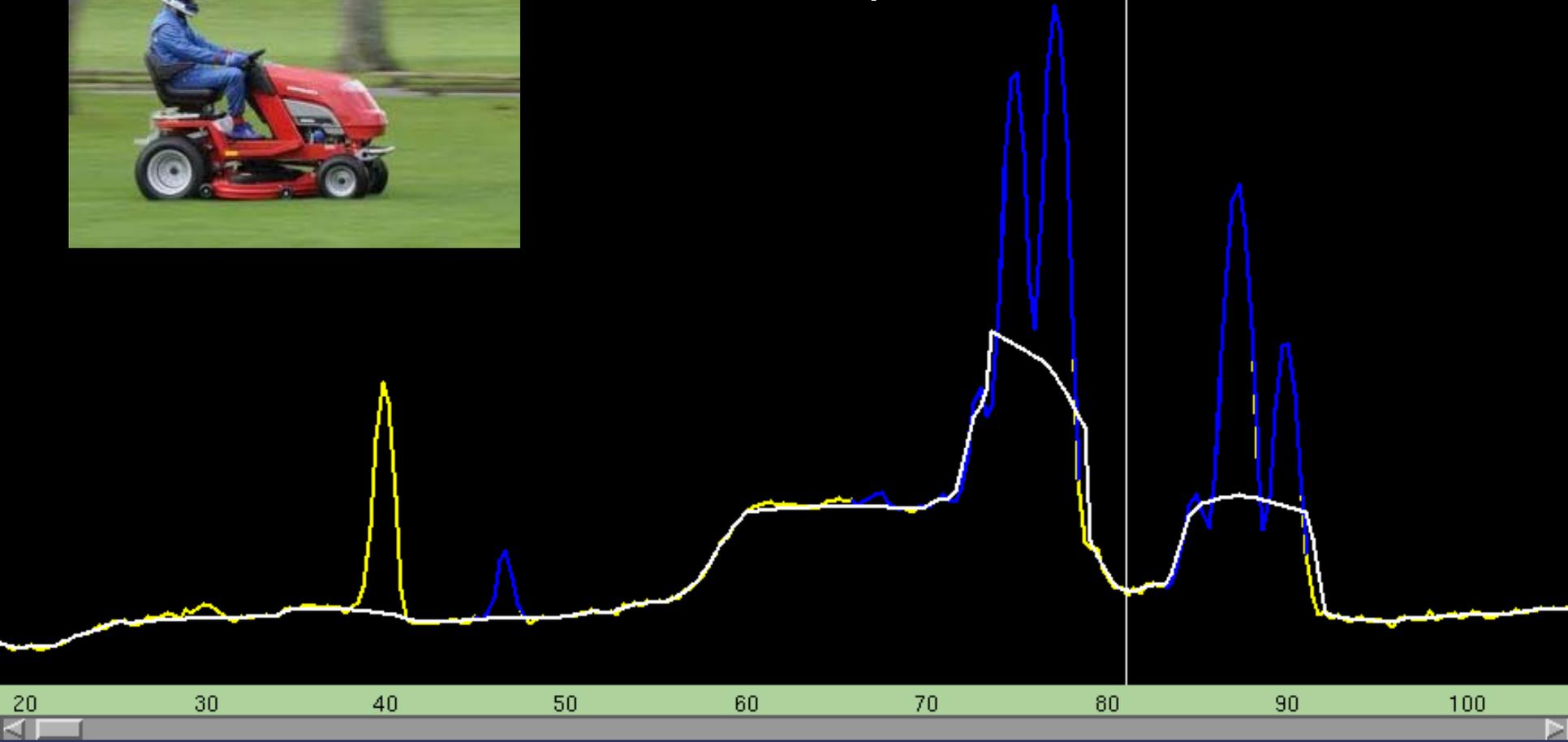


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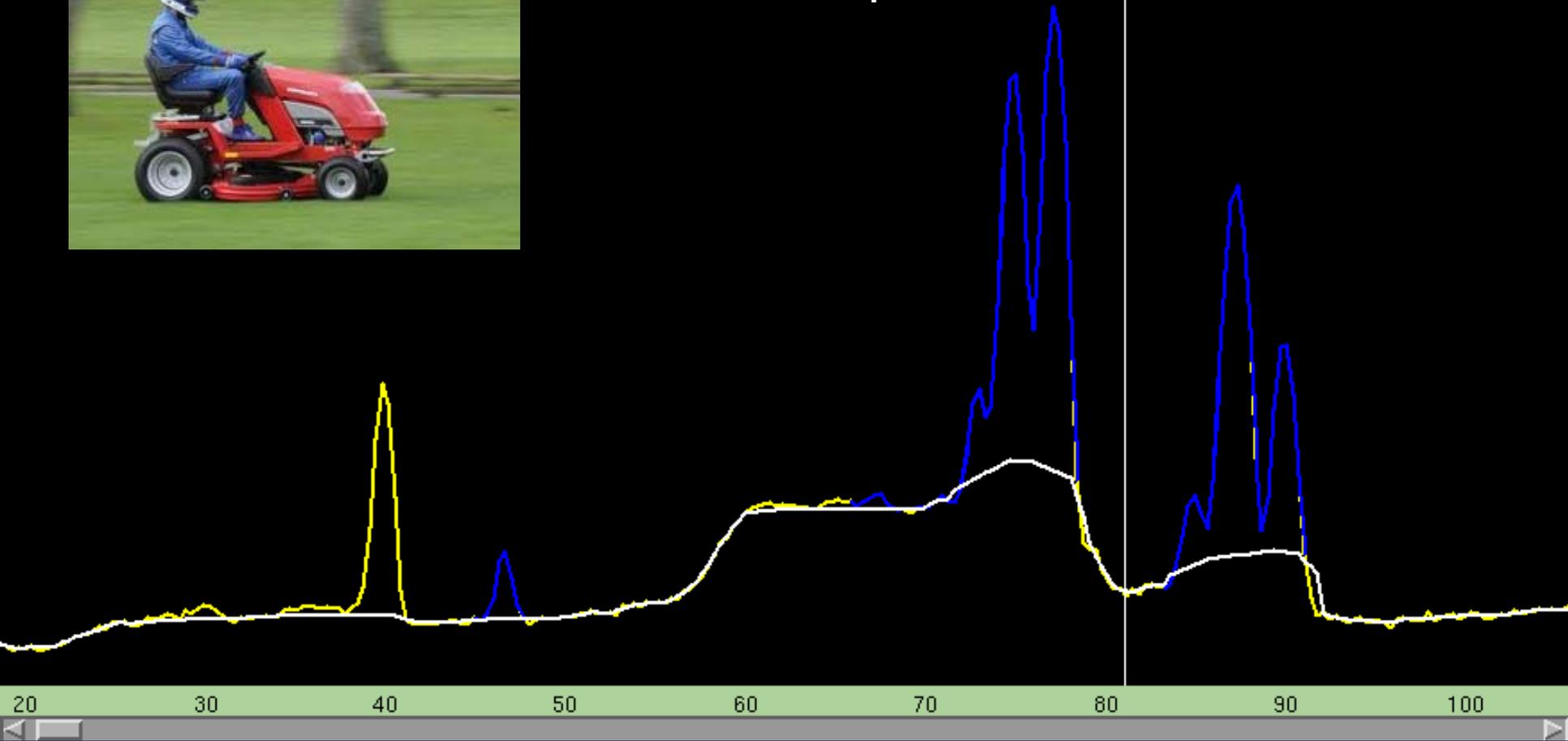


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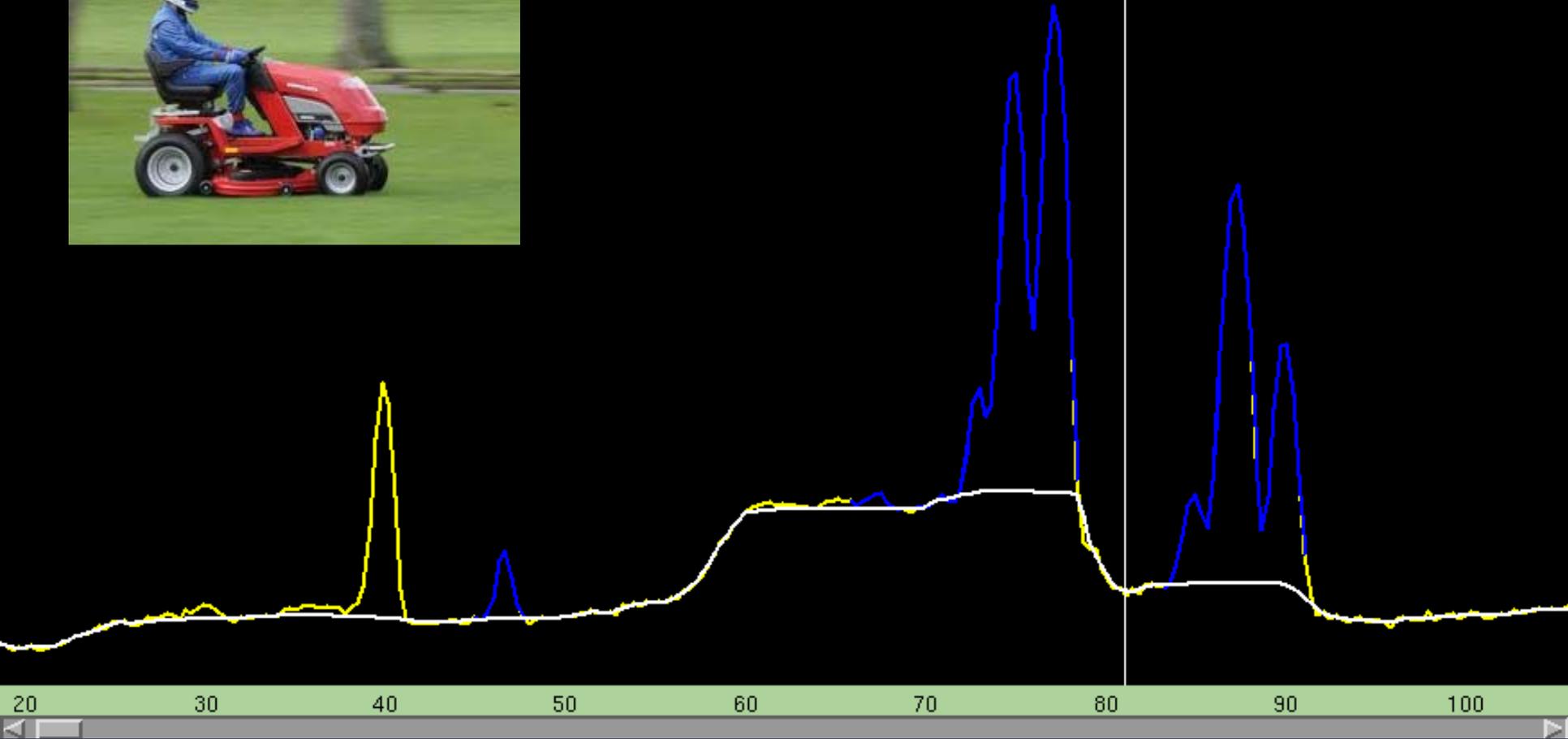


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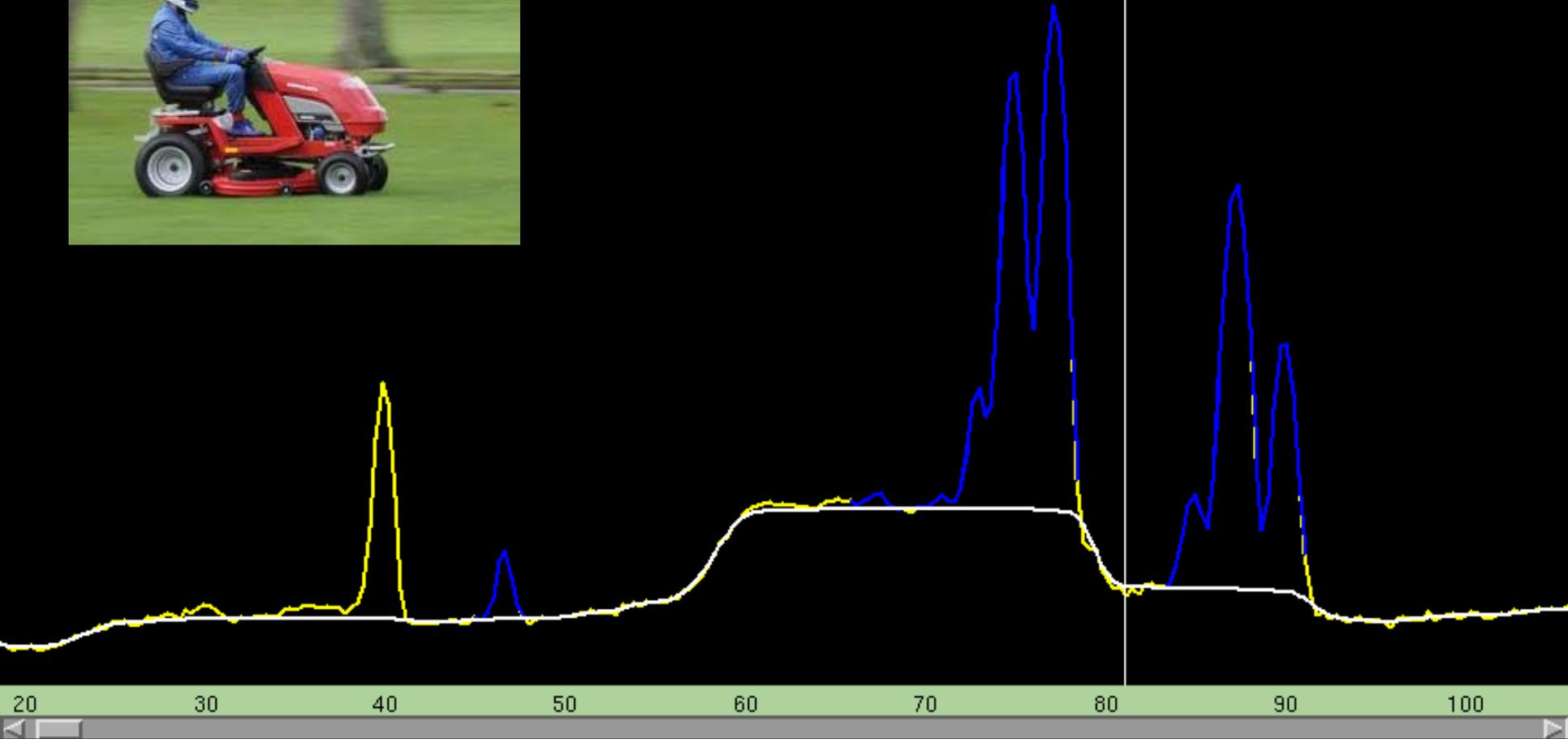


Lawnmower operation

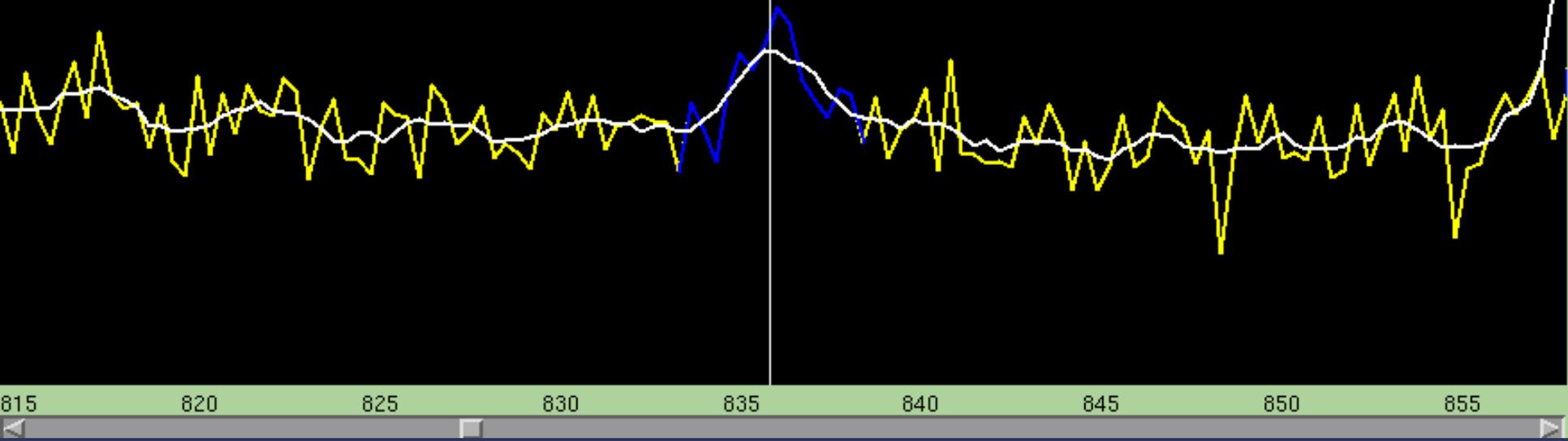
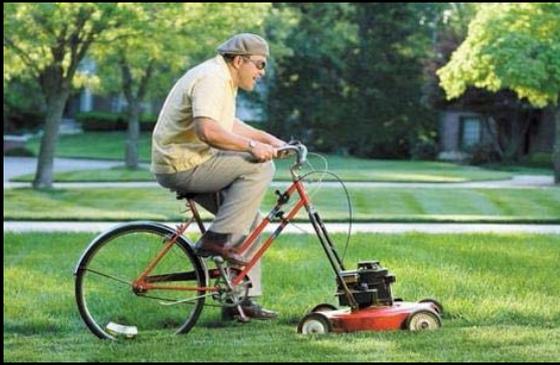




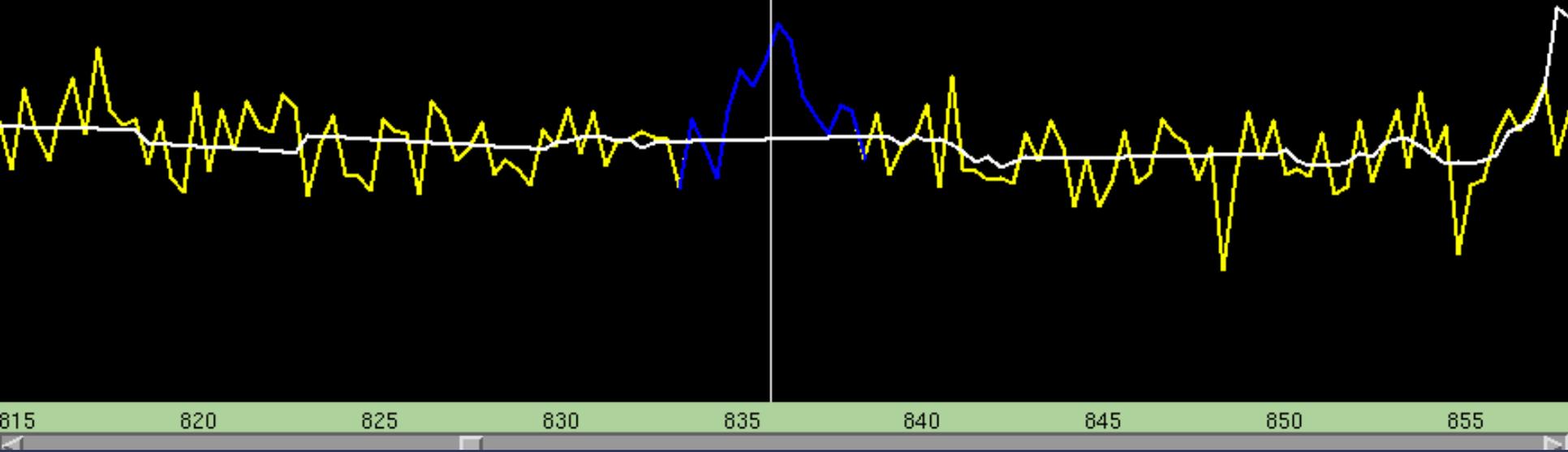
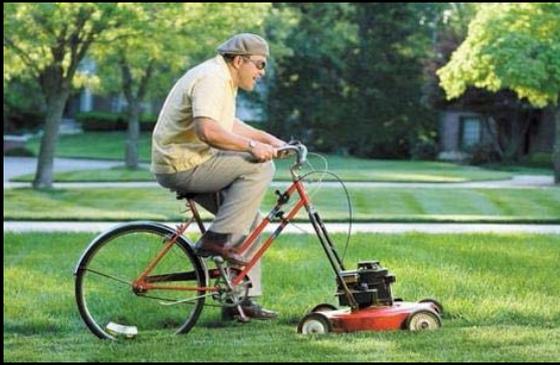
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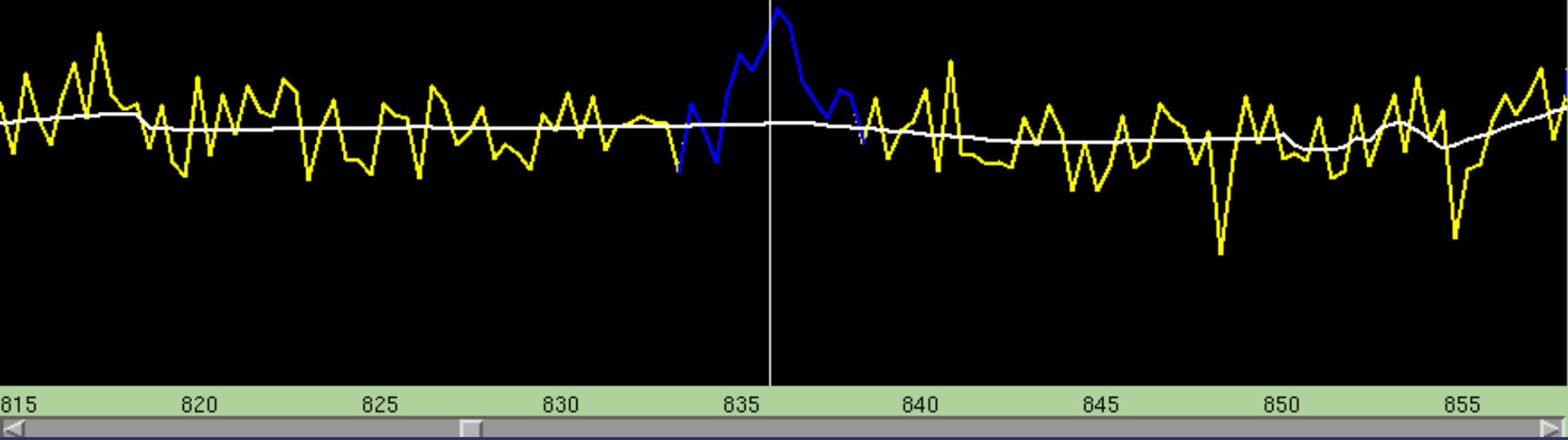
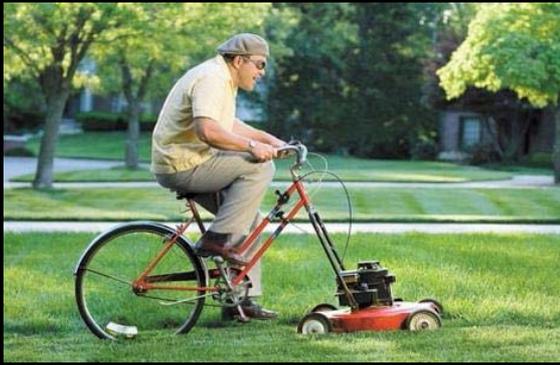
Lawnmower operation



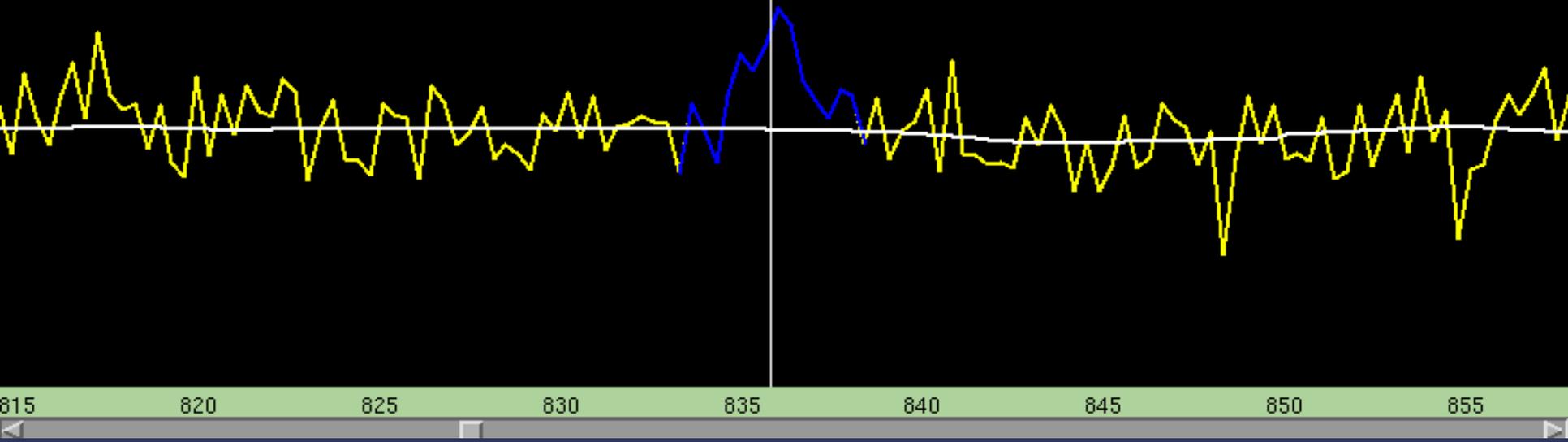
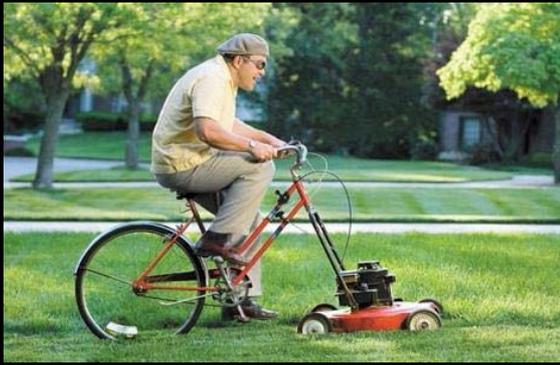
Lawnmower operation

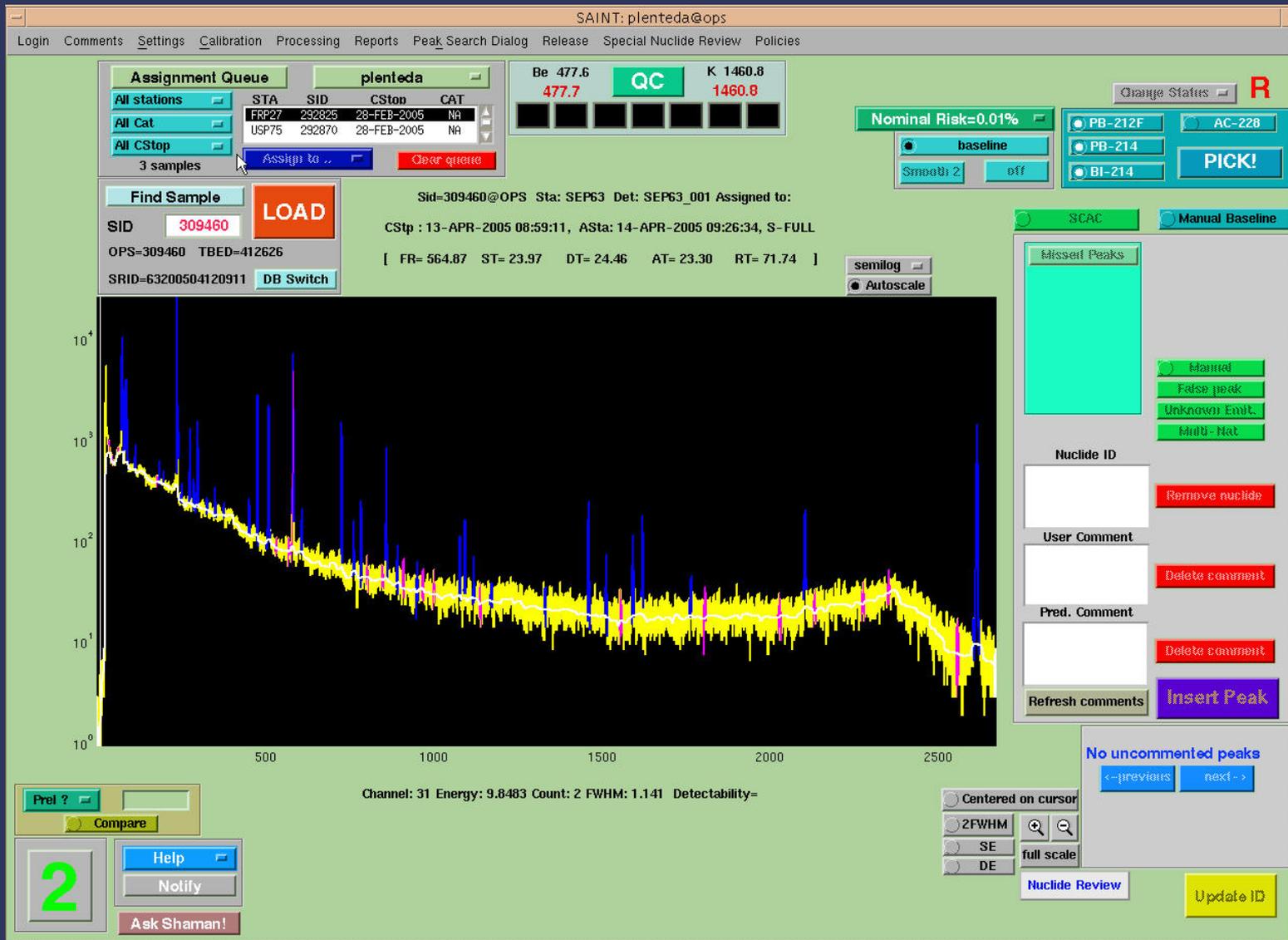


Lawnmower operation

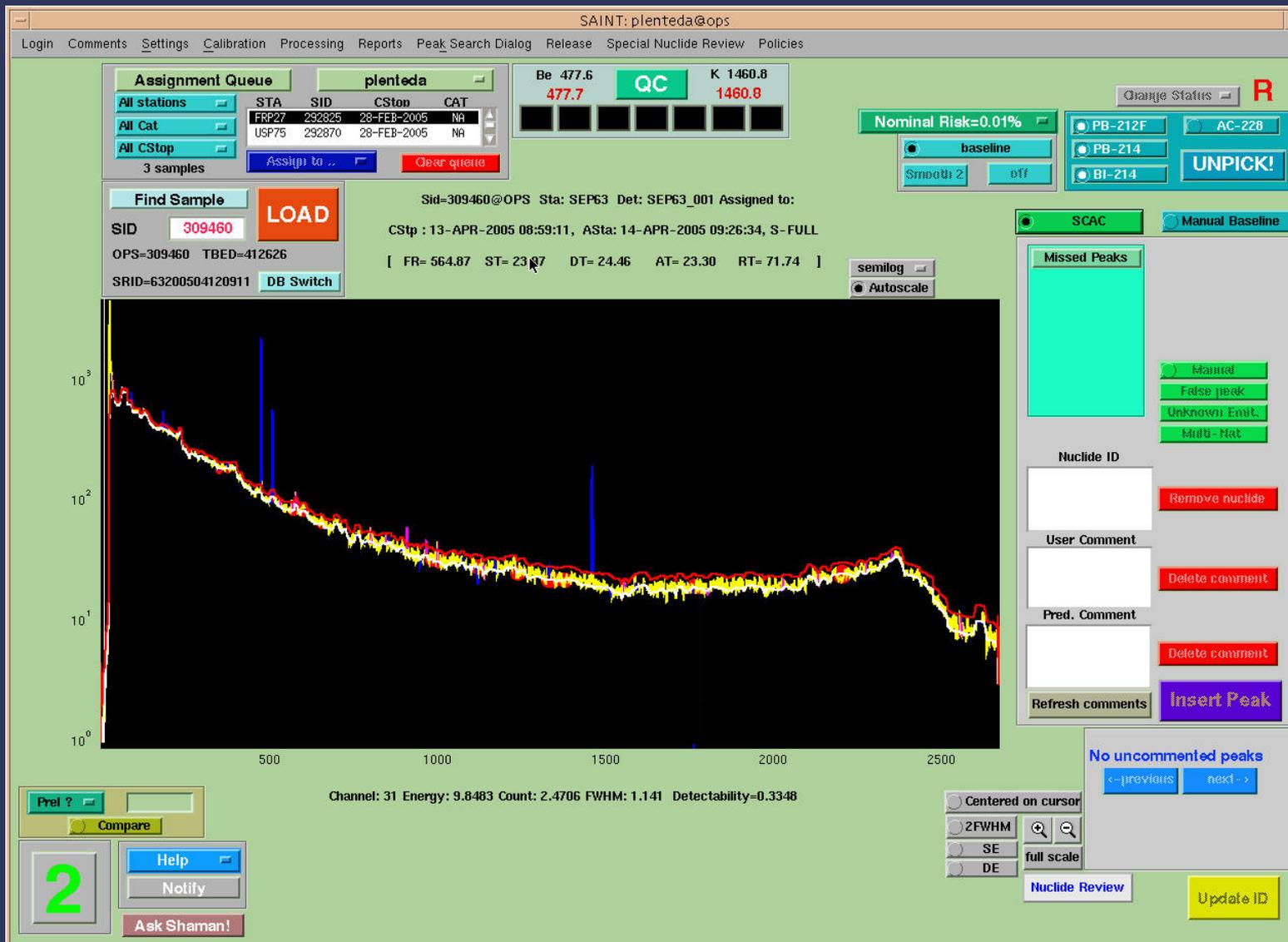


Lawnmower operation





SAINT opening window



SAINT

SCAC-LCC



SAINT

SCAC-LCC

TBED

SID 176123

LOAD

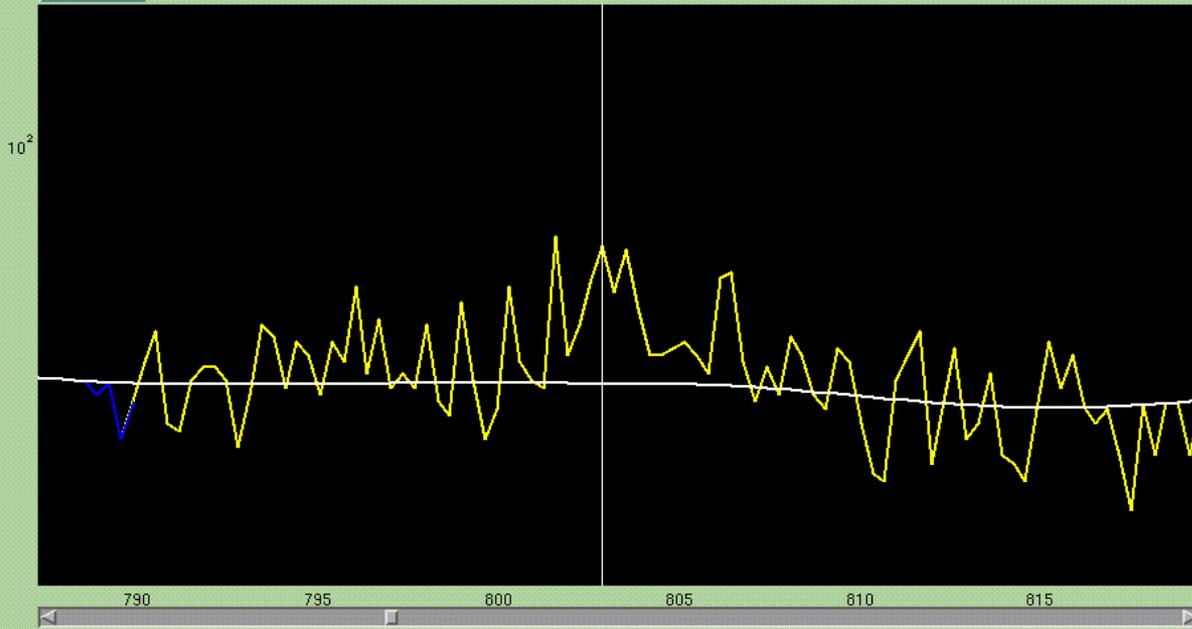
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CStp : 22-AUG-2003 08:55:05, ASta: 23-AUG-2003 09:23:05

[FR 532.4062, ST 22.65, DT 24.4667, AT 23.43]

OPS=124608 TBED=176123

Switch



1

PB-212

UNPICK!

Manual Strip

L Unstrip! R

baseline

Smooth 8 off

LCC SCAC

alfa-beta=0.05%

Type2

196.07

336.197

401.501

9 Anthro-T2

<- ->

Update ID

Autoscale

semilog

Type1 Nat-T2 Manual

Unknown Emit.

Auto Real Peak

Remove nuclei

User

Delete comment

Predefined

Delete comment

Refresh comments

Centered on cursor

2FWHM

full scale



5 uncommented peaks

<-previous next->

TBED

SID 176123

LOAD

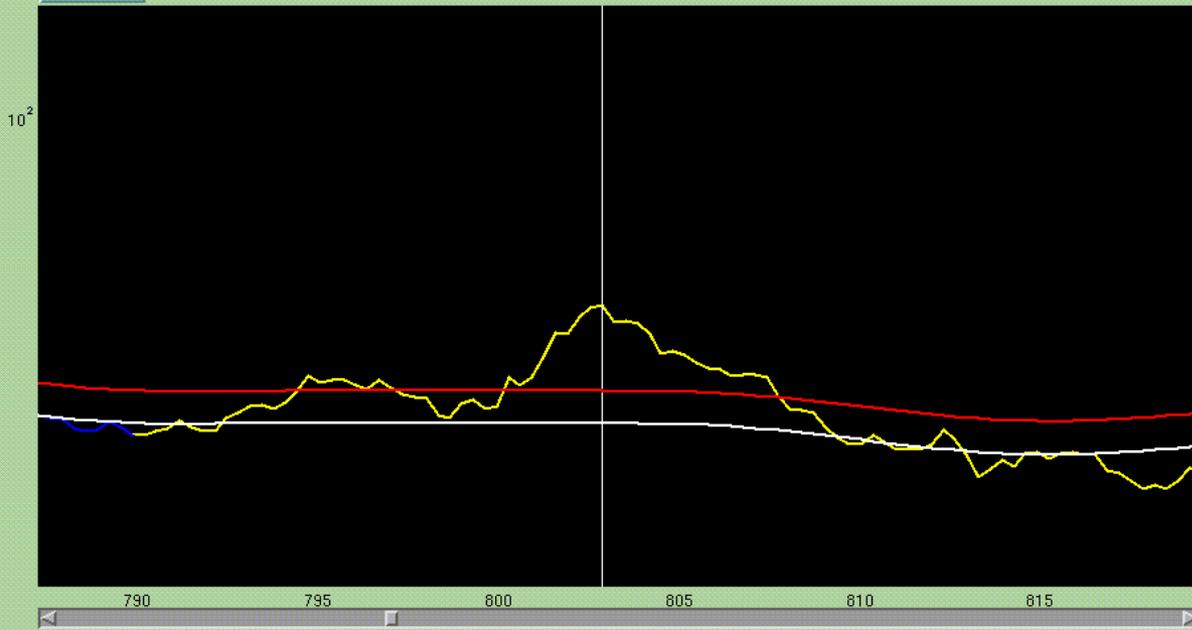
Sid=176123@TBED Sta: SEP63 Det: SEP63_001 Assigned to: unassigned

CStp : 22-AUG-2003 08:55:05, ASta: 23-AUG-2003 09:23:05

[FR 532.4062, ST 22.65, DT 24.4667, AT 23.43]

OPS=124608 TBED=176123

Switch



Possible association: TE-129, SB-128, PB-206, CS-134

1

PB-212

UNPICK!

Manual Strip

L Unstrip! R

baseline

Smooth 8 off

LCC SCAC

Type2

125.802

132.476

143.219

alfa=beta=5.0%

Update ID

Autoscale

42 Anthro-T2

< >

semilog

Type1 Nat-T2 Manual

Unknown Emit.

Auto Real Peak

Remove nuclei

User

Delete comment

Predefined

Delete comment

Refresh comments

Centered on cursor

2FWHM

full scale



5 uncommented peaks

<-previous next->

TBED

SID 176123

LOAD

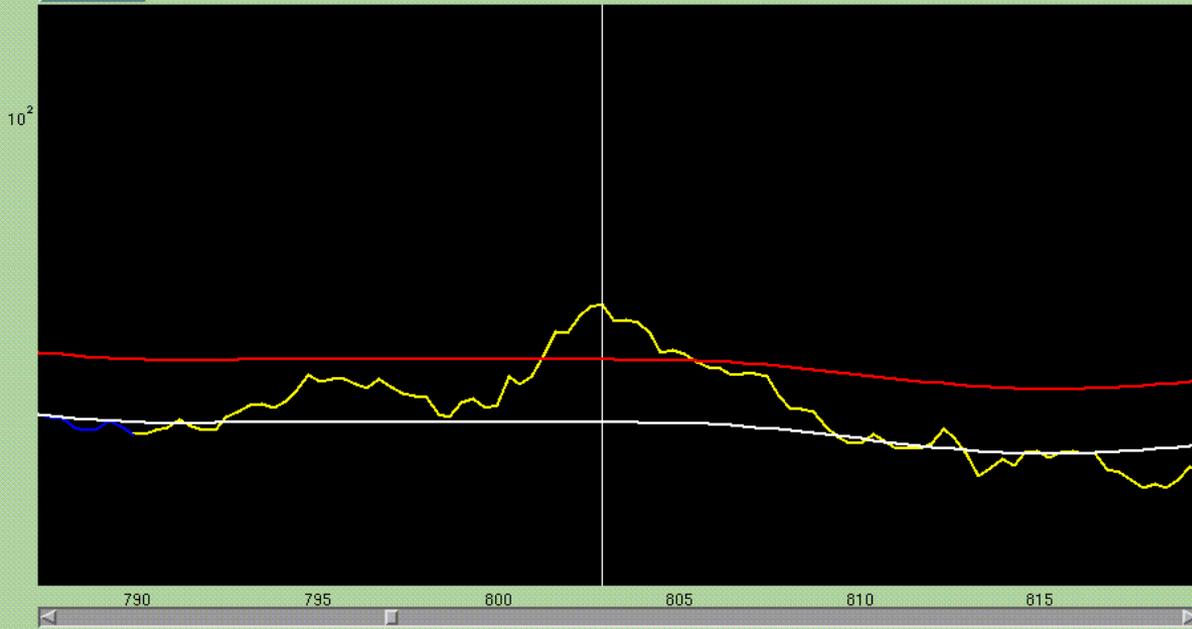
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[FR 532.4062, ST 22.65, DT 24.4667, AT 23.43]

OPS=124608 TBED=176123

Switch



Channel: 2467 Energy: 802.8837 Count: 65.2219 FWHM: 2.3412 Pick_Coeff: -

Possible association: TE-129, SB-128, PB-206, CS-134

1

PB-212

UNPICK!

Manual Strip

L Unstrip! R

baseline

Smooth 8 off

LCC SCAC

alfa=beta=0.05%

Type2

196.07

336.197

401.501

9 Anthro-T2

< >

Update ID

Autoscale

semilog

Type1 Nat-T2 Manual

Unknown Emit. Real Peak

User

Remove nuclei

Predefined

Delete comment

Refresh comments

Delete comment

Centered on cursor

2FWHM

full scale



5 uncommented peaks

<-previous next->

TBED

SID 176123

LOAD

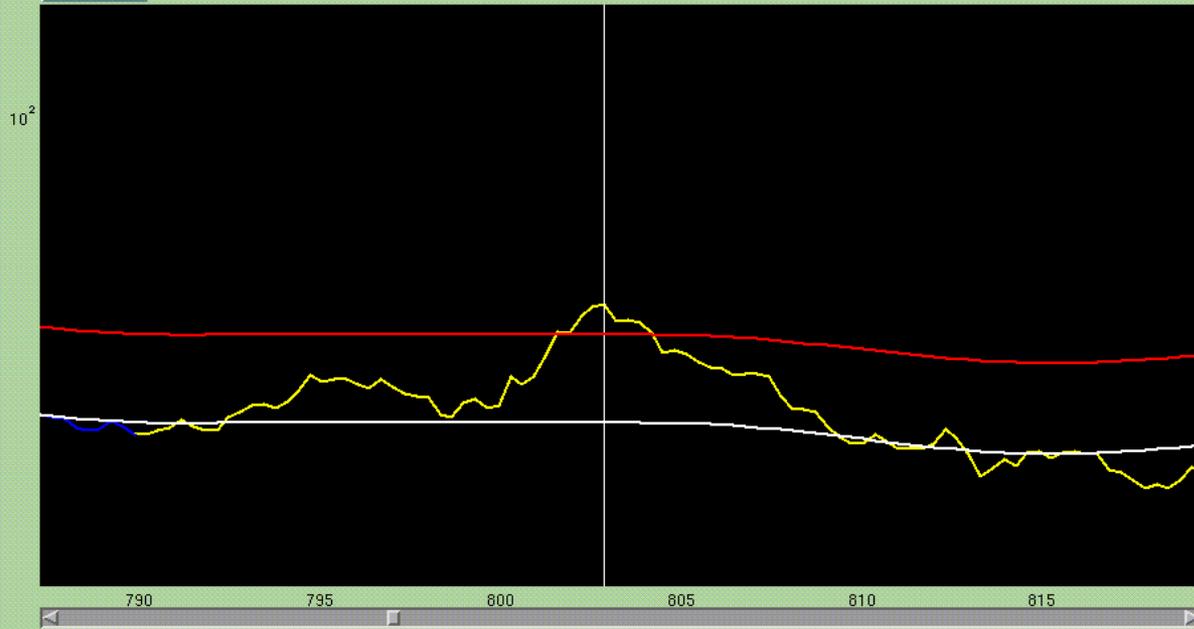
Sid=176123@TBED Sta: SEP63 Det: SEP63_001 Assigned to: unassigned

CStp : 22-AUG-2003 08:55:05, ASta: 23-AUG-2003 09:23:05

[FR 532.4062, ST 22.65, DT 24.4667, AT 23.43]

OPS=124608 TBED=176123

Switch



Channel: 2467 Energy: 802.8837 Count: 65.2219 FWHM: 2.3412 Pick_Coeff: -

Possible association: TE-129, SB-128, PB-206, CS-134

1

PB-212

UNPICK!

Manual Strip

L Unstrip! R

baseline

Smooth 8 off

LCC SCAC

alpha=beta=0.0001%

Type2

802.884

2204.87

2 Anthro-T2

<- ->

Update ID

Autoscale

semilog

Type1 Nat-T2 Manual

Unknown Emit. Real Peak

Remove nuclei

User

Delete comment

Predefined

Delete comment

Refresh comments

Centered on cursor

2FWHM

full scale



5 uncommented peaks

<-previous next->

Thank you

















How do we decide what risks we should take?

1. The PEAKMAKER and PLATEMAKER Mathematica programs provide interfaces between mathematics and what we think we see on the display.
2. Gain experience. Murray has worked on this for a while in Romano's L_c -GUI and I think I hear Murray vote for an α between 0.0001% and 0.001%.
3. To keep the quality pressure on the stations everything stays the same if the SCA width is set to 1.25 FWHM, m to \odot and α to 0.05%, when the barium-140 detection limit is analyzed.

The “Lawnmower” algorithm



To solve the first problem we should “cut” as most we can the spectrum peaks before smoothing. The following algorithm can be used:

For each channel j we define a channel interval $j - \delta^1, j + \delta^2$ where δ^1 and δ^2 are the equivalent in channels of $2 * \text{FWHM}(j)$.

If the channel j happens to be the one with maximum counts in this interval, it is a good candidate to be the centroid of a potential peak.

In this case the original spectrum in the selected interval is replaced with a straight line from $j - \delta^1$ to $j + \delta^2$.

After 2 or 3 times of application of this algorithm all the peaks will be cut out.



Coupling the smoothing and the Lawnmower algorithm we have a very good baseline for all the energies with the exception of some structures like saw teeth where only the smoothing is applied. Compton edges are preserved.

In the case of multiplets the width of the structure is larger than FWHM, therefore the loop smooth-Lawnmower will give a final structure of a plateau far from the right baseline. To solve this problem, on the X-rays region, the Lawnmower algorithm is applied many times before smoothing.

Application: the recipe



Due to the large amount of multiplets in the X-rays region I distinguish two regions:

- 1) 0 - 110 keV**
- 2) 110 - E_{\max}**

1) 2 loops with 1 smoothing and 80 Lawnmowers each.

2) 4 loops with 1 smoothing and 4 Lawnmowers each.

The smoothing algorithm is the most straightforward way to calculate the baseline, it consists in setting the counts to a given channel as the average of the neighbouring channels for a given interval width.

This algorithm, applied in the baseline calculation for gamma spectra, works fine in the absence of peaks structure, since the baseline should not consider the peak counts but only the ones coming from the background.

The baseline itself should follow the shape of certain spectrum structures not recognizable as peaks like Compton edges or saw teeth from (n,n') reactions.

The smoothing

The smoothing algorithm is the most straightforward way to calculate the baseline, it consists in setting the counts to a given channel as the average of the neighbouring channels for a given interval width.

The baseline itself should follow the shape of certain spectrum structures not recognizable as peaks like Compton edges or saw teeth from (n,n') reactions.